

# Advanced Quantum Mechanics

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Winter term 2019/20

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## Chapter 1: Recapitulation: structure of the quantum theory

### 1.1 State space

\* classical mechanics: maximal information = values  $\{q_i, p_i\} \in$  phase space (coordinates  $q_i$ , momenta  $p_i$ )

\* quantum mechanics: state vector  $|\psi\rangle$  in a typically  $\infty$ -dimensional state space  $\mathcal{H}$  known as Hilbert space (axioms: see books, Nolting / Weinberg)

Dirac's "bra-ket" notation:  $|\psi\rangle \in \mathcal{H}$  ("ket")

$$\text{Norm } \|\psi\rangle\|^2 = (\psi, \psi)$$

$$\text{Scalar product: } (\psi, \psi) \in \mathbb{C}$$

$$\text{Properties: linearity } (\chi, [\alpha\psi + \beta\varphi]) = \alpha(\chi, \psi) + \beta(\chi, \varphi)$$

$$\text{symmetry } (\psi, \varphi)^* = (\varphi, \psi)$$

$$\text{positivity } (\psi, \psi) > 0 \text{ for } \psi \neq \underbrace{0}_{\text{zero vector}}$$

Special case:

$$\mathcal{H} = \mathcal{L}^2(\mathbb{R}^3) \text{ square-integrable functions on } \mathbb{R}^3$$

$$(\psi, \varphi) = \int d^3x \varphi^*(\vec{x}) \psi(\vec{x})$$

Describe one non-relativistic particle.

Remark:  $|\psi\rangle$  and  $\lambda|\psi\rangle$  ( $\lambda \neq 0, \lambda \in \mathbb{C}$ ) describe the same state. ②

Physical state is a one-dimensional complex subspace of  $\mathcal{H}$  (equivalence class, projective geometry)

Convention: take representative

$\| |\psi\rangle \| = 1 \rightarrow$  still  $\{ e^{i\alpha} |\psi\rangle \mid \alpha \in \mathbb{R} \}$  one state

\* Dual space

$\mathcal{H}^*$  is the space of linear functionals  $\mathcal{H} \rightarrow \mathbb{C}$

For each "ket"  $|\varphi\rangle$  there is a "bra"  $\langle\varphi|$  such that

$$\langle\varphi|\psi\rangle = (\langle\varphi|, |\psi\rangle)$$

This defines a mapping  $\mathcal{H} \rightarrow \mathcal{H}^*$

$$|\varphi\rangle \mapsto \langle\varphi|$$

It is anti-linear:  $(\lambda|\varphi\rangle) \mapsto \lambda^* \langle\varphi|$

Proof:  $(\lambda|\varphi\rangle, |\psi\rangle) = (|\psi\rangle, \lambda|\varphi\rangle)^* = \lambda^* (|\psi\rangle, |\varphi\rangle)$  □

\* Continuum states

states of the type  $\langle p | p' \rangle = \delta(p - p')$ ,  $p, p' \in \mathbb{R}$

$$\Rightarrow \langle p | p \rangle = \delta(0) = \infty$$

not a physical state (distributions)

Superposition of continuum states can form a physical state

Example: plane waves  $e^{i\vec{p} \cdot \vec{x}}$  enlarge the space  $L^2(\mathbb{R}^3)$

Continuum states also have "bra-ket" mapping

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\* In the following we write the scalar product in the "bra-ket" notation:

$$\langle x | \psi \rangle \equiv (|x\rangle, |\psi\rangle).$$

## 1.2 Observable

Consider an operator  $A: \mathcal{H} \rightarrow \mathcal{H}$

$$|\psi\rangle \mapsto A|\psi\rangle$$

$A$  is linear if  $A(|\psi\rangle + |\varphi\rangle) = A|\psi\rangle + A|\varphi\rangle$ ,  $A(\lambda|\psi\rangle) = \lambda A|\psi\rangle$

The adjoint  $A^\dagger$  of any operator  $A$  (linear or not) is defined through:

$$\begin{array}{ccc} \mathcal{H} & : & |\psi\rangle \mapsto A^\dagger|\psi\rangle =: |\chi\rangle \\ \downarrow & & \downarrow \\ \mathcal{H}^* & : & \langle\psi| \mapsto \langle\psi|A = \langle\chi| \end{array}$$

$$\Leftrightarrow \langle\varphi|A^\dagger|\psi\rangle = \langle\psi|A|\varphi\rangle^*$$

Proof: define  $|\chi\rangle = A^\dagger|\psi\rangle$ , so  $\langle\chi| = \langle\psi|A$ :

$$\langle\varphi|A^\dagger|\psi\rangle = \langle\varphi|\chi\rangle \stackrel{\substack{\uparrow \\ \text{property of scalar product}}}{=} \langle\chi|\varphi\rangle^* = \langle\psi|A|\varphi\rangle^* \quad \square$$

It follows:  $(\lambda A)^\dagger = \lambda^* A^\dagger$ .

Basis  $\{|i\rangle\}$  of  $\mathcal{H}$ : any  $|\psi\rangle = \sum_i c_i |i\rangle$ .  $A$  is linear

$$(A^\dagger)_{ij} = \langle i|A^\dagger|j\rangle = \langle j|A|i\rangle^* = (A_{ji})^*$$

$\Leftrightarrow$  matrices, adjoint = transposed complex conjugate

$A$  is Hermitian or self-adjoint if  $A$  is linear and  $A=A^\dagger$ .

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Physical measurable quantities are represented by Hermitian operators.

Expectation value:

$$\langle A \rangle_{|\psi\rangle} = \frac{\langle \psi | A | \psi \rangle}{\langle \psi | \psi \rangle} = \langle \psi | A | \psi \rangle \text{ if } \langle \psi | \psi \rangle = 1$$

$$A = A^\dagger \Rightarrow \langle A \rangle_{|\psi\rangle} \text{ is real.}$$

### 1.2.1 Discrete spectrum

A Hermitian with discrete eigenvectors

$$A|i\rangle = a_i|i\rangle, \quad i \in \mathbb{N}, \quad \langle i|i\rangle = 1.$$

\*  $a_i \in \mathbb{R}$

Proof:  $a_i = \langle i|a_i|i\rangle = \langle i|A|i\rangle \stackrel{A=A^\dagger}{=} \langle i|A|i\rangle^* = a_i^*$  □

\* Eigenvectors are orthogonal.

Proof:  $\langle i|A|j\rangle = a_j \langle i|j\rangle = \langle j|A^\dagger|i\rangle^* = a_i \langle i|j\rangle$

If  $a_i \neq a_j \Rightarrow \langle i|j\rangle = 0$ .

If eigenvalue  $a_i = a_j$  is degenerate one can choose multiple eigenvectors to the same  $a_i$  such that  $\langle i|j\rangle = \delta_{ij}$  (Gram-Schmidt). □

\* Completeness (Proof is non-trivial):

$$\sum_i \underbrace{|i\rangle \langle i|}_{\text{projector onto } |i\rangle} = \underbrace{\mathbb{1}}_{\text{identity operator}}$$

$\{|i\rangle\}$  form a complete orthonormal set (CON)

$$\Rightarrow |\psi\rangle = \underbrace{\sum_i |i\rangle \langle i|\psi\rangle}_{\mathbb{1}} = \sum_i \underbrace{\langle i|\psi\rangle}_{\text{component w.r.t. CON}} |i\rangle$$

$$A|\psi\rangle = \sum_i A|i\rangle \langle i|\psi\rangle = \sum_i a_i |i\rangle \langle i|\psi\rangle$$

$$\Rightarrow A = \sum_i a_i |i\rangle \langle i| \quad \text{spectral decomposition}$$

$$A^2 = \sum_{i,j} a_i a_j |i\rangle \langle i|j\rangle \langle j| = \sum_i a_i^2 |i\rangle \langle i|$$

Definition:  $f(A) := \sum_i f(a_i) |i\rangle \langle i|$  function of operator  $A$

\* Physical interpretation:

$\{a_i\}$  are the possible quantized measurement values of  $A$

$$\begin{aligned} \langle A \rangle_{|\psi\rangle} &= \sum_i a_i \langle \psi|i\rangle \langle i|\psi\rangle = \sum_i a_i |\langle i|\psi\rangle|^2 \quad (\langle \psi|\psi\rangle = 1) \\ &= \sum_i a_i p_i \end{aligned}$$

$p_i = |\langle i|\psi\rangle|^2 = \text{probability to measure value } a_i \text{ for state } |\psi\rangle$

$$1 = \langle \psi|\psi\rangle = \sum_i p_i, \quad p_i \geq 0$$

\* Complete set of observables:

$A$  and  $B$  form a complete set of observables if:

1.  $[A, B] = AB - BA = 0$

2. for a pair of eigenvalues  $a_i$  of  $A$  and  $b_i$  of  $B$  there exists only one eigenvector  $|i\rangle$  (no further degeneracy)

Property 1. implies that there exists a basis of simultaneous eigenvectors  $A|i\rangle = a_i|i\rangle$ ,  $B|i\rangle = b_i|i\rangle$ .

The definition can be generalised for more than two observables.

Example: Hydrogen atom

Hamilton operator  $H$ , angular momentum  $\vec{L}^2$  and its component  $L_3$  form a complete set.

\* Preparation of a quantum system in a definite state:

$$|\psi\rangle \xrightarrow[\substack{\text{measurement of} \\ \text{a complete set} \\ A, B, \dots, M}]{\text{filters values } (a_i, b_i, \dots, m_i)}$$

⇒ the system is afterwards in the prepared state  $|a_i b_i \dots m_i\rangle$

⇒ collapse of the wave function  $|\psi\rangle$ .

### 1.2.2. Position and momentum

Examples of observables with a continuum spectrum (range of values)

For a non-relativistic particle there are position operators

$$\vec{\hat{x}} = (\hat{x}_1, \hat{x}_2, \hat{x}_3)$$

and associated continuum eigenstates  $|\vec{y}\rangle$  with

$$\vec{\hat{x}} |\vec{y}\rangle = \vec{y} |\vec{y}\rangle, \quad \hat{x}_j |\vec{y}\rangle = y_j |\vec{y}\rangle, \quad j \in 1, 2, 3, \quad y_j \in \mathbb{R}.$$

Spectrum of  $\hat{x}_j = \mathbb{R}$ .

$$\hat{x}_j^\dagger = \hat{x}_j, \quad \text{Hermitian.}$$

\*  $\hat{x}_1, \hat{x}_2, \hat{x}_3$  are simultaneously diagonal in the basis  $|\vec{y}\rangle$

$$\Rightarrow [\hat{x}_i, \hat{x}_j] = 0$$

\* 3 components can be measured simultaneously

\* eigenvectors are indexed by the eigenvalue ( $\vec{y}$ )

$\{\hat{x}_1, \hat{x}_2, \hat{x}_3\}$  is a complete set

\* continuum states:  $\langle \vec{y} | \vec{z} \rangle = \delta^{(3)}(\vec{y} - \vec{z}) = \prod_{i=1}^3 \delta(y_i - z_i)$

\* completeness:  $\int d^3 y |\vec{y}\rangle \langle \vec{y}| = \mathbb{1}$

$$|\psi\rangle = \int d^3y |\vec{y}\rangle \langle \vec{y} | \psi \rangle$$

$\langle \vec{y} | \psi \rangle =: \psi(\vec{y})$  : wave function, Schrödinger representation

\* isomorphism  $\mathcal{H} \cong \mathcal{L}^2(\mathbb{R}^3)$  :

$$\langle \varphi | \psi \rangle = \langle \varphi | \underbrace{\int d^3y |\vec{y}\rangle \langle \vec{y} |}_{\mathbb{1}} | \psi \rangle = \int d^3y \varphi^*(\vec{y}) \psi(\vec{y})$$

$$\langle \psi | \psi \rangle < \infty \Leftrightarrow \int d^3y |\psi(\vec{y})|^2 < \infty$$

\* action of the position operator in the Schrödinger representation:

$$|\psi\rangle \longmapsto \hat{x}_j |\psi\rangle$$

$$\psi(\vec{y}) = \langle \vec{y} | \psi \rangle \longmapsto \langle \vec{y} | \hat{x}_j | \psi \rangle = y_j \langle \vec{y} | \psi \rangle = y_j \psi(\vec{y})$$

multiplication operator. Interpretation:

$$\langle \psi | \hat{x}^2 | \psi \rangle = \int d^3y \langle \psi | \vec{y} \rangle \langle \vec{y} | \hat{x}^2 | \psi \rangle = \int d^3y \vec{y} |\langle \vec{y} | \psi \rangle|^2$$

$$g(\vec{y}) = |\langle \vec{y} | \psi \rangle|^2$$

$g(\vec{y}) d^3y$  = probability that the particle is in a small volume  $d^3y$  centered at  $\vec{y}$

$$\int d^3y g(\vec{y}) = \langle \psi | \psi \rangle = 1$$

Observable:  $P_{\Delta V} = \int_{\Delta V} d^3x |\vec{x}\rangle \langle \vec{x}|$

$$\langle \psi | P_{\Delta V} | \psi \rangle = \int_{\Delta V} d^3x g(\vec{x}) \quad \text{probability to}$$

find the particle in a volume  $\Delta V$

$$\approx g(\vec{x}) \underbrace{|\Delta V|}_{\text{volume}} \quad \text{if } |\Delta V| \text{ small}$$

and  $g$  smooth

Preparation of a state: localize a particle in  $\Delta V$  through measurement  $\Rightarrow$  collapse of  $|\psi\rangle$  to  $\hat{P}_{\Delta V} |\psi\rangle =$  state after the measurement ⑧

Momentum operator  $\hat{\vec{p}}$ :

$$\langle \vec{x} | \hat{p}_j | \psi \rangle = \frac{\hbar}{i} \frac{\partial}{\partial x_j} \langle \vec{x} | \psi \rangle = \frac{\hbar}{i} \frac{\partial}{\partial x_j} \psi(\vec{x})$$

$$\hbar = \frac{h}{2\pi}, \quad h = \text{Planck's constant} = 6.6262 \times 10^{-34} \text{ J}\cdot\text{s}$$

$\hat{p}_j$  is Hermitian:

$$\begin{aligned} \langle \vec{x} | \hat{p}_j^\dagger | \psi \rangle &= \langle \psi | \hat{p}_j | \vec{x} \rangle^* = \left[ \int d^3y \langle \psi | y \rangle \frac{\hbar}{i} \frac{\partial}{\partial y_j} \langle y | \vec{x} \rangle \right]^* \quad \leftarrow \text{partial integration} \\ &= \left[ \int d^3y \left( -\frac{\hbar}{i} \frac{\partial}{\partial y_j} \langle \psi | y \rangle \right) \langle y | \vec{x} \rangle \right]^* \\ &= \int d^3y \langle \vec{x} | y \rangle \frac{\hbar}{i} \frac{\partial}{\partial y_j} \langle y | \psi \rangle = \langle \vec{x} | \hat{p}_j | \psi \rangle \end{aligned}$$

Definition of  $\hat{p}_j$  implies  $[\hat{p}_j, \hat{x}_k] = \frac{\hbar}{i} \delta_{jk}$ :

$$\begin{aligned} \langle \vec{y} | \hat{p}_j \hat{x}_k | \psi \rangle &= \frac{\hbar}{i} \frac{\partial}{\partial y_j} \langle \vec{y} | \hat{x}_k | \psi \rangle = \frac{\hbar}{i} \frac{\partial}{\partial y_j} y_k \langle \vec{y} | \psi \rangle \\ &= \frac{\hbar}{i} \delta_{jk} \langle \vec{y} | \psi \rangle + y_k \frac{\hbar}{i} \frac{\partial}{\partial y_j} \langle \vec{y} | \psi \rangle \\ &= \langle \vec{y} | \frac{\hbar}{i} \delta_{jk} + \hat{x}_k \hat{p}_j | \psi \rangle \end{aligned}$$

- operator identity

-  $\hat{p}_1, \hat{x}_1$  cannot be simultaneously measured

Eigenstates of  $\hat{\vec{p}}$ :  $\hat{p}_j | \vec{q} \rangle = q_j | \vec{q} \rangle$ :

$$\langle \vec{x} | \hat{p}_j | \vec{q} \rangle = q_j \langle \vec{x} | \vec{q} \rangle \Leftrightarrow \frac{\hbar}{i} \frac{\partial}{\partial x_j} \langle \vec{x} | \vec{q} \rangle = q_j \langle \vec{x} | \vec{q} \rangle$$

$$\Rightarrow \langle \vec{x} | \vec{q} \rangle = N \cdot e^{\frac{i}{\hbar} \vec{q} \cdot \vec{x}}$$



$$\langle \vec{q}' | \vec{q} \rangle = \int d^3x \langle \vec{q}' | \vec{x} \rangle \langle \vec{x} | \vec{q} \rangle = |N|^2 \int d^3x e^{\frac{i}{\hbar} (\vec{q}' - \vec{q}) \cdot \vec{x}} \quad (9)$$

$$= \delta^{(3)}(\vec{q}' - \vec{q}) \quad \text{if } N = (2\pi\hbar)^{-\frac{3}{2}}$$

$$\rightarrow \langle \vec{x} | \vec{q} \rangle = \frac{1}{(2\pi\hbar)^{3/2}} e^{\frac{i}{\hbar} \vec{q} \cdot \vec{x}}$$

$$\left. \begin{aligned} \langle \vec{x} | \psi \rangle &= \psi(\vec{x}) \\ \langle \vec{q} | \psi \rangle &=: \tilde{\psi}(\vec{q}) \end{aligned} \right\} \psi \text{ and } \tilde{\psi} \text{ are related by Fourier transform}$$

$$\tilde{\psi}(\vec{q}) = \int d^3x \langle \vec{q} | \vec{x} \rangle \langle \vec{x} | \psi \rangle = \frac{1}{(2\pi\hbar)^{3/2}} \int d^3x \psi(\vec{x}) e^{-\frac{i}{\hbar} \vec{q} \cdot \vec{x}}$$

### 1.2.3 Heisenberg Uncertainty Principle

discrete spectrum;  $\langle \psi | \psi \rangle = 1$ :

$$p_i = |\langle i | \psi \rangle|^2 = \text{probability to measure the value } a_i \text{ for the observable } A = \sum_i a_i |i\rangle \langle i|$$

Bemerkung: in case of degenerate orthonormal states  $|i, k\rangle$  with the same eigenvalue  $a_i$ :  $p_i = \sum_k |\langle i, k | \psi \rangle|^2$

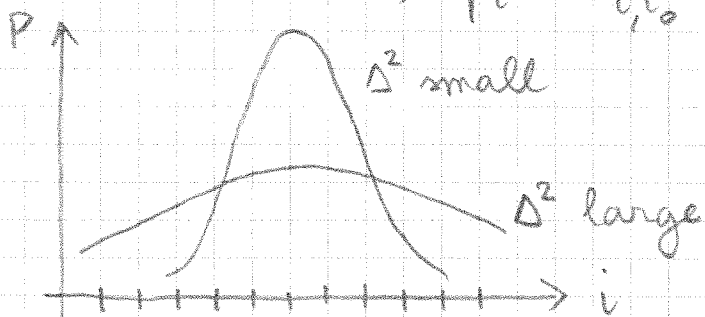
Mean value ("expectation value"):  $\bar{a} = \sum_i p_i a_i = \langle \psi | A | \psi \rangle = \langle A \rangle_{|\psi\rangle}$

Uncertainty, defined as the root mean square deviation:

$$\Delta_{A, |\psi\rangle}^2 = \sum_i p_i (a_i - \bar{a})^2 \geq 0$$

$$\Delta_{A, |\psi\rangle}^2 = 0 \Rightarrow p_i (a_i - \bar{a})^2 = 0 \quad \forall i$$

$$\Rightarrow p_i = \delta_{i, i_0} \Rightarrow |\psi\rangle = (\text{phase}) \cdot |i_0\rangle$$



$$\begin{aligned}
\Delta_{A,|\psi\rangle}^2 &= \sum_i p_i (a_i - \bar{a})^2 = \sum_i |\langle i|\psi\rangle|^2 (a_i - \bar{a})^2 \\
&= \langle\psi|\left[\sum_i |i\rangle\langle i| (a_i - \bar{a})^2\right]|\psi\rangle \\
&= \langle\psi|(A - \bar{a})^2|\psi\rangle \\
&= \|(A - \bar{a})|\psi\rangle\|^2 = \text{norm squared of } (A - \bar{a})|\psi\rangle \\
&= \langle\psi|A^2 - 2\bar{a}A + \bar{a}^2|\psi\rangle \\
&= \langle\psi|A^2|\psi\rangle - (\langle\psi|A|\psi\rangle)^2
\end{aligned}$$

Generalized version of the Heisenberg uncertainty principle:  
 when  $[A, B] = iC$  ( $\Rightarrow C = C^\dagger$ ) for observables  $A, B$ ,  
 then the following holds for any normed state  $|\psi\rangle$ ,  $\langle\psi|\psi\rangle = 1$ :

$$\Delta_{A,|\psi\rangle}^2 \cdot \Delta_{B,|\psi\rangle}^2 \geq \frac{1}{4} (\langle\psi|C|\psi\rangle)^2$$

Proof: see exercises.

Define  $a = A - \bar{a}$ ,  $a = a^\dagger$ ;  $b = B - \bar{b}$ ,  $b = b^\dagger$

Schwarz inequality: for any two states  $|x\rangle$  and  $|x'\rangle$

$$|\langle x'|x\rangle|^2 \leq \langle x'|x'\rangle \langle x|x\rangle$$

Let  $|x\rangle = b|\psi\rangle$ ,  $|x'\rangle = a|\psi\rangle$ :

$$|\langle\psi|ab|\psi\rangle|^2 \leq \langle\psi|a^2|\psi\rangle \langle\psi|b^2|\psi\rangle = \Delta_{A,|\psi\rangle}^2 \cdot \Delta_{B,|\psi\rangle}^2$$

$$\begin{aligned}
|\langle\psi|ab|\psi\rangle|^2 &= \langle\psi|ab|\psi\rangle \cdot \langle\psi|ab|\psi\rangle^* = \langle\psi|ab|\psi\rangle \cdot \langle\psi|\underbrace{ba}|\psi\rangle \\
&= (ab)^\dagger
\end{aligned}$$

$$\begin{aligned}
ab &= \frac{1}{2} \{a, b\} + \frac{1}{2} [a, b] \quad (\{a, b\} = ab + ba \text{ anti-commutator}) \\
ba &= \frac{1}{2} \{a, b\} - \frac{1}{2} [a, b]
\end{aligned}$$

$$\begin{aligned}
|\langle\psi|ab|\psi\rangle|^2 &= \langle\psi|\frac{1}{2}\{a, b\} + \frac{1}{2}[a, b]|\psi\rangle \cdot \langle\psi|\frac{1}{2}\{a, b\} - \frac{1}{2}[a, b]|\psi\rangle \\
&= (\langle\psi|\frac{1}{2}\{a, b\}|\psi\rangle)^2 - (\langle\psi|\frac{1}{2}[a, b]|\psi\rangle)^2
\end{aligned}$$

$$|\langle \psi | ab | \psi \rangle|^2 = \frac{1}{4} \underbrace{(\langle \psi | \{a, b\} | \psi \rangle)^2}_{\text{real, } \geq 0} + \frac{1}{4} \underbrace{(\langle \psi | i[a, b] | \psi \rangle)^2}_{\text{real, } \geq 0}$$

since  $\{a, b\}$  and  $i[a, b]$  are Hermitian.

$$i[a, b] = i[(A - \bar{a}), (B - \bar{b})] = i[A, B] = -C$$

$$\Rightarrow \Delta_{A, |\psi\rangle}^2 \cdot \Delta_{B, |\psi\rangle}^2 \geq |\langle \psi | ab | \psi \rangle|^2 \geq \frac{1}{4} (\langle \psi | C | \psi \rangle)^2$$

Special case:  $A = \hat{x}_k$ ,  $B = \hat{p}_k$ ,  $[\hat{x}_k, \hat{p}_k] = i\hbar$  ( $C = \hbar \cdot 1$ )

$$\Rightarrow \Delta_{\hat{x}_k, |\psi\rangle}^2 \cdot \Delta_{\hat{p}_k, |\psi\rangle}^2 \geq \frac{\hbar^2}{4}$$

\*  $\hat{x}_k, \hat{p}_k$  cannot be simultaneously measured with arbitrary precision

\*  $\hat{x}_k$ -eigenstate  $\Rightarrow \hat{p}_k$  completely undetermined and vice-versa

### 1.3 Dynamics (time evolution)

Hamilton operator  $H = H^\dagger$  determines the time evolution through Schrödinger's equation:

$$i\hbar \frac{d}{dt} |\psi(t)\rangle = H |\psi(t)\rangle$$

\* initial condition  $|\psi(t_0)\rangle$  through preparation

\*  $|\psi(t)\rangle$  unique solution

Time evolution of the expectation value of an observable  $A$ :

$$\langle A \rangle_{|\psi(t)\rangle} = \frac{\langle \psi(t) | A | \psi(t) \rangle}{\langle \psi(t) | \psi(t) \rangle}$$

$$\frac{d}{dt} |\psi(t)\rangle = -\frac{i}{\hbar} H |\psi(t)\rangle \longrightarrow \langle \psi(t) | \frac{d}{dt} H^\dagger = \langle \psi(t) | \frac{i}{\hbar} H$$

$$\Rightarrow \frac{d}{dt} \langle \psi(t) | \psi(t) \rangle = \frac{i}{\hbar} \langle \psi(t) | H | \psi(t) \rangle - \frac{i}{\hbar} \langle \psi(t) | H | \psi(t) \rangle = 0$$

$$\begin{aligned} \frac{d}{dt} \langle \psi(t) | A | \psi(t) \rangle &= \frac{i}{\hbar} \langle \psi(t) | H A | \psi(t) \rangle - \frac{i}{\hbar} \langle \psi(t) | A H | \psi(t) \rangle \\ &\quad + \langle \psi(t) | \frac{\partial A}{\partial t} | \psi(t) \rangle \\ &= -\frac{i}{\hbar} \langle \psi(t) | [A, H] | \psi(t) \rangle + \langle \psi(t) | \frac{\partial A}{\partial t} | \psi(t) \rangle \end{aligned}$$

Choosing  $\langle \psi(t_0) | \psi(t_0) \rangle = 1 \Rightarrow \langle \psi(t) | \psi(t) \rangle = 1 \quad \forall t \Rightarrow$

$$\frac{d}{dt} \langle A \rangle_{\psi(t)} = -\frac{i}{\hbar} \langle \psi(t) | [A, H] | \psi(t) \rangle + \langle \psi(t) | \frac{\partial A}{\partial t} | \psi(t) \rangle$$

If  $[A, H] = 0$  and  $\frac{\partial A}{\partial t} = 0$  then  $\langle A \rangle_{\psi(t)}$  is constant and  $A$  is called a conserved quantity.

Stationary states

If  $\frac{\partial H}{\partial t} = 0$  :  $H = \sum_n E_n |n\rangle \langle n|$  (discrete spectrum)

$E_n$  : allowed energy values

(can be generalised to continuum or mixed spectrum)

$$i\hbar \frac{d}{dt} |\psi(t)\rangle = H |\psi(t)\rangle \quad \frac{\partial H}{\partial t} = 0 \Rightarrow |\psi(t)\rangle = e^{-\frac{i}{\hbar} H (t-t_0)} |\psi(t_0)\rangle$$

$$|\psi(t)\rangle = \sum_n \underbrace{e^{-\frac{i}{\hbar} E_n (t-t_0)}}_{\text{phases}} |n\rangle \langle n | \psi(t_0) \rangle$$

Determination of  $\{|n\rangle, E_n\} \Leftrightarrow$  solution of dynamics

If  $|\psi(t_0)\rangle = |m\rangle$  (energy eigenstate), then

$$|\psi(t)\rangle = e^{-\frac{i}{\hbar} E_m (t-t_0)} |m\rangle \quad \text{stationary state} \\ \text{(constant physical state)}$$

If  $\frac{\partial H}{\partial t} = 0$ ,  $\frac{\partial A}{\partial t} = 0$  and  $[H, A] = 0$  there exists a basis of simultaneous eigenvectors

$$A |n\rangle = a_n |n\rangle \quad \text{and} \quad H |n\rangle = E_n |n\rangle.$$

If  $|\psi(t_0)\rangle = |m\rangle$  then  $|\psi(t)\rangle$  remains eigenstate of  $A$  and  $H$  with eigenvalues  $(a_m, E_m)$ .  $a_n$  are called "good quantum numbers"

Schrödinger's equation in the Schrödinger representation for a non-relativistic particle:

$$H = \frac{\vec{p}^2}{2m} + V(\vec{x}) \quad (\text{correspondence principle})$$

$$\psi(\vec{x}, t) = \langle \vec{x} | \psi(t) \rangle$$

$$i\hbar \langle \vec{x} | \frac{d}{dt} |\psi(t)\rangle = \langle \vec{x} | \left[ \frac{\vec{p}^2}{2m} + V(\vec{x}) \right] |\psi(t)\rangle$$

$$i\hbar \frac{\partial}{\partial t} \psi(\vec{x}, t) = \left[ -\frac{\hbar^2}{2m} \Delta + V(\vec{x}) \right] \psi(\vec{x}, t)$$

$$\Delta = \vec{\nabla}^2 = \frac{\partial^2}{\partial x_1^2} + \frac{\partial^2}{\partial x_2^2} + \frac{\partial^2}{\partial x_3^2} \quad \text{Laplace operator}$$

So far:

- a) Schrödinger picture:
- states are time-dependent
  - observables are time-independent (operators)

also possible:

- b) Heisenberg picture:
- states time-independent
  - observables time-dependent
  - same time-dependence of expectation values

- c) Dirac or interaction picture:

$$H = \underbrace{H_0}_{\text{time-indep.}} + \underbrace{H_1(t)}_{\text{(time-dep.) perturbation ("interaction")}}$$

Mixed form: time-dependence of

- states determined from  $H_1$
- observables determined from  $H_0$

# 1.4 Classic problems

Exactly solvable potentials  $V(\vec{x}) \propto \vec{x}^2, \frac{1}{|\vec{x}|}$

## 1.4.1 The Harmonic Oscillator

$\vec{x} \rightarrow x$ , one-dimensional:

$$H = \frac{\hat{p}^2}{2m} + \frac{1}{2} m \omega^2 \hat{x}^2$$

$$E_n = (n + \frac{1}{2}) \hbar \omega$$

$$\langle x | n \rangle = d_n H_n(\xi) e^{-\frac{1}{2} \xi^2}, \quad \xi = \sqrt{\frac{m\omega}{\hbar}} x$$

normalization  $\nearrow$        $\nwarrow$  Hermite polynomial

- solution of differential equation + boundary conditions (normed states!  $\Leftrightarrow$  exponentially falling solution)  
or

- algebraic method:

$$H = \hbar \omega (a^\dagger a + \frac{1}{2})$$

$$a = \frac{i}{\sqrt{2m\omega\hbar}} \hat{p} + \sqrt{\frac{m\omega}{2\hbar}} \hat{x}, \quad a^\dagger = \frac{-i}{\sqrt{2m\omega\hbar}} \hat{p} + \sqrt{\frac{m\omega}{2\hbar}} \hat{x}$$

$$[a, a^\dagger] = 1 \quad \text{commutation relation}$$

$$a |0\rangle = 0, \quad E_0 = \frac{1}{2} \hbar \omega \quad \text{ground state}$$

$$\left. \begin{aligned} |n\rangle &\propto a^\dagger |n-1\rangle \propto (a^\dagger)^n |0\rangle \\ &\quad \uparrow \text{creation operator} \\ |n-1\rangle &\propto a |n\rangle \\ &\quad \uparrow \text{annihilation operator} \end{aligned} \right\} \begin{array}{l} \text{lowering (a) and} \\ \text{raising (a}^\dagger\text{) operators} \end{array}$$

$|n\rangle$ : state with  $n$  quanta,  
energy =  $E_n = n \cdot \hbar \omega + \underbrace{\frac{1}{2} \hbar \omega}_{\text{zero point, ground state energy}}$

Remark: in three dimension the harmonic oscillator is

$$H = \frac{\vec{p}^2}{2m} + \frac{1}{2} m\omega^2 \vec{x}^2$$

$$= \sum_{i=1}^3 \left[ \frac{\hat{p}_i^2}{2m} + \frac{1}{2} m\omega^2 \hat{x}_i^2 \right] \quad \text{sum of three one-dimensional oscillators}$$

$$\Rightarrow E_n = E_{n_1} + E_{n_2} + E_{n_3} = \left( n_1 + n_2 + n_3 + \frac{3}{2} \right) \hbar\omega$$

$$\Psi_n(\vec{x}) = \Psi_{n_1}(x_1) \cdot \Psi_{n_2}(x_2) \cdot \Psi_{n_3}(x_3) \quad (\text{separable solution})$$

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$$\psi_n(\vec{x}) = \psi_{n_1}(x_1) \cdot \psi_{n_2}(x_2) \cdot \psi_{n_3}(x_3) \quad (\text{separable solution})$$

### 1.4.2 The Hydrogen Atom

$$H = \frac{\vec{p}^2}{2m} - \frac{e^2}{4\pi\epsilon_0} \frac{1}{|\vec{x}|}$$

$e = 1.60219 \cdot 10^{-19} \text{ C}$   
 $\epsilon_0 = 8.85419 \cdot 10^{-12} \frac{\text{C}}{\text{V}\cdot\text{m}}$

reduced mass  $m = m_e \left( 1 + \frac{m_e}{m_p} \right)^{-1} \approx m_e$

since  $\frac{m_p}{m_e} = 1836.11$

Schrödinger's equation for stationary states  $\psi(\vec{x})e^{-iEt}$ :

$$\left( -\frac{\hbar^2}{2m} \Delta - \frac{e^2}{4\pi\epsilon_0} \frac{1}{|\vec{x}|} \right) \psi(\vec{x}) = E \psi(\vec{x})$$

Polar coordinates  $\psi(\vec{x}) \rightarrow \psi(r, \theta, \varphi)$ :

$$\Delta = \frac{1}{r} \frac{\partial^2}{\partial r^2} r - \underbrace{\frac{\vec{L}^2}{r^2 \hbar^2}}_{\text{centrifugal term}}$$

$\vec{L} = \vec{x} \times \vec{p}$  angular momentum, in polar coordinates:

$$-\frac{1}{\hbar^2} \vec{L}^2 = \frac{1}{\sin^2 \theta} \left[ \sin \theta \frac{\partial}{\partial \theta} \left( \sin \theta \frac{\partial}{\partial \theta} \right) + \frac{\partial^2}{\partial \varphi^2} \right]$$

$$L_3 = \frac{\hbar}{i} \frac{\partial}{\partial \varphi}$$

$\{H, L_3, \vec{L}^2\}$  form a complete set of observables:



$$\Psi_{n\ell m}(r, \theta, \varphi) = R_{n\ell}(r) Y_{\ell m}(\theta, \varphi) \quad n > \ell$$

$$\vec{L}^2 Y_{\ell m} = \ell(\ell+1) \hbar^2 Y_{\ell m}$$

$$L_3 Y_{\ell m} = m \hbar Y_{\ell m} \quad m = -\ell, -\ell+1, \dots, \ell-1, \ell$$

$Y_{\ell m}(\theta, \varphi)$  : spherical harmonics

complete orthonormal set on the unit sphere  $(\theta, \varphi)$  :

$$(f, g) = \int d\Omega f^* g, \quad \int d\Omega = \int_0^{2\pi} d\varphi \int_0^\pi d\theta \sin\theta$$

$$(Y_{\ell m}, Y_{\ell' m'}) = \delta_{\ell\ell'} \delta_{mm'}$$

$$Y_{\ell m} = \sqrt{\frac{2\ell+1}{4\pi} \frac{(\ell-m)!}{(\ell+m)!}} \underbrace{P_\ell^m(\cos\theta)}_{\text{associated Legendre polynomials,}} e^{im\varphi}$$

polynomials of degree  $\ell$  in  $\sin\theta$  and  $\cos\theta$

- radial function for bound states :

$E < 0$  , discrete spectrum

$$R_{n\ell} = \underbrace{d_{n\ell}}_{\text{normalization}} e^{-\frac{r}{na_B}} \left(\frac{2r}{na_B}\right)^\ell \underbrace{L_{n-\ell}^{2\ell+1}\left(\frac{2r}{na_B}\right)}_{\text{associated Laguerre polynomials}}$$

degree  $n-\ell-1 \geq 0$

$$a_B = \frac{4\pi\epsilon_0 \hbar^2}{e^2 m} = 0.529 \text{ \AA} \quad (1 \text{ \AA} = 10^{-10} \text{ m}) \quad \text{Bohr radius}$$

$$H \Psi_{n\ell m} = E_n \Psi_{n\ell m}$$

$$E_n = \frac{-E_R}{n^2} \quad \text{Balmer formula}$$

$n = 1, 2, 3, \dots$  principal quantum number

degeneracy =  $n^2$

$$E_R = \frac{\hbar^2}{2m a_B^2} = 13.605 \text{ eV} \quad \text{Rydberg energy}$$

Transitions  $n \rightarrow m$ :  $\hbar \nu_{nm} = -E_R \left( \frac{1}{n^2} - \frac{1}{m^2} \right) \quad n > m$

frequency of the emitted photon = quantum of light

This formula explains the observed spectral lines.

- scattering states for  $E > 0$

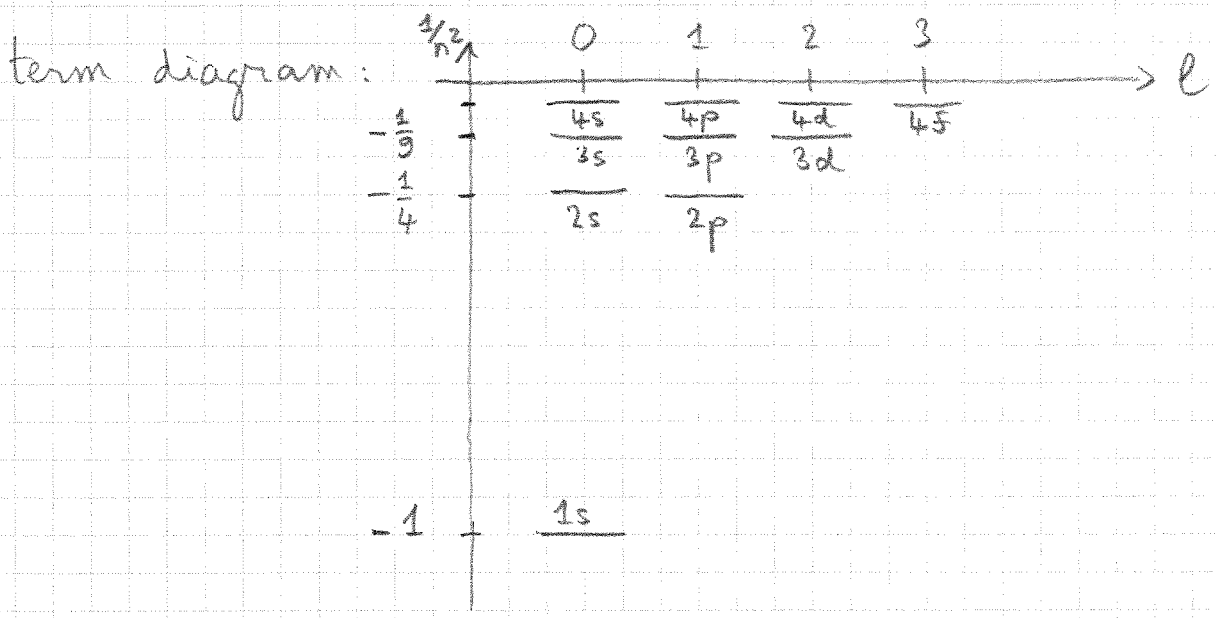
Remarks:

- symmetry of the  $\frac{1}{r}$  potential is actually  $SO(4)$ , not only  $SO(3)$   
Conservation of the Runge-Lenz-Sauli vector  $\vec{A}$ , explains  $n^2$  degeneracy of  $E_n$

- nomenclature:

$n \leftrightarrow$  electron shell:  $n=1$  ("K shell"),  $n=2$  ("L shell"),  
 $n=3$  ("M shell"), ...

$l = (0, 1, 2, 3, \dots) \leftrightarrow (s, p, d, f, \dots)$  - orbitals



- atomic nucleus with positive charge  $Z \cdot e$   
 $\Rightarrow$  replace the argument  $r$  of  $R_{ne}(r)$  by  $Z \cdot r$

## 1.5 Composite systems

State spaces  $\mathcal{H}_1$  and  $\mathcal{H}_2$

$$\mathcal{H} = \underbrace{\mathcal{H}_1 \otimes \mathcal{H}_2}_{\text{direct product space}}, \quad |t_1\rangle \otimes |t_2\rangle \in \mathcal{H}$$

$$\begin{aligned} \{|i\rangle\} & \text{ CON on } \mathcal{H}_1 \\ \{|\alpha\rangle\} & \text{ CON on } \mathcal{H}_2 \end{aligned} \Rightarrow \{|i\rangle \otimes |\alpha\rangle\} \text{ CON on } \mathcal{H}$$

$$|t\rangle \in \mathcal{H}: \quad |t\rangle = \sum_{i,\alpha} C_{i\alpha} |i\rangle \otimes |\alpha\rangle$$

$$\dim(\mathcal{H}) = \dim(\mathcal{H}_1) \cdot \dim(\mathcal{H}_2)$$

Operators:  $A_i: \mathcal{H}_i \rightarrow \mathcal{H}_i$

$$\begin{aligned} A_1 & \longrightarrow A_1 \otimes \mathbb{1} \\ A_2 & \longrightarrow \mathbb{1} \otimes A_2 \end{aligned} \left. \vphantom{\begin{aligned} A_1 \\ A_2 \end{aligned}} \right\} \begin{array}{l} \text{operators on } \mathcal{H}, \text{ for example} \\ \text{observables of the composite system} \end{array}$$

$$(A \otimes B) |t\rangle = \sum_{i,\alpha} C_{i\alpha} (A|i\rangle) \otimes (B|\alpha\rangle)$$

Example: 2 particles:

$$\text{basis } \{|\vec{x}_1\rangle \otimes |\vec{x}_2\rangle = |\vec{x}_1 \vec{x}_2\rangle\}$$

$$\psi(\vec{x}_1, \vec{x}_2) = \langle \vec{x}_1 \vec{x}_2 | t \rangle$$

$$\begin{aligned} H &= \frac{\vec{p}_1^2 \otimes \mathbb{1}}{2m_1} + \frac{\mathbb{1} \otimes \vec{p}_2^2}{2m_2} + V(\vec{x}_1 \otimes \mathbb{1} - \mathbb{1} \otimes \vec{x}_2) \\ &\triangleq \frac{\vec{p}_1^2}{2m_1} + \frac{\vec{p}_2^2}{2m_2} + V(\vec{x}_1 - \vec{x}_2) \end{aligned}$$

Example: particle with spin

$$\mathcal{H}_1 \cong \mathcal{L}^2(\mathbb{R}^3), \quad \mathcal{H}_2 = \mathbb{C}^2$$

$$\text{basis } \{|\vec{x}\rangle \otimes |\alpha\rangle = |\vec{x}\alpha\rangle\}, \quad \alpha = \pm$$

$$\psi_\alpha(\vec{x}) = \langle \vec{x}\alpha | t \rangle$$

Hamilton operator for a particle with spin  $\rightarrow$  see Chapter 3  
(Pauli equation)

## 2. The Quantum Theory of the Angular Momentum

### 2.1 Symmetry and Conserved Observable

Consider an active transformation  $\mathcal{G}$  of the system:

$$|\psi\rangle \xrightarrow{\mathcal{G}} |\psi'\rangle$$

active  $\Leftrightarrow \mathcal{G}$  transforms the system but not the system of coordinates

$$|\psi\rangle \Leftrightarrow \text{ray } \{e^{i\alpha} |\psi\rangle, \alpha \in \mathbb{R}\}, \langle \psi | \psi \rangle = 1$$

$\mathcal{G}$  represent a symmetry,  $\mathcal{G}$  must preserve probabilities:

if  $|\psi\rangle \xrightarrow{\mathcal{G}} |\psi'\rangle$  and  $|\varphi\rangle \xrightarrow{\mathcal{G}} |\varphi'\rangle$  then

$$|\langle \varphi' | \psi' \rangle|^2 = |\langle \varphi | \psi \rangle|^2$$

This is a condition on rays.

A transformation  $\mathcal{G}$  is a symmetry of the dynamics if

$$|\psi(t_0)\rangle \xrightarrow{\mathcal{G}} |\psi'(t_0)\rangle$$

$$\downarrow$$

$$|\psi(t)\rangle$$

$$\xrightarrow{\mathcal{G}} |\psi'(t)\rangle$$

yields the same result

$$\downarrow: \text{solve its } \frac{d}{dt} |\psi\rangle = H |\psi\rangle$$

Theorem of Wigner about the representation of symmetries:  
there are just two ways to represent such a transformation  $\mathcal{G}$ :

a) by a linear unitary operator:

$$\langle \varphi' | \psi' \rangle = \langle \varphi | \psi \rangle$$

$$|\psi'\rangle = U |\psi\rangle, \quad |\varphi'\rangle = U |\varphi\rangle, \quad U^\dagger U = \mathbb{1} = U U^\dagger$$

$$U (\alpha |\psi\rangle + \beta |\chi\rangle) = \alpha U |\psi\rangle + \beta U |\chi\rangle, \quad \alpha, \beta \in \mathbb{C}$$

b) by an antilinear antiunitary operator:

$$\langle \varphi' | \psi' \rangle = \langle \varphi | \psi \rangle^*$$

antilinear:  $U(\alpha|\psi\rangle + \beta|\chi\rangle) = \alpha^* U|\psi\rangle + \beta^* U|\chi\rangle$

for antiunitary operator the definition of adjoint is changed to

$$(\langle U^\dagger \varphi |, |\psi\rangle) = (\langle \varphi |, U|\psi\rangle)^* \Rightarrow U^\dagger U = \mathbb{1}$$

Symmetries represented by antiunitary operators all involve a change in the direction of time flow (time reversal)

In the following: unitary transformations

$$U: \mathcal{H} \rightarrow \mathcal{H}, |\psi\rangle \xrightarrow{U} |\psi'\rangle = U|\psi\rangle$$

If  $U(\alpha)$  depends on  $\alpha \in \mathbb{R}$  with  $U(0) = \mathbb{1}$  (for example  $\alpha =$  rotation angle) then:

$$\begin{aligned} |\psi'(t)\rangle &= e^{-\frac{i}{\hbar} Ht} |\psi'(0)\rangle = e^{-\frac{i}{\hbar} Ht} U(\alpha) |\psi(0)\rangle \quad \rightarrow \downarrow \\ &= U(\alpha) e^{-\frac{i}{\hbar} Ht} |\psi(0)\rangle \quad \downarrow \rightarrow \end{aligned}$$

same result  $\Leftrightarrow [U(\alpha), H] = 0 \quad \forall \alpha$

If  $\alpha$  is infinitesimally small:

$$U(\alpha) = \mathbb{1} + \alpha \cdot T + \mathcal{O}(\alpha^2); \quad T = \left. \frac{dU}{d\alpha} \right|_{\alpha=0}$$

$$\begin{aligned} U U^\dagger &= (\mathbb{1} + \alpha \cdot T) (\mathbb{1} + \alpha \cdot T^\dagger) + \mathcal{O}(\alpha^2) \\ &= \mathbb{1} + \alpha (T + T^\dagger) + \mathcal{O}(\alpha^2) \end{aligned}$$

$$U U^\dagger = \mathbb{1} \Rightarrow T = -T^\dagger$$

We define  $A = iT$ ,  $A = A^\dagger$ .  $A$  is known as the generator of the symmetry

Symmetry of the dynamics  $\Leftrightarrow [A, H] = 0 \quad (\Leftrightarrow \frac{d}{d\alpha} [U(\alpha), H] = 0)$

$\Rightarrow A$  is a conserved observable

Many if not all of the operators representing observables in quantum mechanics are the generators of symmetries.  
Measurement apparatus for  $A$  is still to be found!

Compare to Noether Theorem of classical mechanics:

$$[A, H] = 0 \longrightarrow \{A, H\} = 0 \text{ Poisson bracket}$$

$$\frac{dH}{dx} = \{A, H\} : \text{A is generator of the symmetry transformation}$$

$$\frac{dA}{dt} = \{H, A\} : \text{Hamilton function H is generator of time evolution}$$

$$\{H, A\} = \sum_i \frac{\partial H}{\partial p_i} \frac{\partial A}{\partial q_i} - \frac{\partial A}{\partial p_i} \frac{\partial H}{\partial q_i}$$

## 2.1.1 Translations

Example:  $n$  particles,  $\mathcal{H} = \mathcal{H}_1 \otimes \mathcal{H}_2 \otimes \dots \otimes \mathcal{H}_n$

Basis  $|\vec{x}_1 \vec{x}_2 \dots \vec{x}_n\rangle$

Potential  $V = V(\vec{x}_i - \vec{x}_j)$  only depends on difference of coordinates

$\Rightarrow$  invariance under translation  $\vec{x}_i \rightarrow \vec{x}_i + \vec{a} \quad \forall i$

$$\text{Define } \langle \vec{x}_1 \dots \vec{x}_n | \psi_{\vec{a}} \rangle := \langle \vec{x}_1 - \vec{a} \dots \vec{x}_n - \vec{a} | \psi \rangle$$

$$\psi_{\vec{a}}(\vec{x}_1, \dots, \vec{x}_n) \quad \psi(\vec{x}_1 - \vec{a}, \dots, \vec{x}_n - \vec{a})$$

$$|\psi_{\vec{a}}\rangle =: U(\vec{a}) |\psi\rangle$$

Remarks:

1) sign: if  $|\psi\rangle$  has a peak at  $\vec{x}_j = 0$  then  $|\psi_{\vec{a}}\rangle$  is peaked at  $\vec{x}_j = \vec{a}$

$$\begin{aligned} 2) \text{ unitarity: } \langle \varphi | U^\dagger(\vec{a}) U(\vec{a}) | \psi \rangle &= \int d^3x_1 \dots d^3x_n \langle \varphi | U^\dagger(\vec{a}) | \vec{x}_1 \dots \vec{x}_n \rangle \langle \vec{x}_1 \dots \vec{x}_n | U(\vec{a}) | \psi \rangle \\ &= \int d^3x_1 \dots d^3x_n \langle \varphi_{\vec{a}} | \vec{x}_1 \dots \vec{x}_n \rangle \langle \vec{x}_1 \dots \vec{x}_n | \psi_{\vec{a}} \rangle \\ &= \int d^3x_1 \dots d^3x_n \langle \varphi | \vec{x}_1 - \vec{a} \dots \vec{x}_n - \vec{a} \rangle \langle \vec{x}_1 - \vec{a} \dots \vec{x}_n - \vec{a} | \psi \rangle = \langle \varphi | \psi \rangle \end{aligned}$$

by change of variables  $\vec{x}'_j = \vec{x}_j - \vec{a}$ ,  $d^3x'_j = d^3x_j$   
holds for all  $|\psi\rangle, |\varphi\rangle \Rightarrow U^\dagger(\vec{a}) U(\vec{a}) = \mathbb{1}$

analogously  $U(\vec{a}) U(\vec{a})^\dagger = \mathbb{1}$ :

$$|\varphi_{\vec{a}}\rangle = U(\vec{a}) |\varphi\rangle \Rightarrow |\varphi\rangle = U^\dagger(\vec{a}) |\varphi_{\vec{a}}\rangle \Rightarrow U^\dagger(\vec{a}) = U(-\vec{a})$$

$U^\dagger U = \mathbb{1}$

generator,  $\vec{a} = \vec{\epsilon}$  infinitesimal:

$$\langle \vec{x}_1 \dots \vec{x}_n | \psi_{\vec{\epsilon}} \rangle = \psi(\vec{x}_1 - \vec{\epsilon}, \dots, \vec{x}_n - \vec{\epsilon}) \stackrel{\text{Taylor expansion}}{=} \left( 1 - \vec{\epsilon} \cdot \sum_{k=1}^n \vec{\nabla}_{\vec{x}_k} \right) \psi(\vec{x}_1, \dots, \vec{x}_n)$$

$$= \langle \vec{x}_1 \dots \vec{x}_n | \left( 1 - \frac{i}{\hbar} \vec{\epsilon} \cdot \sum_{k=1}^n \vec{P}_k \right) |\psi\rangle$$

$$U(\vec{\epsilon}) \approx \mathbb{1} - \frac{i}{\hbar} \vec{\epsilon} \cdot \vec{P} \quad \left( \mathbb{1} \otimes \dots \otimes \vec{P}_k \otimes \dots \otimes \mathbb{1} \right)$$

$\vec{P} = \sum_{k=1}^n \vec{P}_k$  is the total momentum

kth particle

If the potential  $V$  is invariant under translations:  $V = V(\vec{x}_i - \vec{x}_j)$   
 $\Rightarrow [H, \vec{P}] = 0$  (please verify)

Finite translation:

$$U(\vec{a}) = \lim_{N \rightarrow \infty} \underbrace{U\left(\frac{\vec{a}}{N}\right)}_N^N = \lim_{N \rightarrow \infty} \left( 1 - \frac{i}{\hbar} \frac{\vec{a} \cdot \vec{P}}{N} \right)^N = e^{-\frac{i}{\hbar} \vec{a} \cdot \vec{P}}$$

N times translation by  $\frac{\vec{a}}{N}$

compare  $e^{a \frac{d}{dx}} f(x) = f(x) + a f'(x) + \frac{1}{2} a^2 f''(x) + \dots = f(x+a)$

Example:  $n=2$ ,  $V(\vec{x}_1 - \vec{x}_2)$

$$\psi(\vec{x}_1, \vec{x}_2) = e^{i\vec{k} \cdot \frac{m_1 \vec{x}_1 + m_2 \vec{x}_2}{m_1 + m_2}} \varphi(\vec{x}_1 - \vec{x}_2)$$

$$P_j \psi = \left[ (\hat{P}_1)_j + (\hat{P}_2)_j \right] \psi = \frac{\hbar}{i} \left( \frac{\partial}{\partial x_1} + \frac{\partial}{\partial x_2} \right) \psi = \hbar k_j \psi$$

$P_j \varphi = 0$

$P_1, P_2, P_3, H, \dots$  diagonal  $\Rightarrow$  center-of-mass motion is described by a plane wave

Translation eigenstate:  $U(\vec{a}) |\psi\rangle = e^{-i\vec{a}\cdot\vec{K}} |\psi\rangle$

Time translation:

it is a symmetry if  $\frac{\partial H}{\partial t} = 0$ ;  $H$  itself is the generator:

$$\begin{array}{ccc}
 |\psi(t_0)\rangle & \longrightarrow & |\psi(t_0+\tau)\rangle = \underbrace{e^{-\frac{i}{\hbar}\tau H}}_{U(\tau)} |\psi(t_0)\rangle \\
 \downarrow & & \downarrow \\
 |\psi(t)\rangle & \dashrightarrow & |\psi(t+\tau)\rangle
 \end{array}$$

Schrödinger equation for  $\frac{\partial H}{\partial t} = 0 \Rightarrow |\psi(t)\rangle = e^{-\frac{i}{\hbar}(t-t_0)H} |\psi(t_0)\rangle$   
 $H \Leftrightarrow$  energy, generator of time translation  $= \underbrace{U(t-t_0)}_{\text{time translation}} |\psi(t_0)\rangle$

### 2.1.2 Rotations

$$\vec{x} \mapsto R\vec{x}, \quad x_j \mapsto \sum_{k=1}^3 R_{jk} x_k \equiv R_{jk} x_k$$

$$RR^T = \mathbb{1}, \quad \det(R) = +1 \quad (\text{Einstein's summation convention})$$

$R \in SO(3)$  = special orthogonal group in 3 dimensions

Infinitesimal rotations:

$$\vec{x} \mapsto \vec{x} + \delta \vec{n} \times \vec{x} + O(\delta^2) : \vec{n} \text{ with } \vec{n}^2 = 1 : \text{rotation axis} \\
 \delta : \text{rotation angle}$$

$$x_i \mapsto x_i + \delta \epsilon_{ijk} n_j x_k = x_i + \delta n_j (T_j)_{ik} x_k \Leftrightarrow$$

$$\vec{x} \mapsto \vec{x} + \delta n_j T_j \vec{x}, \quad \{T_j\} : 3 \times 3 \text{ matrix generators} \\
 (T_j)_{ik} = \epsilon_{ijk}$$

$$R \approx \mathbb{1} + \delta \vec{n} \cdot \vec{T}$$



$$T_1 = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & -1 \\ 0 & 1 & 0 \end{pmatrix}, \quad T_2 = \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ -1 & 0 & 0 \end{pmatrix}, \quad T_3 = \begin{pmatrix} 0 & -1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \quad (24)$$

$(T_i)^T = -T_i$  antisymmetric  
 $\{T_i\}$  satisfy the commutation relations

$$[T_i, T_j] = \epsilon_{ijk} T_k$$

Proof:  $[T_i, T_j]_{mn} = \epsilon_{mir} \epsilon_{rjn} - \epsilon_{mjr} \epsilon_{rin} = \epsilon_{mir} \epsilon_{jnr} - (i \leftrightarrow j)$

$$= (\delta_{mj} \delta_{in} - \delta_{mn} \delta_{ij}) - (i \leftrightarrow j)$$

$$= \delta_{mj} \delta_{in} - \delta_{mi} \delta_{jn} = \epsilon_{kmn} \epsilon_{kji} = \epsilon_{ijk} \epsilon_{mkn}$$

$$= \epsilon_{ijk} (T_k)_{mn} \quad \square$$

Commutation relations characterizes the group of rotations.

Rotations induce transformations in the Hilbert space  $\mathcal{H}$ :

$$\langle \vec{x}_1 \dots \vec{x}_n | \Psi_R \rangle := \langle R^{-1} \vec{x}_1 \dots R^{-1} \vec{x}_n | \Psi \rangle \quad \mathcal{H}_1 \otimes \dots \otimes \mathcal{H}_n$$

infinitesimal  $\rightarrow \approx \langle (\vec{x}_1 - \delta \vec{n} \times \vec{x}_1) \dots (\vec{x}_n - \delta \vec{n} \times \vec{x}_n) | \Psi \rangle$

$$\rightarrow \approx \left( 1 - \sum_{k=1}^n \delta (\vec{n} \times \vec{x}_k) \cdot \vec{\nabla}_{\vec{x}_k} \right) \langle \vec{x}_1 \dots \vec{x}_n | \Psi \rangle$$

$$= \vec{n} \cdot \left( \vec{x}_k \times \vec{\nabla}_{\vec{x}_k} \right)$$

$$\Rightarrow |\Psi_R\rangle = U(R) |\Psi\rangle$$

$$= \left( 1 - \frac{i}{\hbar} \delta \vec{n} \cdot \sum_{k=1}^n \vec{x}_k \times \vec{p}_k \right) |\Psi\rangle$$

$(1 \otimes \dots \otimes \underbrace{\vec{x}_k \times \vec{p}_k}_{k\text{th factor (particle)}} \otimes \dots \otimes 1)$

infinitesimal:  $U(R) = 1 - \frac{i}{\hbar} \delta \vec{n} \cdot \vec{L}$

$\vec{L} = \sum_{k=1}^n \vec{x}_k \times \vec{p}_k$  is the total angular momentum

System of  $n$  particles has a rotation symmetry if  $V(\vec{x}_1, \dots, \vec{x}_n) = V(R\vec{x}_1, \dots, R\vec{x}_n)$ , then

$$[U(R), H] = 0 \text{ and}$$

$$[L_j, H] = 0 \quad j=1,2,3$$

$U(R)$  is unitary,  $\vec{L}$  is Hermitian ( $\hat{x}_k \hat{p}_l = \hat{p}_l \hat{x}_k$  if  $k \neq l$ )

Components of  $\vec{L}$  have the same commutation relations as  $i\hbar \vec{T}$ :

$$[i\hbar T_i, i\hbar T_j] = i\hbar \epsilon_{ijk} (i\hbar T_k)$$

$$[L_i, L_j] = i\hbar \epsilon_{ijk} L_k \quad (\text{please verify})$$

This is not by chance:  $U(R)$  = representation of  $SO(3)$  on  $\mathcal{H}$

$\{L_i\}$  cannot be simultaneously diagonal,  $[L_i, L_j] \neq 0$

$$[L_j, H] = 0 \Rightarrow [f(L_j), H] = 0 \quad (\text{only one component } j)$$

$\{L_3, \vec{L}^2 = L_1^2 + L_2^2 + L_3^2, H\}$  commute pairwise  $\Rightarrow$  have a common set of eigenvectors (see hydrogen atom)

## 2.2 Transformation of observables

active transformation  $\mathcal{B}: |\psi\rangle \xrightarrow{\mathcal{B}} |\psi'\rangle = U|\psi\rangle$

If the measurement apparatus ( $\leftrightarrow$  observables) remain unchanged:

$$\langle \psi | A | \psi \rangle \xrightarrow{\mathcal{B}} \langle \psi' | A | \psi' \rangle = \langle \psi | U^\dagger A U | \psi \rangle$$

For some set of observables  $\{A_i\}$ ,  $U^\dagger A_i U$  ( $\langle \psi | A_i | \psi' \rangle$ ) can be again expressed in terms of  $\{A_i\}$  ( $\{\langle \psi | A_i | \psi \rangle\}$ ).

Example:  $\vec{x}$ , rotations  $U \cong \mathbb{1} - \frac{i}{\hbar} \delta \vec{n} \cdot \vec{L}$ :

$$U^\dagger \hat{x}_j U \cong \left( \mathbb{1} + \frac{i}{\hbar} \delta \vec{n} \cdot \vec{L} \right) \hat{x}_j \left( \mathbb{1} - \frac{i}{\hbar} \delta \vec{n} \cdot \vec{L} \right)$$

$$\cong \hat{x}_j + \frac{i}{\hbar} \delta n_k [L_k, \hat{x}_j]$$

$$[L_k, \hat{x}_j] = \epsilon_{klm} [\hat{x}_l \hat{p}_m, \hat{x}_j] \underset{\substack{\text{components of } \hat{x} \text{ commute}}}{=} \epsilon_{klm} \hat{x}_l [\hat{p}_m, \hat{x}_j] = \frac{\hbar}{i} \epsilon_{klm} \delta_{mj} \hat{x}_l = \frac{\hbar}{i} \epsilon_{klj} \hat{x}_l$$

$$\Rightarrow U^\dagger \hat{x}_j U \cong \hat{x}_j + \delta n_k \epsilon_{klj} \hat{x}_l = (\mathbb{1} + \delta n_k T_k)_{je} \hat{x}_e$$

Finite transformations:

$$U^\dagger \hat{x}_j U = R_{je} \hat{x}_e$$

$\uparrow$  Hilbert space       $\uparrow$  SO(3) matrix

This means:  $\langle \psi | \hat{x}_j | \psi \rangle = R_{je} \langle \psi | \hat{x}_e | \psi \rangle$

"rotated position expectation values in the rotated state"

Definition:  $\vec{\hat{x}}$  is a vector operator.  $\vec{\hat{p}}$  too.

Dynamics is invariant under rotations  $\Leftrightarrow U^\dagger H U = H$

$\Leftrightarrow$  Hamiltonian operator  $H$  is a scalar operator.

Other (than scalar, vector) tensor operators exist.

## 2.3 Groups and Representations

Concept of a group, for example SO(3)

$$\vec{x} \mapsto R_1 \vec{x} \mapsto R_2 R_1 \vec{x}, \quad R_3 = R_2 \cdot R_1 \in SO(3)$$

rotation  $R_i$  is determined by axis  $\vec{n}_i$  ( $|\vec{n}_i|^2 = 1$ ) and angle  $\alpha_i$

$$\Rightarrow \left. \begin{aligned} \alpha_3 &= \alpha_3(\alpha_1, \alpha_2, \vec{n}_1 \cdot \vec{n}_2) \\ \vec{n}_3 &= \vec{n}_3(\alpha_1, \alpha_2, \vec{n}_1, \vec{n}_2) \end{aligned} \right\} \text{complicated functions}$$

$R^{-1} \in SO(3)$  with  $(-\alpha, \vec{n})$ ,  $\mathbb{1}_{3 \times 3} \in SO(3)$  identity

$SO(3)$  is an example of a Lie group.

$$\frac{\partial R}{\partial \alpha} \Big|_{\alpha=0} = \vec{n} \cdot \vec{T}, \quad R = \mathbb{1} + \alpha \vec{n} \cdot \vec{T} + O(\alpha^2)$$

$$\Rightarrow R = \lim_{N \rightarrow \infty} \left( R \left( \frac{\alpha}{N}, \vec{n} \right) \right)^N = \lim_{N \rightarrow \infty} \left( \mathbb{1} + \frac{\alpha}{N} \vec{n} \cdot \vec{T} \right)^N = e^{\alpha \vec{n} \cdot \vec{T}}$$

$$* R^T = e^{\alpha \vec{n} \cdot \vec{T}^T} = e^{-\alpha \vec{n} \cdot \vec{T}} = R^{-1}$$

$$* \det R = e^{\text{tr} \ln R} = e^{\text{tr}(\alpha \vec{n} \cdot \vec{T})} = 1$$

\* each element  $R \in SO(3)$  can be expressed this way

\*  $\{T_1, T_2, T_3\}$  are a basis of the Lie algebra  $so(3)$

\* commutation relations determine the group multiplication (Baker, Campbell, Hausdorff)

$$e^A \cdot e^B = e^C$$

$$C = A + B + \frac{1}{2} [A, B] + \frac{1}{42} \{ [A, [A, B]] + [B, [B, A]] \} + \dots$$

only commutators  $\Rightarrow C \in so(3)$

\* analogously for other Lie groups like  $SU(3)$  (quarks)

Representation: unitary, in Hilbert space:

$$R \mapsto U(R), \quad U: \mathfrak{H} \rightarrow \mathfrak{H}$$

$$U(R_1 \cdot R_2) = U(R_1) \cdot U(R_2)$$

$$U(R^{-1}) = U(R)^{-1} = U(R)^\dagger$$

$$U(\mathbb{1}) = \mathbb{1}$$

Our example,  $n=1$  particles:

$$\begin{aligned} \langle \vec{x} | U(R_2) U(R_1) | \psi \rangle &= \langle R_2^{-1} \vec{x} | U(R_2) | \psi \rangle = \langle R_1^{-1} R_2^{-1} \vec{x} | \psi \rangle \\ &= \langle (R_2 R_1)^{-1} \vec{x} | \psi \rangle = \langle \vec{x} | U(R_2 R_1) | \psi \rangle \end{aligned}$$

Inverse, identity: straight forward.

\*  $\{R\}$  and  $\{U(R)\}$  have the same multiplication rules as specified above

$$\begin{aligned} * \frac{\partial}{\partial \alpha} R |_{\alpha=0} &= \vec{n} = \vec{1} = n_k \vec{T}_k \\ \frac{\partial}{\partial \alpha} U(R) |_{\alpha=0} &= -\frac{i}{\hbar} \vec{n} \cdot \vec{L} \Rightarrow \{ \vec{T}_k \}, \{ \frac{-i}{\hbar} L_k \} \text{ have the algebra of commutation relations} \end{aligned}$$

\*  $R \leftrightarrow \alpha, \vec{n} \ (\vec{n}^2 = 1) \leftrightarrow 3$  real parameters

$U(R)$  is defined in infinite dimensional Hilbert space: do subspaces exist, which are invariant under all  $\{U(R)\}$ ? ( $\Leftrightarrow$  invariant under  $\vec{L}$ ) Irreducible representations (irreps)

### 2.4 Representations of the Rotation Group

Goal: find the irreps in the Hilbert space of the algebra

$$[J_i, J_j] = i\hbar \epsilon_{ijk} J_k$$

Procedure: find the simultaneous eigenvectors of  $J_3, J^2$ :

$$\begin{aligned} [J_i, J_k J_k] &= J_k [J_i, J_k] + [J_i, J_k] J_k = i\hbar (J_k \epsilon_{ike} J_e + \epsilon_{ike} J_e J_k) \\ &= i\hbar \epsilon_{ike} (J_k J_e + J_e J_k) = 0 \end{aligned}$$

antisymmetric symmetric when  $k \leftrightarrow l$

$$\begin{aligned} J^2 |j, m\rangle &= \hbar^2 \alpha_j |j, m\rangle ; \alpha_j \geq 0, \text{ real } (\hbar^2 \text{ because of dimension}) \\ J_3 |j, m\rangle &= \hbar m |j, m\rangle \end{aligned}$$

Values of  $m, \alpha_j$  still undetermined.

$\{|j, m\rangle\}$  is a CON. Since  $[J_i, J^2] = 0$  it follows

$$J_i |j, m\rangle = \sum_{m'} c_{i, m, m'}^j |j, m'\rangle \quad (j \leftrightarrow \alpha_j \text{ unchanged})$$

$$\langle j, m' | j, m \rangle = \delta_{mm'}$$

Instead of  $J_1, J_2 \rightarrow J_{\pm} = J_1 \pm iJ_2, J_3 = J_z$ .

$$\langle j, m' | J_3 | j, m \rangle = \hbar m \delta_{mm'}$$

$$\langle j, m' | J_{\pm} | j, m \rangle = \hbar \times ?$$

$$* \hbar^2 m^2 = \langle j, m | J_3^2 | j, m \rangle \leq \langle j, m | J^2 | j, m \rangle = \hbar^2 \alpha_j \Rightarrow m^2 \leq \alpha_j$$

$$* [J_3, J_{\pm}] = [J_3, J_1 \pm iJ_2] = i\hbar (J_2 \mp iJ_1) = \pm\hbar (J_2 \pm iJ_1) = \pm\hbar J_{\pm}$$

$$\Rightarrow J_3 J_{\pm} |j, m\rangle = ([J_3, J_{\pm}] + J_{\pm} J_3) |j, m\rangle = \hbar (m \pm 1) J_{\pm} |j, m\rangle$$

$\Rightarrow J_{\pm} |j, m\rangle$  is eigenvector of  $J_3$  with eigenvalue  $\hbar (m \pm 1)$

$$\Rightarrow J_3 (J_{\pm})^n |j, m\rangle = \hbar (m \pm n) (J_{\pm})^n |j, m\rangle$$

It is possible that  $(J_{\pm})^n |j, m\rangle = 0$  for some  $n$ .

\* Since  $m^2 \leq \alpha_j, \exists |j, m_{\max}\rangle$  with  $J_+ |j, m_{\max}\rangle = 0$

Convention:  $j = m_{\max}$  ( $j$  is not the eigenvalue  $\alpha_j$ )

$$J_+ |j, j\rangle = 0$$

\* Again, since  $m^2 \leq \alpha_j, \exists n$  such that  $(J_-)^{n+1} |j, j\rangle = 0, (J_-)^n |j, j\rangle \neq 0$

$$|j, j-k\rangle := c_k (J_-)^k |j, j\rangle, \quad k=0, 1, \dots, n$$

dimension of the representation =  $n+1$

$$\langle j, m | j, m' \rangle = \delta_{mm'}, \quad c_k > 0$$

fixes  $c_k$  uniquely, only phase of  $|j, j\rangle$  (global phase for all  $|j, m\rangle$ ) is still free.

\* from the commutation relations  $[J_i, J_j] = i\hbar \epsilon_{ijk} J_k$  it follows

$$J_+ J_- = J^2 - J_3^2 + \hbar J_3$$

$$J_- J_+ = J^2 - J_3^2 - \hbar J_3$$

\* computation of  $j$  ( $= m_{\max}$ ),  $n$ ,  $\alpha_j$  and matrix elements:

$$0 = J_- J_+ |j, j\rangle = \hbar^2 (\alpha_j - j^2 - j) |j, j\rangle \quad (1)$$

with  $m_{\min} = m_{\max} - n$ :

$$0 = J_+ J_- |j, m_{\min}\rangle = \hbar^2 (\alpha_j - m_{\min}^2 + m_{\min}) |j, m_{\min}\rangle \quad (2)$$

$$(1) \text{ and } (2) \Rightarrow j(j+1) = m_{\min}(m_{\min} - 1) = \alpha_j$$

$$m_{\min} = \begin{cases} j+1 \\ -j \end{cases}$$

$$n = m_{\max} - m_{\min} = 2j$$

$$\Rightarrow \begin{cases} m = j, j-1, \dots, -j+1, -j : 2j+1 \text{ values} \\ 2j = n \text{ integer} \\ j = 0, \frac{1}{2}, 1, \frac{3}{2}, \dots : j \text{ (and } m) \text{ must be an integer or a half integer!} \\ \alpha_j = j(j+1) \end{cases}$$

matrix elements:

$$|J_- |j, m\rangle|^2 = \langle j, m | J_+ J_- |j, m\rangle = \langle j, m | J^2 - J_3^2 + \hbar J_3 |j, m\rangle$$

$$= \hbar^2 [j(j+1) - m(m-1)]$$

$$\Rightarrow J_- |j, m\rangle = \hbar \sqrt{j(j+1) - m(m-1)} |j, m-1\rangle \quad \text{real and positive phase}$$

$$= \hbar \sqrt{(j+m)(j-m+1)} |j, m-1\rangle$$

(= 0 for  $m = -j$ )

analogously:  $J_+ |j, m\rangle = \hbar \sqrt{(j-m)(j+m+1)} |j, m+1\rangle$   
 (= 0 for  $m = j$ )

$$\langle j, m' | J_- |j, m\rangle = \sqrt{(j+m)(j-m+1)} \hbar \delta_{m', m-1} \quad \langle j, m' | J_3 |j, m\rangle = \hbar m \delta_{m', m}$$

$$\langle j, m' | J_+ |j, m\rangle = \sqrt{(j-m)(j+m+1)} \hbar \delta_{m', m+1} \quad \langle j, m' | J^2 |j, m\rangle = \hbar^2 j(j+1) \delta_{m', m}$$

### 2.4.1 Spin $j = \frac{1}{2}$

$j=0$  : trivial irrep, one-dimensional :  $|00\rangle$   
 $J^2|00\rangle = 0, J_3|00\rangle = 0$

$j = \frac{1}{2}$  :  $m = \frac{1}{2}, -\frac{1}{2}$  :  $|\frac{1}{2} \frac{1}{2}\rangle =: |+\rangle, |\frac{1}{2} -\frac{1}{2}\rangle =: |-\rangle$ , two-dimensional

$$J_3 \rightarrow \langle jm' | J_3 | jm \rangle = \frac{\hbar}{2} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

$$J_+ \rightarrow = \frac{\hbar}{2} \begin{pmatrix} 0 & 1 \\ 0 & 0 \end{pmatrix} = \begin{pmatrix} \langle + | J_+ | + \rangle & \langle + | J_+ | - \rangle \\ \langle - | J_+ | + \rangle & \langle - | J_+ | - \rangle \end{pmatrix}$$

$$J_- \rightarrow = \frac{\hbar}{2} \begin{pmatrix} 0 & 0 \\ 1 & 0 \end{pmatrix}$$

$$J_1 = \frac{1}{2} (J_+ + J_-) \rightarrow = \frac{\hbar}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$$

$$J_2 = \frac{1}{2i} (J_+ - J_-) \rightarrow = \frac{\hbar}{2} \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$$

$$J_k \rightarrow \langle \frac{1}{2} m' | J_k | \frac{1}{2} m \rangle = \frac{\hbar}{2} (\sigma_k)_{m'm}, \quad k=1,2,3$$

$\{\sigma_1, \sigma_2, \sigma_3\} = \left\{ \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \right\}$  Pauli matrices

\* basis of Hermitian traceless  $2 \times 2$  matrices

$$* \sigma_i \sigma_j = \delta_{ij} + i \epsilon_{ijk} \sigma_k \quad (\text{multiplication table})$$

$$\Rightarrow [\sigma_i, \sigma_j] = 2i \epsilon_{ijk} \sigma_k \quad (\text{commutator algebra})$$

$$\{\sigma_i, \sigma_j\} = 2\delta_{ij}$$

$\Rightarrow \left[ \frac{\hbar}{2} \sigma_i, \frac{\hbar}{2} \sigma_j \right] = i\hbar \epsilon_{ijk} \frac{\hbar}{2} \sigma_k \Rightarrow \left\{ \frac{\hbar}{2} \sigma_k \right\}$  obey algebra of angular momentum

\* two-dimensional irrep of angular momentum

it does not occur among the irreps of  $U(R)$  (orbital angular momentum), see book of J. Münster.



### 2.4.2 Spin 1

$$j = 1, \quad m = 1, 0, -1$$

$$J_- \rightarrow \langle 1 m' | J_- | 1 m \rangle = \hbar \begin{pmatrix} 0 & 0 & 0 \\ \sqrt{2} & 0 & 0 \\ 0 & \sqrt{2} & 0 \end{pmatrix} \left[ \begin{array}{c} m \\ 1 \\ 0 \\ -1 \end{array} \right] \begin{array}{c} \xrightarrow{m} \\ 1 \quad 0 \quad -1 \end{array}$$

$$J_+ \rightarrow = \hbar \begin{pmatrix} 0 & \sqrt{2} & 0 \\ 0 & 0 & \sqrt{2} \\ 0 & 0 & 0 \end{pmatrix}$$

$$J_z \rightarrow = \hbar \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & -1 \end{pmatrix}$$

$$J_1 = \frac{1}{2} (J_+ + J_-) \rightarrow = \hbar \frac{1}{\sqrt{2}} \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}$$

$$J_2 = \frac{1}{2i} (J_+ - J_-) \rightarrow = \hbar \frac{i}{\sqrt{2}} \begin{pmatrix} 0 & -1 & 0 \\ 1 & 0 & -1 \\ 0 & 1 & 0 \end{pmatrix}$$

Spin 1 representation matrices, fulfill algebra of angular momentum.

$i\hbar \vec{T}$ ,  $(T_i)_{jk} = \epsilon_{jik}$  fulfill algebra of angular momentum and are  $3 \times 3$  matrices, too. Yet a different representation? No, they are equivalent up to a unitary similarity transformation

# 3. The Electron Spin

## 3.1. Particles in the Electromagnetic Field

The electromagnetic field can be expressed in terms of potentials  $(\phi, \vec{A})$  as

$$\vec{E} = -\vec{\nabla}\phi - \frac{\partial \vec{A}}{\partial t} \quad \vec{\nabla} \times \vec{E} = -\frac{\partial \vec{B}}{\partial t} \quad (\text{Faraday's law})$$

$$\vec{B} = \vec{\nabla} \times \vec{A} \quad \Rightarrow \quad \vec{\nabla} \cdot \vec{B} = 0 \quad (\text{absence of free magnetic poles})$$

\*  $\Rightarrow$ : automatically fulfill Maxwell's equation

\*  $\vec{E}, \vec{B}$  are invariant under the gauge transformations:

$$\phi \mapsto \phi' = \phi - \frac{\partial \alpha}{\partial t} \quad \alpha = \alpha(\vec{x}, t) \text{ real function}$$

$$\vec{A} \mapsto \vec{A}' = \vec{A} + \vec{\nabla}\alpha$$

$$H = \frac{1}{2m} \left( \vec{p} - q \vec{A} \right)^2 + q\phi \quad : \text{Hamiltonian of a non-relativistic particle in the electromagnetic field}$$

operator  $\vec{A}(\vec{x}, t)$

$\psi(\vec{x}, t)$  is a solution of Schrödinger's equation with  $\phi, \vec{A} \iff$

$\psi'(\vec{x}, t) = \Omega(\vec{x}, t) \psi(\vec{x}, t)$  is a solution of Schrödinger's equation with gauge transformed  $\phi', \vec{A}'$  and

$$\Omega(\vec{x}, t) = e^{i \frac{q}{\hbar} \alpha(\vec{x}, t)} \quad : \text{gauge transformation}$$

$$\rho = |\psi|^2, \quad \vec{j} = \frac{\hbar}{2mi} \left\{ \psi^* \left( \vec{\nabla} - \frac{iq\vec{A}}{\hbar} \right) \psi - \psi \left( \vec{\nabla} + \frac{iq\vec{A}}{\hbar} \right) \psi^* \right\}$$

$q \cdot \rho = \text{charge density}$   
 $q \cdot \vec{j} = \text{electrical current density}$  } are invariant under gauge transformations

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0$$

electron in a homogeneous magnetic field  $\vec{B}(\vec{x}) = \vec{B} = \text{const}$ ,  $q = -e$  (24)

$$\vec{A} = \frac{1}{2} \vec{B} \times \vec{x} \Rightarrow \vec{\nabla} \times \vec{A} = \vec{B}$$

$$(\vec{\hat{p}} + e\vec{A})^2 = \vec{\hat{p}}^2 + e^2 \vec{A}^2 + e(\vec{\hat{p}} \cdot \vec{A} + \vec{A} \cdot \vec{\hat{p}})$$

last term: it holds  $[\hat{p}_k, f(\vec{x})] = \frac{\hbar}{i} \frac{\partial f}{\partial x_k}(\vec{x})$

$$\begin{aligned} \text{proof: } \langle \vec{y} | [\hat{p}_k, f(\vec{x})] | \psi \rangle &= \langle \vec{y} | \hat{p}_k f(\vec{x}) | \psi \rangle - \langle \vec{y} | f(\vec{x}) \hat{p}_k | \psi \rangle \\ &= \frac{\hbar}{i} \frac{\partial}{\partial y_k} \langle \vec{y} | f(\vec{x}) | \psi \rangle - f(\vec{y}) \frac{\hbar}{i} \frac{\partial}{\partial y_k} \langle \vec{y} | \psi \rangle \\ &= \frac{\hbar}{i} \frac{\partial f}{\partial y_k}(\vec{y}) \langle \vec{y} | \psi \rangle = \langle \vec{y} | \frac{\hbar}{i} \frac{\partial f}{\partial x_k}(\vec{x}) | \psi \rangle \end{aligned}$$

$$\begin{aligned} \Rightarrow \vec{\hat{p}} \cdot \vec{A} &= \vec{A} \cdot \vec{\hat{p}} + [\hat{p}_k, A_k] = \vec{A} \cdot \vec{\hat{p}} + \frac{\hbar}{i} \frac{\partial A_k}{\partial x_k} \\ &= \vec{A} \cdot \vec{\hat{p}} + \frac{\hbar}{i} \vec{\nabla} \cdot \vec{A} \end{aligned}$$

$$\text{due to } \vec{A} = \frac{1}{2} \vec{B} \times \vec{x} : \vec{\nabla} \cdot \vec{A} = \frac{1}{2} \epsilon_{ijk} \frac{\partial}{\partial x_i} B_j x_k = 0 \quad \text{Coulomb-gauge}$$

$\underbrace{\frac{\partial}{\partial x_i} B_j x_k}_{= \delta_{ik}}$

$$\Rightarrow \vec{\hat{p}} \cdot \vec{A} + \vec{A} \cdot \vec{\hat{p}} = 2\vec{A} \cdot \vec{\hat{p}} = (\vec{B} \times \vec{x}) \cdot \vec{\hat{p}} = \vec{B} \cdot (\vec{x} \times \vec{\hat{p}}) = \vec{B} \cdot \vec{L}$$

$$\Rightarrow H = \frac{\vec{\hat{p}}^2}{2m} + \frac{e}{2m} \vec{B} \cdot \vec{L} + \frac{e^2}{8m} (\vec{B} \times \vec{x})^2 - e\phi$$

### 3.2 The Zeeman effect

$\frac{e}{2m} \vec{B} \cdot \vec{L}$  paramagnetic term

$$= -\vec{\mu} \cdot \vec{B} \quad \text{with } \vec{\mu} = -\frac{e}{2m} \vec{L} = \text{magnetic moment } (q = -e)$$

(permanent, results from the motion of the particle, exists also if  $\vec{B} = 0$ )

if  $\vec{\mu}$  is parallel to  $\vec{B} \Rightarrow$  energy decreases

$$\mu_B = \frac{e \hbar}{2m} = 0.579 \times 10^{-4} \frac{eV}{T} = \text{Bohr's magneton}$$

$$\frac{e^2}{8m} (\vec{B} \times \vec{r})^2 \text{ diamagnetic term}$$

(magnetic field changes the particle's motion and induces a extra magnetic moment  $\propto B$ )

$$\text{Orders of magnitude: } \frac{e^2}{m} B^2 a_B^2 \sim (\mu_B B)^2 \frac{m a_B^2}{\hbar^2} \sim \mu_B B \cdot \frac{\mu_B}{E_R} \cdot B$$

$$\frac{\mu_B}{E_R} \sim \frac{10^{-5}}{T}$$

in the lab usually  $B \sim 1T \Rightarrow$  diamagnetic  $\sim 10^{-5}$  x paramagnetic, diamagnetic term can be neglected compared to  $\mu_B B$

$$\text{With } \vec{B} = B \vec{e}_3: H \cong H_0 + \frac{\mu_B}{\hbar} B L_3$$

$$H_0: \text{hydrogen-atom-like, } \phi = \frac{Ze}{4\pi\epsilon_0} \frac{1}{|\vec{r}|}$$

$$H \psi_{nlm} = E_{nlm} \psi_{nlm}$$

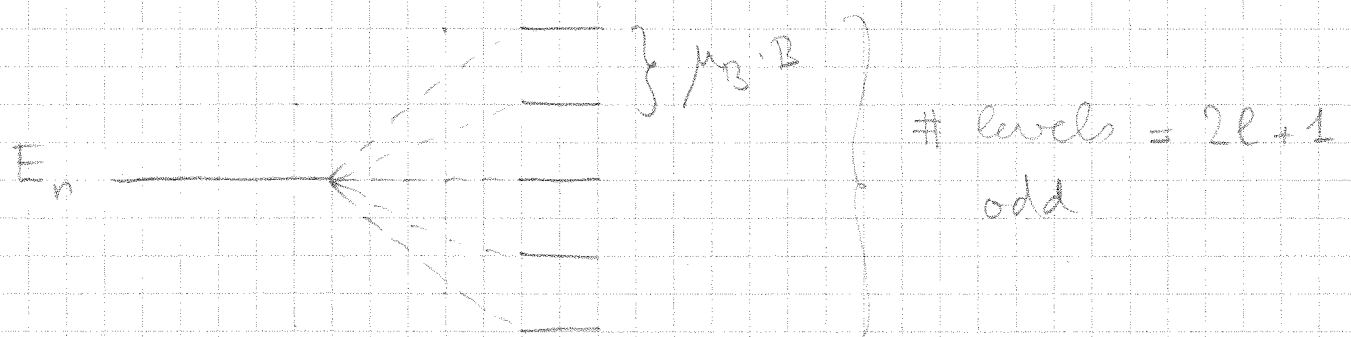
$$E_{nlm} = -\frac{E_R Z^2}{n^2} + \mu_B B m$$

m-quantum number:  $L_3 \psi_{nlm} = \hbar m \psi_{nlm}$

easy because  $\{H_0, L^2, L_3\}$  simultaneously diagonal

\* m-degeneracy of  $E_n$  is split by  $B \neq 0$

\* m = magnetic quantum number



Zeeman effect

Effect was observed in the 1890's by the spectroscopist Pieter Zeeman (splitting of the D spectral lines of sodium)

\* m-degeneracy  $\leftrightarrow$  rotational invariance ( $L_z$  change m) broken by external magnetic field  $\vec{B} \neq 0$

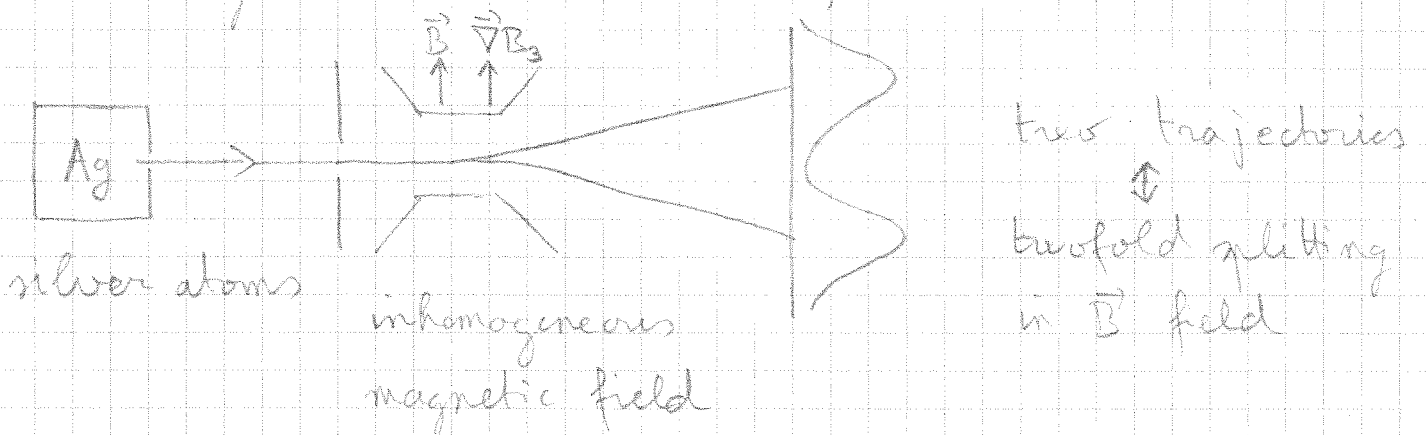
Experimental observations:

- ground state of hydrogen atom has  $l=0, m=0$   
 $\Rightarrow$  no Zeeman splitting

but: energy level is doubled

- Stern - Gerlach experiment:

Walter Gerlach and Otto Stern, in 1922



Splitting is not observed with  $Ag^+$  ions

- in general: even splitting for atoms with an odd number of electrons

- splittings  $g \mu_B B$  with  $g = 0(1) =$  Landé factor or gyromagnetic factor depend on the multiplet

seems like half-integer  $l, m$  although they do not occur for orbital angular momentum  $\vec{L} = \vec{r} \times \vec{p}$

Interpretation by George Uhlenbeck and Samuel Goudsmit, 1925: splitting is due to spin, intrinsic angular momentum of the electron with  $j = \frac{1}{2}$ , independently of

the orbital motion

permanent magnetic moment  $\vec{\mu} = \frac{-e}{2m} g \vec{S}$ ,  $m_s = \pm \frac{\hbar}{2}$

### 3.3 The Electron with Spin $\frac{1}{2}$

$$\mathcal{H} = \mathcal{H}_{\vec{x}} \otimes \mathcal{H}_S$$

$\mathcal{H}_{\vec{x}}$  like so far, basis  $|\vec{x}\rangle$

$\mathcal{H}_S \cong \mathbb{C}^2$  Hilbert space where spin operator  $\vec{S}$  acts

Basis  $\{|\frac{1}{2} m\rangle, m = +\frac{1}{2}, -\frac{1}{2}\}$

$$[S_i, S_j] = i\hbar \epsilon_{ijk} S_k$$

$$S^2 |\frac{1}{2} m\rangle = \hbar^2 \frac{1}{2} (\frac{1}{2} + 1) |\frac{1}{2} m\rangle = \hbar^2 \frac{3}{4} |\frac{1}{2} m\rangle$$

$$S_3 |\frac{1}{2} m\rangle = \hbar m |\frac{1}{2} m\rangle$$

simplified notation:  $|+\rangle = |\frac{1}{2} \frac{1}{2}\rangle$ ,  $|-\rangle = |\frac{1}{2} -\frac{1}{2}\rangle$

$$\langle \epsilon' | \vec{S} | \epsilon \rangle = \frac{\hbar}{2} (\vec{\sigma})_{\epsilon'\epsilon}, \quad \epsilon', \epsilon = +, -$$

$\vec{\sigma}$ : Pauli matrices

basis in  $\mathcal{H}$ :  $\{|\vec{x} \epsilon\rangle = |\vec{x}\rangle \otimes |\epsilon\rangle\}$

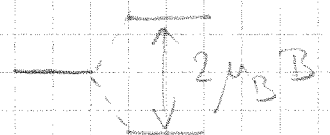
$$|+\rangle = \int d^3x \{ |\vec{x} +\rangle \psi_+(\vec{x}) + |\vec{x} -\rangle \psi_-(\vec{x}) \}$$

$\begin{pmatrix} \psi_+(\vec{x}) \\ \psi_-(\vec{x}) \end{pmatrix}$  two-component wave function, Pauli-spinor

$$H = H_{\text{Schrödinger}} \otimes \mathbb{1} + \frac{\mu_B}{\hbar} g_e \vec{B} \cdot (\mathbb{1} \otimes \vec{S})$$

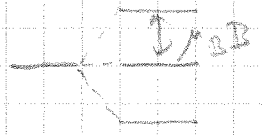
- yields the correct splittings

-  $g_e \approx 2$ :  $\Rightarrow$  eigenvalues of  $g_e S_3$ :  $\pm 2 \cdot \frac{\hbar}{2}$

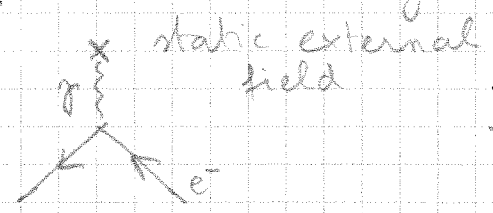


$g_e \approx 2 \Rightarrow$  spacing of energy levels  $\approx$  like for

$l=1: \mu_B B m, m = +1, 0, -1$



later: relativistic Dirac - theory of the electron yields  $g_e = 2$   
Corrections due to processes like the emission and absorption of photons shift  $g_e = 2.002322...$  (quantum electrodynamics)



$+ \dots g = 2 + \frac{\alpha}{\pi} + \dots$

Schwinger, 1948

fine structure constant  $\frac{e^2}{4\pi\epsilon_0 \hbar c} \approx \frac{1}{137}$

remark: there are many other particles with spin  $\frac{1}{2}$ , for example the proton:

$\vec{\mu}_s = g_p \frac{\mu_p}{\hbar} \vec{S}$

$\mu_p = \frac{e\hbar}{2m_p} \approx \frac{\mu_B}{1836}$

$g_p \approx 5.58$  (complicated internal structure)

$g_n \approx -3.82$  (neutron)

### 3.4 The Pauli Equation

$i\hbar \frac{d}{dt} |\psi(t)\rangle = H |\psi(t)\rangle$

position-spin basis:  $\psi(\vec{x}, t) = \begin{pmatrix} \psi_+(\vec{x}, t) \\ \psi_-(\vec{x}, t) \end{pmatrix}$

$H = \frac{1}{2m} \left( \hbar \vec{\nabla} + e \vec{A}(\vec{x}, t) \right)^2 - e \phi(\vec{x}, t) + \mu_B \vec{B} \cdot \vec{\sigma} \quad (g_e = 2)$

It can be shown:

$$H = \frac{1}{2m} \left[ \vec{\sigma} \cdot \left( \frac{\hbar}{i} \vec{\nabla} + e \vec{A} \right) \right]^2 - e\phi$$

For bound states:  $1 = \int d^3x \psi^\dagger \psi = \int d^3x (|\psi_+|^2 + |\psi_-|^2)$

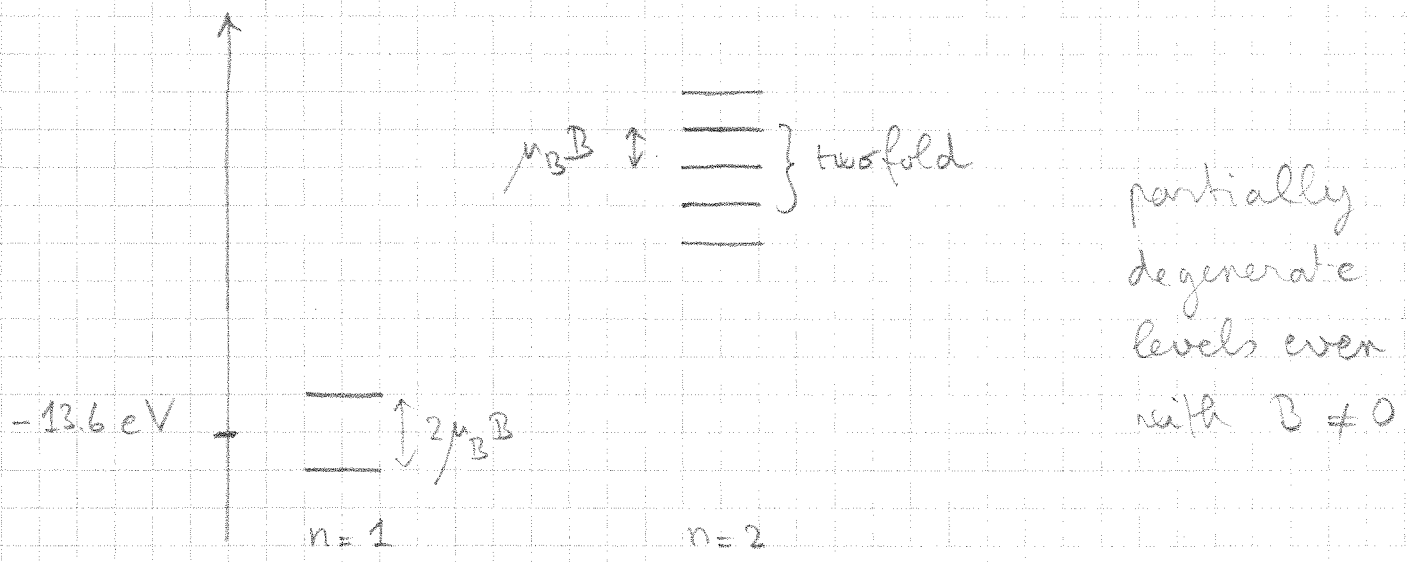
$|\psi_\pm|^2$  = probability density for electron with  $S_z = \pm \frac{\hbar}{2}$

Eigenfunctions ( $\vec{B} = B \vec{e}_3$ , no diamagnetic term, hydrogen-atom-like as in 3.2):

$$R_{ne} Y_{lm} \begin{pmatrix} 1 \\ 0 \end{pmatrix}, R_{ne} Y_{lm} \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

$$E = -\frac{E_R Z^2}{n^2} + \underbrace{\mu_B B}_{\text{Zeeman}} (m \pm 1)$$

$2l+1$  states with  $S_z = +\frac{\hbar}{2}$   
and  $2l+1$  states with  $S_z = -\frac{\hbar}{2}$



degeneracy for  $B=0$ :  $2n^2$

Complete set of commuting observables in the Pauli theory:

$$H, \vec{L}^2, L_3, \vec{S}^2 = \frac{\hbar^2}{4} \vec{\sigma}^2 = \frac{3\hbar^2}{4} \cdot \mathbb{1}, S_3$$

Relativistic ( $\frac{v}{c}$  not negligible):  $[H, L_3] \neq 0, [H, S_3] \neq 0$   
because H contains a term  $\vec{S} \cdot \vec{L} \cdot O\left(\frac{v^2}{c^2}\right)$

$$\rightarrow H, \vec{J}^2, J_3, \vec{L}^2, \vec{S}^2 = \frac{3\hbar^2}{4} \cdot \mathbb{1}$$



$$\vec{J} = \vec{L} + \vec{S} = \vec{L} \otimes \mathbb{1} + \mathbb{1} \otimes \vec{S}, \quad [\vec{J}, \vec{S} \cdot \vec{L}] = 0 \quad (40)$$

### 3.5 Rotations of Spinors

We show: the total angular momentum  $\vec{J} = \vec{L} + \vec{S}$  is the generator of rotations in the Pauli theory:

$$[L_i, L_j] = i\hbar \epsilon_{ijk} L_k \quad [S_i, S_j] = i\hbar \epsilon_{ijk} S_k$$

$$[S_i, L_j] = 0 \quad (\text{different factors of tensor product})$$

$$\Rightarrow [J_i, J_j] = i\hbar \epsilon_{ijk} J_k$$

For  $R = \mathbb{1} + \delta \vec{n} \cdot \vec{T}$ :  $|\psi'\rangle \approx \left( \mathbb{1} - \delta \frac{i}{\hbar} \vec{n} \cdot \vec{J} \right) |\psi\rangle$  infinitesimal

$$\text{finite: } |\psi'\rangle = e^{-\frac{\alpha}{\hbar} \vec{n} \cdot \vec{J}} |\psi\rangle = e^{-\frac{\alpha}{\hbar} \vec{n} \cdot \vec{L}} e^{-\frac{\alpha}{\hbar} \vec{n} \cdot \vec{S}} |\psi\rangle$$

$$=: U(R) |\psi\rangle$$

$$\langle \vec{x}' | \psi' \rangle = \langle R^{-1} \vec{x}' | \epsilon | e^{-\frac{\alpha}{\hbar} \vec{n} \cdot \vec{S}} | \psi \rangle = \langle R^{-1} \vec{x}' | \otimes \langle \epsilon | e^{-\frac{\alpha}{\hbar} \vec{n} \cdot \vec{S}} | \psi \rangle$$

$$= \int d^3 y \sum_{\epsilon'} \langle R^{-1} \vec{x}' | \otimes \langle \epsilon | e^{-\frac{\alpha}{\hbar} \vec{n} \cdot \vec{S}} | \epsilon' \rangle \otimes | y \rangle \langle y \epsilon' | \psi \rangle$$

$$= \sum_{\epsilon'} \langle \epsilon | e^{-\frac{\alpha}{\hbar} \vec{n} \cdot \vec{S}} | \epsilon' \rangle \langle R^{-1} \vec{x}' | \epsilon' | \psi \rangle$$

$$= \sum_{\epsilon'} \left( e^{-i \frac{\alpha}{2} \vec{n} \cdot \vec{\sigma}} \right)_{\epsilon \epsilon'} \langle R^{-1} \vec{x}' | \epsilon' | \psi \rangle$$

$$\Rightarrow \begin{pmatrix} \psi'_+ (\vec{x}') \\ \psi'_- (\vec{x}') \end{pmatrix} = e^{-i \frac{\alpha}{2} \vec{n} \cdot \vec{\sigma}} \begin{pmatrix} \psi_+ (R^{-1} \vec{x}') \\ \psi_- (R^{-1} \vec{x}') \end{pmatrix}$$

\* transformation of the argument and of the spinor components

$$* (\vec{n} \cdot \vec{\sigma})^2 = n_i n_j \sigma_i \sigma_j = n_i n_i = 1 \quad \text{since } \sigma_i \sigma_j = \delta_{ij} + i \epsilon_{ijk} \sigma_k$$

$$\Rightarrow e^{-i \frac{\alpha}{2} \vec{n} \cdot \vec{\sigma}} = \cos \left( \frac{\alpha}{2} \right) \mathbb{1} - i \sin \left( \frac{\alpha}{2} \right) \vec{n} \cdot \vec{\sigma}$$

Special case  $\alpha = 2\pi$ :  $e^{-i\pi \vec{n} \cdot \vec{\sigma}} = \cos(\pi) \mathbb{1} = -\mathbb{1}$

$\Rightarrow U(R) |\psi\rangle = -|\psi\rangle$  under rotation by  $2\pi$

\*  $\vec{S}$  is a vector operator  $\Leftrightarrow U^\dagger(R) S_i U(R) = R_{ik} S_k$

Proof: infinitesimal:

$$U^\dagger(R) S_i U(R) \approx \left( \mathbb{1} + \delta \frac{i}{\hbar} \vec{n} \cdot \vec{J} \right) S_i \left( \mathbb{1} - \delta \frac{i}{\hbar} \vec{n} \cdot \vec{J} \right)$$

$$= S_i + \delta \frac{i}{\hbar} n_j [J_j, S_i]$$

$\vec{J}$  commutes with  $\vec{S} \Rightarrow$   $= S_i + \delta \frac{i}{\hbar} n_j [S_j, S_i]$

$$= S_i - \delta n_j \epsilon_{jik} S_k = S_i + \delta n_j (\epsilon_{ijl})_{lk} S_k$$

$$= (\delta_{ik} + \delta (\vec{n} \cdot \vec{T})_{ik}) S_k$$

$$\stackrel{||}{=} R_{ik} S_k \quad \square$$

$\Rightarrow$  terms like  $\langle \psi | \vec{B} \cdot \vec{S} | \psi \rangle$  are invariant (scalar product of two vectors), when

\*  $|\psi\rangle \mapsto U(R) |\psi\rangle$

\* external field  $\vec{B} \mapsto R\vec{B}$

# 4. Addition of Angular Momenta

## 4.1 Addition of Two Angular Momenta

### 4.1.1 Motivation

Angular momenta  $\vec{j} \equiv (\vec{j} \otimes \mathbb{1})$  and  $\vec{j}' \equiv (\mathbb{1} \otimes \vec{j}')$  in state space  $\mathcal{H} \otimes \mathcal{H}'$  with

$$[j_i, j_k] = i\hbar \epsilon_{ikl} j_l, \text{ similarly } \vec{j}', [j_i, j'_k] = 0$$

Examples:

- $\vec{j} = \vec{S}_1, \vec{j}' = \vec{S}_2$  spins of two electrons in the ground state of the Helium atom, no orbital angular momentum
- $\vec{j} = \vec{L}, \vec{j}' = \vec{S}$  for the Pauli electron

Presence of interaction between angular momenta usually has the effect that they are not separately conserved

Example: term  $\vec{j} \cdot \vec{j}'$  in  $H \Rightarrow [\vec{j}, H] \neq 0, [\vec{j}', H] \neq 0$

Total angular momentum:  $\vec{J} = \vec{j} + \vec{j}'$  in  $\mathcal{H} \otimes \mathcal{H}'$

$$[\vec{J}, \vec{j} \cdot \vec{j}'] = 0$$

$$[j_i, j_k j_k] = [j_i, j_k] j_k + j_k [j_i, j_k] = i\hbar \epsilon_{ikl} (j_l j_k + j_k j_l) = 0$$

$\vec{J}$  commutes with  $H \rightarrow$  diagonalize  $J^2$  and  $J_z$ !

Problem: how do we express the eigenstates of  $\vec{J}$  in terms of the eigenstates of  $\vec{j}$  and  $\vec{j}'$ ?

Tensor-product basis  $|rjm\rangle \otimes |r'j'm'\rangle = |rr'jj'mm'\rangle$

diagonalizes  $J^2, J_z, j_1^2, j_2^2, \dots$

Short hand notation:  $|r r' j j' m m'\rangle \rightarrow |m m'\rangle$   
 since we stay in the eigenspace with values  $\hbar^2 j(j+1)$  for  $\vec{j}^2$   
 and  $\hbar^2 j'(j'+1)$  for  $\vec{j}'^2$  and  $r, r'$  for ... ( $[\vec{j}, \vec{j}'^2] = 0 = [\vec{j}', \vec{j}^2]$ )

$\Rightarrow \{|m m'\rangle\}$  dimension  $(2j+1) \cdot (2j'+1)$

$$J_3 |m m'\rangle = (j_3 + j'_3) |m m'\rangle = \hbar (m + m') |m m'\rangle$$

$$\begin{aligned} \vec{j}^2 |m m'\rangle &= (\vec{j}^2 + \vec{j}'^2 + 2\vec{j} \cdot \vec{j}') |m m'\rangle \\ &= [\hbar^2 j(j+1) + \hbar^2 j'(j'+1) + (j_+ j'_- + j_- j'_+ + 2\hbar^2 m m')] |m m'\rangle \\ &= a |m m'\rangle + b |m+1 m'-1\rangle + c |m-1 m'+1\rangle \\ &\quad \text{not diagonal} \end{aligned}$$

$$2\vec{j} \cdot \vec{j}' = j_+ j'_- + j_- j'_+ + 2j_3 j'_3, \quad j_{\pm} = j_1 \pm i j_2, \quad j'_{\pm} = j'_1 \pm i j'_2$$

Construct  $|j M\rangle = \sum_{m, m'} |m m'\rangle \underbrace{\langle m m' | j M \rangle}_{\text{Clebsch-Gordan coefficients}}$

$$\vec{j}^2 |j M\rangle = \hbar^2 j(j+1) |j M\rangle \quad \text{Clebsch-Gordan coefficients}$$

$$J_3 |j M\rangle = \hbar M |j M\rangle \quad \text{which } j, M \text{ values?}$$

4.1.2  $j = \frac{1}{2} + j' = \frac{1}{2}$

In the tensor basis  $|1\rangle := |++\rangle, |2\rangle := |+-\rangle, |3\rangle := |-+\rangle, |4\rangle := |--\rangle$

$$\langle i | \frac{J_3}{\hbar} | k \rangle = \begin{pmatrix} 1 & & & \\ & 0 & & \\ & & 0 & \\ & & & -1 \end{pmatrix}$$

$$j_+ | - m' \rangle = \hbar | + m' \rangle$$

$$j_- | + m' \rangle = \hbar | - m' \rangle \quad \text{etc.}$$

(44)

$$\langle l | \frac{1}{\hbar^2} \vec{J}^2 | k \rangle = 2 \times \frac{3}{4} \delta_{ik} + \begin{pmatrix} \frac{1}{2} & & & \\ & \begin{pmatrix} -\frac{1}{2} & 1 \\ 1 & -\frac{1}{2} \end{pmatrix} & & \\ & & \frac{1}{2} & \\ & & & 0 \end{pmatrix} \begin{matrix} \langle + - | 2j_3 j_3' | + - \rangle \\ \langle + - | j_+ j_- | - + \rangle \end{matrix}$$

$$= \begin{pmatrix} 2 & 0 & 0 & 0 \\ 0 & 1 & 1 & 0 \\ 0 & 1 & 1 & 0 \\ 0 & 0 & 0 & 2 \end{pmatrix}$$

$\Rightarrow |+- \rangle$  and  $|-- \rangle$  have  $j=1 \Rightarrow j(j+1) = 2$  therefore

$$|j=1, M=1 \rangle = |++ \rangle = |1 \rangle$$

$$|j=1, M=-1 \rangle = |-- \rangle = |4 \rangle$$

$$\begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \begin{pmatrix} 1 \\ 1 \end{pmatrix} = 2 \begin{pmatrix} 1 \\ 1 \end{pmatrix}, \quad \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} \begin{pmatrix} 1 \\ -1 \end{pmatrix} = 0$$

$$\Rightarrow \frac{1}{\sqrt{2}} (|+- \rangle - |-+ \rangle) = \frac{1}{\sqrt{2}} (|2 \rangle - |3 \rangle) = |j=0, M=0 \rangle$$

antisymmetric  $\Leftrightarrow$  total spin is zero

$$\frac{1}{\sqrt{2}} (|+- \rangle + |-+ \rangle) = \frac{1}{\sqrt{2}} (|2 \rangle + |3 \rangle) = |j=1, M=0 \rangle$$

symmetric

Standard procedure to construct  $\vec{J}^2$  multiplets:

$$J_- |++ \rangle = (j_- + j_-') |++ \rangle = \hbar |-+ \rangle + \hbar |+- \rangle \doteq c \cdot |j=1, M=0 \rangle$$

with  $c > 0$ .  $|j=0, M=0 \rangle$  is a state orthogonal to  $|j=1, M=0 \rangle$ .

$\Rightarrow$  several  $J_-$ -values, 0 and 1, when adding  $\frac{1}{2} + \frac{1}{2}$ .

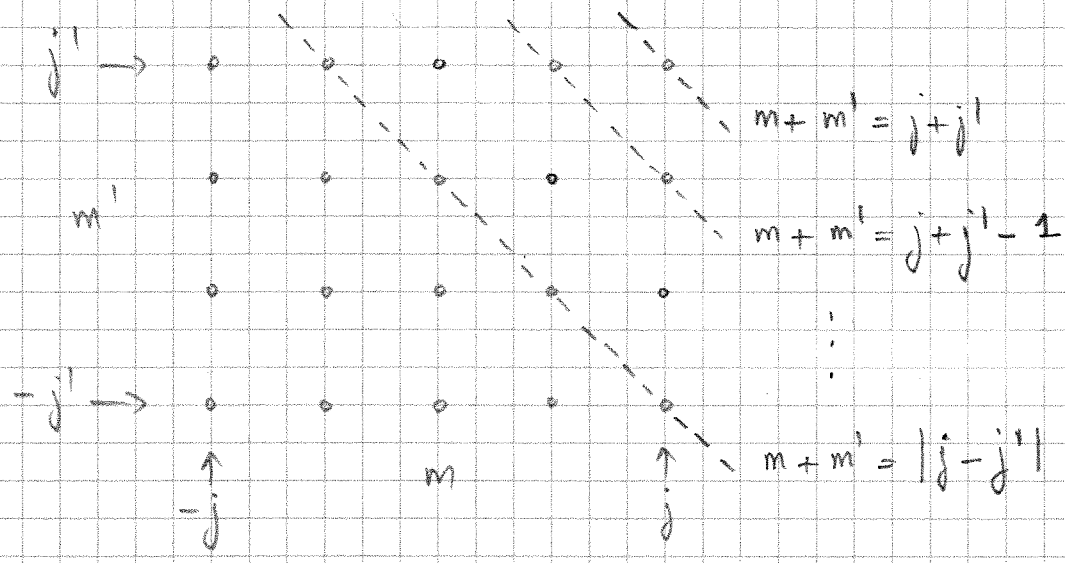
### 4.1.3 General Case

Since  $J_3 |mm' \rangle = \hbar (m+m') |mm' \rangle$  it holds:

$$M = m + m' \Rightarrow -j - j' \leq M \leq j + j'$$

$$\Rightarrow j \leq j + j'$$

$$\langle j M | m m' \rangle \propto \delta_{M, m+m'} \quad (\text{calculate } \langle j M | Y_3 | m m' \rangle)$$



\* maximal M :  $|m=j, m'=j'\rangle = |j=j+j', M=j\rangle$

"highest" state of the multiplet  $|j M\rangle, M=j, \dots, -j$ , with  $j=j+j'$

apply  $(Y_-)^k$  on  $|j j\rangle \rightarrow$  all  $|j M\rangle$  ( $j=j+j'$  fixed)

$$Y_{\pm} |j M\rangle = \hbar \sqrt{(j \mp M)(j \pm M + 1)} |j M \pm 1\rangle, \quad Y_{\pm} = j_1 \pm i j_2$$

4.1.4  $\left\{ \begin{aligned} Y_- |j j\rangle &= \hbar \sqrt{2j} |j j-1\rangle \\ &= (j_- + j'_-) |m=j, m'=j'\rangle = \hbar \sqrt{2j} |j-1, j'\rangle + \hbar \sqrt{2j'} |j, j'-1\rangle \end{aligned} \right.$

\* choose  $|j=j+j'-1, M=j\rangle$  as a linear combination of the states on the "line" with  $m+m' = j+j'-1$  such that

$$Y_+ |j=j+j'-1, M=j\rangle = 0$$

$\Rightarrow$  highest state  $|j M=j\rangle$  of the multiplet with  $j=j+j'-1$ , other states in the multiplet obtained by  $Y_-$

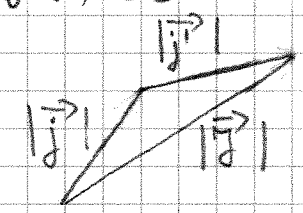
\* multiplets  $\gamma = j+j', \gamma = j+j'-1, \gamma = j+j'-2, \dots$   
 as long as there is one more state on the lines with  $M = m+m'$  fixed when  $J_-$  is applied ( $M \rightarrow M-1$ )

\* algorithm terminates when a corner is reached  
 $\Leftrightarrow M = |j-j'| =$  smallest value of  $\gamma$

\*  $|j, M\rangle$  with  $\gamma = j+j', j+j'-1, \dots, |j-j'|$  and  $M = \gamma, \gamma-1, \dots, -\gamma$  (for each  $\gamma$ )

\* complete set of states (given  $j, j'$ ): for example  $j'' \leq j$

$$\sum_{\gamma=j-j'}^{j+j'} (2\gamma+1) = (2j'+1) \frac{1}{2} [2(j+j')+1 + 2(j-j')+1] = (2j'+1)(2j+1)$$



\*  $|j-j'| \leq \gamma \leq |j+j'|$

$$\Rightarrow \sqrt{\gamma(\gamma+1)} < \sqrt{j(j+1)} + \sqrt{j'(j'+1)} \quad \text{triangle inequality}$$

$$j=j' = \frac{1}{2} \Rightarrow \gamma = 0, 1, \quad \sqrt{2} < \sqrt{\frac{3}{4}} + \sqrt{\frac{3}{4}} = \sqrt{3}$$

### 1.4 Computation of Clebsch-Gordan coefficients

$$|j, M\rangle = \sum_{m, m'} |m, m'\rangle \langle m, m' | j, M\rangle$$

$\langle m, m' | j, M\rangle$ : unitary matrix, size:  $[(2j+1)(2j'+1)]^2$   
 indices,  $(2j+1) \cdot (2j'+1)$  values

$$\sum_{m, m'} \langle m, m' | j, M\rangle^* \langle m, m' | j', M'\rangle = \delta_{j, j'} \delta_{M, M'}$$

$$\sum_{j, M} \langle m, m' | j, M\rangle^* \langle n, n' | j, M\rangle = \delta_{m, n} \delta_{m', n'}$$

Remarks: it will be shown that  $\langle m, m' | j, M\rangle$  can be chosen to be real

Phase convention:

- inside each "ladder" the phase is fixed:

$$Y_- |Y M\rangle = c |Y M-1\rangle \quad \text{with } c \text{ real and positive}$$

all states  $|1\rangle |^2 = 1 \Rightarrow c = +\hbar \sqrt{(Y+M)(Y-M+1)}$

- the overall phase of the multiplet is not fixed

convention of spherical harmonics:

$$Y_{l0}(0,0) = \text{real positive}$$

convention of Clebsch-Gordan coefficients:

$$\langle m m' | Y Y \rangle = \text{real positive for all } |j-j'| \leq Y \leq j+j'$$

at  $\underline{m=j}, m'=Y-j$

→ "entry point on the  $Y$ -ladder"

⇒ all  $\langle m m' | Y M \rangle$  real

This phase convention is asymmetric under  $\vec{j} \leftrightarrow \vec{j}'$

Finally:  $\boxed{Y = j + j'}$

$$|Y Y\rangle = |m=j, m'=j'\rangle \Rightarrow \langle j j' | Y Y \rangle = +1$$

Application of  $Y_- = j_- + j'_-$  on both sides:

$$j_- |Y Y\rangle = \hbar \sqrt{(Y+Y)(Y-Y+1)} |Y Y-1\rangle = \hbar \sqrt{2Y} |Y Y-1\rangle$$

$$= (j_- + j'_-) |m=j, m'=j'\rangle = \hbar \sqrt{2j} |j-1, j'\rangle + \hbar \sqrt{2j'} |j, j'-1\rangle$$

$$\Rightarrow |Y Y-1\rangle = \sqrt{\frac{j}{Y}} |j-1, j'\rangle + \sqrt{\frac{j'}{Y}} |j, j'-1\rangle$$

$$\langle j-1, j' | Y = j+j', M = Y-1 \rangle = \sqrt{\frac{j}{j+j'}}$$

$$\langle j, j'-1 | \quad \quad \quad \rangle = \sqrt{\frac{j'}{j+j'}}$$

etc. until  $\langle -j, -j' | Y = j+j', M = -Y \rangle = +1$



Afterwards :  $J = j + j' - 1$

$|j j\rangle = \alpha |j j'-1\rangle + \beta |j-1 j'\rangle$  ,  $\alpha > 0$  real (our phase convention)  
 $\alpha^2 + \beta^2 = 1$

$0 \stackrel{!}{=} J_+ |j j\rangle = \beta j_+ |j-1 j'\rangle + \alpha j'_+ |j j'-1\rangle$   
 $= (\hbar\beta\sqrt{2j} + \hbar\alpha\sqrt{2j'}) |j j'\rangle$

$\Rightarrow \alpha = \sqrt{\frac{j}{j+j'}}$  ,  $\beta = -\sqrt{\frac{j'}{j+j'}}$

$\langle j-1 j' | J = j + j' - 1 M = j \rangle = -\sqrt{\frac{j'}{j+j'}}$

$\langle j j'-1 | \quad \quad \quad \rangle = \sqrt{\frac{j}{j+j'}}$

other  $|j M\rangle$  from application of  $J_- = j_- + j'_-$

- tables (cf. Weinberg's book)
- tricks, recursion formulae, symmetries between Clebsch-Gordan coefficients
- computer

4.1.5 Example :  $1 + \frac{1}{2} \rightarrow$  exercises

for example Pauli electron with  $l=1, s=\frac{1}{2}, \vec{J} = \vec{L} + \vec{S}$  :

$|ms\rangle$  ,  $m = 1, 0, -1$  ,  $s = \frac{1}{2}, -\frac{1}{2}$

Dimension = dim = 6

Expect :  $J = \frac{3}{2}$  (dim=4) and  $J = \frac{1}{2}$  (dim=2)

$J = \frac{3}{2}$  :

$|J = \frac{3}{2} M = \frac{3}{2}\rangle = |m=1 s=\frac{1}{2}\rangle$

on the left :  $J_- | \frac{3}{2} \frac{3}{2} \rangle = \hbar\sqrt{3} | \frac{3}{2} \frac{1}{2} \rangle$

on the right :  $(L_- + S_-) | 1 \frac{1}{2} \rangle = \hbar\sqrt{2} | 0 \frac{1}{2} \rangle + \hbar | 1 -\frac{1}{2} \rangle$

$$\Rightarrow |j = \frac{3}{2} \quad M = \frac{1}{2}\rangle = \sqrt{\frac{2}{3}} |0 \quad \frac{1}{2}\rangle + \sqrt{\frac{1}{3}} |1 \quad -\frac{1}{2}\rangle$$

on the left:  $J_- | \frac{3}{2} \quad \frac{1}{2}\rangle = \hbar \sqrt{2} | \frac{3}{2} \quad -\frac{1}{2}\rangle$

on the right:  $(L_- + S_-) \left( \sqrt{\frac{2}{3}} |0 \quad \frac{1}{2}\rangle + \sqrt{\frac{1}{3}} |1 \quad -\frac{1}{2}\rangle \right) =$   
 $= \sqrt{\frac{2}{3}} \left( \hbar \sqrt{2} | -1 \quad \frac{1}{2}\rangle + \hbar |0 \quad -\frac{1}{2}\rangle \right) + \sqrt{\frac{1}{3}} \hbar \sqrt{2} |0 \quad -\frac{1}{2}\rangle$   
 $= \frac{2}{\sqrt{3}} \hbar | -1 \quad \frac{1}{2}\rangle + 2 \sqrt{\frac{2}{3}} \hbar |0 \quad -\frac{1}{2}\rangle$

$$\Rightarrow |j = \frac{3}{2} \quad M = -\frac{1}{2}\rangle = \sqrt{\frac{2}{3}} |0 \quad -\frac{1}{2}\rangle + \sqrt{\frac{1}{3}} | -1 \quad \frac{1}{2}\rangle$$

on the left:  $J_- | \frac{3}{2} \quad -\frac{1}{2}\rangle = \hbar \sqrt{3} | \frac{3}{2} \quad -\frac{3}{2}\rangle$

on the right:  $(L_- + S_-) \left( \sqrt{\frac{2}{3}} |0 \quad -\frac{1}{2}\rangle + \sqrt{\frac{1}{3}} | -1 \quad \frac{1}{2}\rangle \right) =$   
 $= \sqrt{\frac{2}{3}} \left( \hbar \sqrt{2} | -1 \quad -\frac{1}{2}\rangle \right) + \sqrt{\frac{1}{3}} \left( \hbar | -1 \quad -\frac{1}{2}\rangle \right) = \sqrt{3} \hbar | -1 \quad -\frac{1}{2}\rangle$

$$\Rightarrow |j = \frac{3}{2} \quad M = -\frac{3}{2}\rangle = | -1 \quad -\frac{1}{2}\rangle$$

$|j = \frac{1}{2}\rangle$ :

Claim:  $|j = \frac{1}{2} \quad M = \frac{1}{2}\rangle = \sqrt{\frac{2}{3}} |1 \quad -\frac{1}{2}\rangle - \sqrt{\frac{1}{3}} |0 \quad \frac{1}{2}\rangle$

Proof:  $J_+ \quad \quad \quad = (L_+ + S_+) |j = \frac{1}{2} \quad M = \frac{1}{2}\rangle$   
 $= \sqrt{\frac{2}{3}} \left( \hbar |1 \quad \frac{1}{2}\rangle \right) - \sqrt{\frac{1}{3}} \left( \sqrt{2} \hbar |1 \quad \frac{1}{2}\rangle \right) = 0 \quad \square$

$$\Rightarrow |j = \frac{1}{2} \quad M = -\frac{1}{2}\rangle = \frac{1}{\hbar} J_- |j = \frac{1}{2} \quad M = \frac{1}{2}\rangle$$

$$= \sqrt{\frac{2}{3}} \left( \sqrt{2} |0 \quad -\frac{1}{2}\rangle \right) - \sqrt{\frac{1}{3}} \left( \sqrt{2} | -1 \quad \frac{1}{2}\rangle + |0 \quad -\frac{1}{2}\rangle \right)$$

$$= \sqrt{\frac{1}{3}} |0 \quad -\frac{1}{2}\rangle - \sqrt{\frac{2}{3}} | -1 \quad \frac{1}{2}\rangle$$

## 4.2 Addition of $> 2$ Angular Momenta

$$\vec{J} = \vec{j}_1 + \vec{j}_2 + \vec{j}_3$$

Tensor-product basis  $|j_1 j_2 j_3 m_1 m_2 m_3\rangle$

Coupling  $-\vec{j}_1 + \vec{j}_2 = \vec{J}_{12}$

$-\vec{J}_{12} + \vec{j}_3 = \vec{J}$

$$|j_1 j_2 j_3 J_{12} M_{12} m_3\rangle =$$

$$\sum_{m_1, m_2} |j_1 j_2 j_3 m_1 m_2 m_3\rangle \underbrace{\langle j_1 j_2 m_1 m_2 | j_1 j_2 J_{12} M_{12} \rangle}_{\text{Clebsch-Gordan from } \vec{j}_1 + \vec{j}_2 = \vec{J}_{12}}$$

Clebsch-Gordan from  $\vec{j}_1 + \vec{j}_2 = \vec{J}_{12}$ , does not depend on  $m_3$

Then:

$$|j_1 j_2 j_3 J_{12} J M\rangle =$$

$$\sum_{M_{12}, m_3} |j_1 j_2 j_3 J_{12} M_{12} m_3\rangle \underbrace{\langle J_{12} j_3 M_{12} m_3 | J_{12} j_3 J M \rangle}_{\text{Clebsch-Gordan from } \vec{J}_{12} + \vec{j}_3 = \vec{J}}$$

Clebsch-Gordan from  $\vec{J}_{12} + \vec{j}_3 = \vec{J}$

$$= \sum_{\substack{m_1 m_2 m_3 \\ M_{12}}} |j_1 j_2 j_3 m_1 m_2 m_3\rangle \langle j_1 j_2 m_1 m_2 | j_1 j_2 J_{12} M_{12} \rangle \langle J_{12} j_3 M_{12} m_3 | J_{12} j_3 J M \rangle$$

$$J_{12} = |j_1 - j_2| \dots j_1 + j_2$$

$$J = |J_{12} - j_3| \dots J_{12} + j_3 \Rightarrow \text{maximal } J = j_1 + j_2 + j_3$$

Alternatively:

Coupling  $\vec{j}_1 = \vec{j}_2 + \vec{j}_3 \rightarrow |j_1 j_2 j_3 J_{13} J M\rangle,$

or  $\vec{j}_2 = \vec{j}_1 + \vec{j}_3 \rightarrow |j_1 j_2 j_3 J_{23} J M\rangle$

$$|j_1 j_2 j_3 J_{23} J M\rangle = \sum_{J_{12}} |j_1 j_2 j_3 J_{12} J M\rangle \cdot \underbrace{\langle j_1 j_2 j_3 J_{12} J | j_1 j_2 j_3 J_{23} J \rangle}_{\text{independent of } M}$$

Change of coupling  $\rightarrow$  recoupling: Racah W-coefficients, Wigner 6-j symbols

4 angular momenta  $\rightarrow$  9-j symbols

### 4.3 The Wigner-Eckart Theorem

#### 4.3.1 Selection Rules

System with angular momentum  $\vec{J}$  and basis  $|rjm\rangle$  ( $r$  other quantum numbers)

$A$  is a scalar observable:  $[A, \vec{J}] = 0$

Implications for matrix elements (ME):

$$\langle rjm | A | r'j'm' \rangle = a_j(r, r') \delta_{jj'} \delta_{mm'}$$

Proof: for example  $\delta_{jj'}$ :

$$\begin{aligned} 0 &= \langle rjm | \vec{J}^2 A - A \vec{J}^2 | r'j'm' \rangle \\ &= (j(j+1) - j'(j'+1)) \hbar^2 \langle rjm | A | r'j'm' \rangle \\ \Rightarrow \text{ME is zero unless } j=j' \end{aligned}$$

$\rightarrow$  several ME = 0, others are equal

$\Rightarrow$  predictions through symmetry without solving the problem

$\vec{V}$  is a vector operator:  $[J_i, V_k] = i\hbar \epsilon_{ikl} V_l$

for example  $\vec{x}, \vec{p}, \vec{L}, \vec{S}$

$$\Leftrightarrow \langle \psi_R | V_i | \psi_R \rangle = \langle \psi | U^\dagger(R) V_i U(R) | \psi \rangle = R_{ik} \langle \psi | V_k | \psi \rangle$$

Proof:  $U(R) = e^{-\frac{i}{\hbar} \alpha \vec{n} \cdot \vec{J}}$ , infinitesimal ( $\delta = \alpha$ ):

$$\begin{aligned} U(R)^\dagger V_i U(R) &\approx \left(1 + \frac{i}{\hbar} \delta \vec{n} \cdot \vec{J}\right) V_i \left(1 - \frac{i}{\hbar} \delta \vec{n} \cdot \vec{J}\right) \\ &\approx V_i + \frac{i}{\hbar} \delta n_k \underbrace{[J_k, V_i]} \\ &= i \hbar \epsilon_{kie} V_e = -i \hbar \epsilon_{ike} V_e = -i \hbar (T_k)_{ie} V_e \\ &= (\delta_{ie} + \delta (\vec{n} \cdot \vec{T})_{ie}) V_e \equiv R_{ie} V_e \end{aligned}$$

Proportionality between ME of  $\vec{J}$  and  $\vec{V}$ : it holds

$$\langle r' j' m' | \vec{V} | r j m \rangle = v(r, r', j) \langle j' m' | \vec{J} | j m \rangle$$

Proof: application of algebra, see Cohen-Tannoudji volume 2

### 4.3.2 Irreducible Tensor Operators

Other than scalar, vector operators: there exists more general operators  $T_q^k$  with a well-defined transformation under rotation analogous to the states  $|j m\rangle \rightarrow |k q\rangle$ :

$$[J_3, T_q^k] = \hbar q T_q^k$$

$$[J_\pm, T_q^k] = \hbar \sqrt{k(k+1) - q(q\pm 1)} T_{q\pm 1}^k, \quad -k \leq q \leq k$$

vector operator corresponds to  $k=1$ :

$$T_0^1 = V_3, \quad T_\pm^1 = \mp \frac{1}{\sqrt{2}} (V_1 \pm iV_2)$$

Proof: for example

$$\begin{aligned} [J_+, T_-^1] &= [J_1 + iJ_2, \frac{1}{\sqrt{2}} (V_1 - iV_2)] = \frac{1}{\sqrt{2}} ([J_1, -iV_2] + [iJ_2, V_1]) \\ &= \frac{1}{\sqrt{2}} (-i \hbar \epsilon_{123} V_3 + i \hbar \epsilon_{213} V_3) = \hbar \sqrt{2} V_3 = \hbar \sqrt{2} T_0^1 \end{aligned}$$

$\Rightarrow T_q^k |j m\rangle$  have the same transformation properties under rotations as  $|k q\rangle \otimes |j m\rangle$   
(addition of angular momenta "k+j")

$$\Rightarrow \sum_{q,m} T_k^q |j,m\rangle \langle (kj) q m | j M \rangle$$

dropped so far

rotates like  $|j M\rangle$  ( $k+j \geq M \geq |k-j|$ )

$$\text{d. } |j M\rangle = \sum_{m,m'} |(jj') mm'\rangle \langle (jj') mm' | j M \rangle$$

### 4.3.3 The Theorem

$$\langle r j M | T_q^k | r' j' M' \rangle =$$

$$\frac{1}{\sqrt{j+1}} \langle r j || T^k || r' j' \rangle \cdot \langle (j' k) M' q | j M \rangle$$

$\uparrow$  convention       $\uparrow$  "reduced matrix elements", independent of  $M, M', q$        $\uparrow$  Clebsch-Gordan,  $\propto \delta_{q, M-M'}$

→ ratios of ME are determined in terms of Clebsch-Gordan coefficients

→ selection rules, for example  $k+j' \geq j \geq |k-j'|$

## 5. Approximation Methods

Recapitulation of time-independent perturbation theory:

$$H(\lambda) = H_0 + \lambda H_1 \quad (\lambda \rightarrow 1 \text{ at the end})$$

$$H_0 |k\rangle^{(0)} = E_0 |k\rangle^{(0)}, \quad k=1, \dots, d \rightarrow d\text{-fold degeneracy}$$

$$H(\lambda) |k, \lambda\rangle = E^{(k)}(\lambda) |k, \lambda\rangle$$

Perturbation of a set of degenerate states  $\{|k\rangle^{(0)}\}$ :

$$\begin{cases} E^{(k)}(\lambda) = E_0 + E_1^{(k)} \cdot \lambda + E_2^{(k)} \cdot \lambda^2 + \dots \\ |k, \lambda\rangle = |k\rangle^{(0)} + |k\rangle^{(1)} \cdot \lambda + |k\rangle^{(2)} \cdot \lambda^2 + \dots \end{cases} \quad k=1, \dots, d$$

The states  $\{|k\rangle^{(0)}\}$  have to be chosen such that

$$\langle l | H_1 | k \rangle^{(0)} = \text{diagonal} = \underline{\underline{E_1^{(k)}}} \cdot \delta_{kl}$$

( $\rightarrow$  diagonalize the matrix  $\langle l | H_1 | k \rangle^{(0)}$  if it is not the case)

$$\Rightarrow |k\rangle^{(1)} = - \left( H_0 - E_0 \right)^{-1} \left( H_1 - E_1^{(k)} \right) |k\rangle^{(0)}$$

$\uparrow$   
 okay, since  $\langle l | H_1 - E_1^{(k)} | k \rangle^{(0)} = 0$

Proof: insert the expansions in  $\lambda$ , compare powers of  $\lambda$

$$\text{till 1st order: } (H_0 + \lambda H_1) (|k\rangle^{(0)} + \lambda |k\rangle^{(1)}) =$$

$$(E_0 + \lambda E_1^{(k)}) (|k\rangle^{(0)} + \lambda |k\rangle^{(1)})$$

$$\Rightarrow O(\lambda^1): H_1 |k\rangle^{(0)} + H_0 |k\rangle^{(1)} = E_1^{(k)} |k\rangle^{(0)} + E_0 |k\rangle^{(1)}$$

$$(H_0 - E_0) |k\rangle^{(1)} = - (H_1 - E_1^{(k)}) |k\rangle^{(0)}$$

$$\text{Explicitly: } |k\rangle^{(1)} = \sum_m' \frac{|m\rangle \langle m | H_1 - E_1^{(k)} | k \rangle^{(0)}}{E_0 - E_m}$$

$\square$   
 sum over eigenstate of  $H_0$  other than  $\{|k\rangle^{(0)}\}$ .

# 5.1 Time-dependent Perturbation Theory

$$H(t) = H_0 + H_1(t)$$

$$\frac{\partial H_0}{\partial t} = 0 ;$$

$$H_0 |n\rangle = \epsilon_n |n\rangle$$

$$\langle n|m\rangle = \delta_{nm} , \quad \mathbb{1} = \sum_n |n\rangle \langle n|$$

$H_0$ : stationary system

- here: discrete spectrum

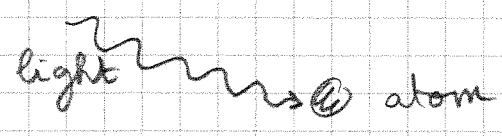
- later: continuum spectrum of final states

$H_1(t)$ : perturbation

- activation operation, for example  $H_1(t) = 0$  for  $t \leq 0$

- external fields

transition probabilities between stationary states without  $H_1$ , for example radiative transitions



## 5.1.1 Perturbation Theory in the Schrödinger Picture

$$H_1(t) = 0 \text{ for } t \leq 0$$

$$|\psi(t=0)\rangle = |i\rangle \text{ initial state}$$

→ solve  $i\hbar \frac{d}{dt} |\psi(t)\rangle = (H_0 + H_1(t)) |\psi(t)\rangle$  for  $t \geq 0$

$$P_{i \neq f}(t) = |\langle f | \psi(t) \rangle|^2 = \text{transition probability}$$

Matrix representation in the basis of  $H_0$ -eigenstates:

$$|\psi(t)\rangle = \sum_n c_n(t) |n\rangle , \quad c_n(t) = \langle n | \psi(t) \rangle$$



$$\langle m | H_0 | n \rangle = \epsilon_m \delta_{mn}$$

$$\langle m | H_1(t) | n \rangle =: W_{mn}(t)$$

$$\Rightarrow i\hbar \sum_m \dot{c}_m(t) |m\rangle = \sum_m c_m(t) (\epsilon_m |m\rangle + H_1(t) |m\rangle) \quad (\dot{\phantom{x}} \equiv \frac{d}{dt})$$

$$\langle n | \Rightarrow i\hbar \dot{c}_n = \epsilon_n c_n + \sum_m W_{nm}(t) c_m \quad (*)$$

$$c_n(0) = \delta_{ni}$$

$$P_{if} = |c_f(t)|^2$$

Solution of (\*): write  $c_n(t) = e^{-\frac{i}{\hbar} \epsilon_n t} b_n(t)$   
 solution without  $H_1$

Insert into (\*) and multiply by  $e^{+\frac{i}{\hbar} \epsilon_n t}$ :

$$\Rightarrow i\hbar \dot{b}_n = \sum_m W_{nm}(t) e^{i\omega_{nm}t} b_m(t)$$

$$\omega_{nm} = \frac{\epsilon_n - \epsilon_m}{\hbar} = \text{Bohr frequencies}$$

Expansion in powers of  $H_1$  ( $H_1 \rightarrow \lambda H_1$  in  $\lambda$ ):

$$b_n(t) = b_n^{(0)}(t) + b_n^{(1)}(t) + b_n^{(2)}(t) + \dots$$

$$(O(\lambda^0) \quad O(\lambda) \quad O(\lambda^2) \quad \dots)$$

$$i\hbar (\dot{b}_n^{(0)}(t) + \dot{b}_n^{(1)}(t) + \dots) = \sum_m W_{nm}(t) e^{i\omega_{nm}t} (b_m^{(0)}(t) + \dots)$$

$$\Rightarrow O(\lambda^0): \dot{b}_n^{(0)}(t) = 0$$

$$O(\lambda^r): i\hbar \dot{b}_n^{(r)}(t) = \sum_m W_{nm}(t) e^{i\omega_{nm}t} b_m^{(r-1)}(t)$$

Initial condition:  $b_n^{(0)}(0) = \delta_{ni}$ ,  $b_n^{(r)}(0) = 0$  for  $r > 0$ .

$$\Rightarrow O(\lambda^0): b_n^{(0)}(t) = \delta_{ni}$$

$$O(\lambda^1): i\hbar \dot{b}_n^{(1)}(t) = W_{ni}(t) \cdot e^{i\omega_{ni}t}$$

In leading order for  $n \neq i$ :

$$b_n^{(1)}(t) = -\frac{i}{\hbar} \int_0^t dt' W_{ni}(t') e^{i\omega_{ni}t'}$$

$$P_{if}(t) = \underset{f \neq i}{=} |b_f^{(1)}(t)|^2 = \frac{1}{\hbar^2} \left| \int_0^t dt' W_{fi}(t') e^{i\omega_{fi}t'} \right|^2 + O(H_1^3) \quad (f \neq i)$$

## 5.1.2 The Interaction Picture

Time-evolution operator:  $|\psi(t)\rangle = U(t, t_0) |\psi(t_0)\rangle$

defined through:

$$- U(t_0, t_0) = \mathbb{1}$$

$$- i\hbar \frac{\partial}{\partial t} U(t, t_0) = H(t) U(t, t_0) \quad (\text{Schrödinger's equation})$$

$$\Rightarrow U \text{ is unitary: } U^\dagger(t_0, t_0) U(t_0, t_0) = \mathbb{1}$$

$$\frac{\partial}{\partial t} U^\dagger(t, t_0) U(t, t_0) = 0 \quad \left( \Leftrightarrow \frac{\partial}{\partial t} U^\dagger = \frac{i}{\hbar} U^\dagger H \right)$$

similarly for  $U U^\dagger$

$$\Rightarrow U(t_3, t_2) U(t_2, t_1) = U(t_3, t_1)$$

$$\Rightarrow U^{-1}(t, t_0) = U^\dagger(t, t_0) = U(t_0, t)$$

$$\Rightarrow U(t, t_0) = \mathbb{1} - \frac{i}{\hbar} \int_0^t dt' H(t') U(t', t_0)$$

(integral representation)

Special case:  $H = H_0$  time-independent:

$$U(t, t_0) = e^{-\frac{i}{\hbar} H_0 (t-t_0)} = \sum_n e^{-\frac{i}{\hbar} \epsilon_n (t-t_0)} |n\rangle \langle n|$$

In general:

$$U = \mathbb{1} + \sum_{n=1}^{\infty} U^{(n)} \quad \text{with}$$

$$U^{(n)}(t, t_0) = \left(\frac{-i}{\hbar}\right)^n \int_{t_0}^t dt_1 \int_{t_0}^{t_1} dt_2 \cdots \int_{t_0}^{t_{n-1}} dt_n H(t_1) H(t_2) \cdots H(t_n)$$

Dyson perturbation theory (F. J. Dyson, Phys. Rev. 75, 1949)

Proof: recursive use of the integral representation:

$$\begin{aligned} U(t, t_0) &= \mathbb{1} - \frac{i}{\hbar} \int_{t_0}^t dt_1 H(t_1) U(t_1, t_0) \\ &= \mathbb{1} - \frac{i}{\hbar} \int_{t_0}^t dt_1 H(t_1) \left( \mathbb{1} - \int_{t_0}^{t_1} dt_2 H(t_2) U(t_2, t_0) \right) \\ &= \dots \end{aligned}$$

□

We can rewrite

$$U^{(n)}(t, t_0) = \left(\frac{-i}{\hbar}\right)^n \frac{1}{n!} \int_{t_0}^t dt_1 \int_{t_0}^{t_1} dt_2 \cdots \int_{t_0}^{t_{n-1}} dt_n \overline{T} (H(t_1) H(t_2) \cdots H(t_n))$$

$\overline{T}$ : time-ordering operator

$$\overline{T} (H(t_1) \cdots H(t_n)) = H(t_{\pi(1)}) H(t_{\pi(2)}) \cdots H(t_{\pi(n)})$$

where  $\pi$  = permutation of the indices  $1, 2, \dots, n$  such that

$$t_{\pi(1)} \geq t_{\pi(2)} \geq \dots \geq t_{\pi(n)}$$

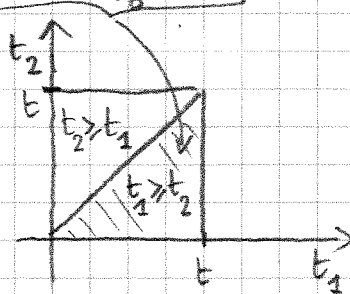
-  $\overline{T} (H(t_1) \cdots H(t_n))$  is symmetric with respect to the sequence of the arguments  $t_1 \dots t_n$  by construction

$$\int_{t_0}^t dt_1 \int_{t_0}^{t_1} dt_2 \cdots \int_{t_0}^{t_{n-1}} dt_n F(t_1, \dots, t_n) = \frac{1}{n!} \int_{t_0}^t dt_1 \int_{t_0}^{t_1} dt_2 \cdots \int_{t_0}^{t_{n-1}} dt_n F$$

$\uparrow$   
F symmetric

Proof:

$$\int_{t_0}^t dt_1 \int_{t_0}^{t_1} dt_2 H(t_1) H(t_2) = \frac{1}{2} \left( \int_{t_0}^t dt_1 \int_{t_0}^{t_1} dt_2 H(t_1) H(t_2) \right) \{t_1 \geq t_2\} + \int_{t_0}^t dt_2 \int_{t_0}^{t_2} dt_1 H(t_2) H(t_1) \{t_2 \geq t_1\}$$

$$= \frac{1}{2} \int_{t_0}^t dt_1 \int_{t_0}^t dt_2 T(H(t_1) H(t_2))$$


generalize in  $n$ ,  $n!$  = number of permutations of  $t_1, \dots, t_n$ .  $\square$

$$\Rightarrow U(t, t_0) = T \left( e^{-\frac{i}{\hbar} \int_{t_0}^t dt' H(t')} \right)$$

So far: Schrödinger picture:  $|\psi_S(t)\rangle$ ,  $\frac{\partial A_S}{\partial t} = 0$  ( $A_S$  = observable)

Heisenberg picture:

$$|\psi_H\rangle = U^\dagger(t, t_0) |\psi_S(t)\rangle = |\psi_S(t_0)\rangle$$

$$A_H(t) = U^\dagger(t, t_0) A_S U(t, t_0)$$

$$i\hbar \frac{\partial A_H}{\partial t} = [A_H(t), H_H(t)] \quad \text{Heisenberg equation}$$

$$\left( \text{if } \frac{\partial H}{\partial t} = 0, \text{ then } H_H = H \right)$$

$$\text{Equal matrix elements: } \langle \psi_H | A_H(t) | \varphi_H \rangle = \langle \psi_S(t) | A_S | \varphi_S(t) \rangle$$

Dirac or Interaction picture:

$$U_0(t, t_0) = e^{-\frac{i}{\hbar} H_0 \cdot (t-t_0)} \quad \text{free dynamics}$$

$$A_D(t) = U_0^\dagger(t, t_0) A_S U_0(t, t_0)$$

$$\text{Require } \langle \psi_D(t) | A_D(t) | \varphi_D(t) \rangle = \langle \psi_S(t) | A_S | \varphi_S(t) \rangle$$

$$\Rightarrow |\psi_D(t)\rangle = U_0^\dagger(t, t_0) |\psi_S(t)\rangle = U_0^\dagger(t, t_0) U(t, t_0) |\psi_H\rangle$$

$$\Rightarrow i\hbar \frac{\partial}{\partial t} |\Psi_D(t)\rangle = U_0^\dagger(t, t_0) \underbrace{(H - H_0)}_{= H_1} U(t, t_0) |\Psi_H\rangle$$

$$=: H_D(t) |\Psi_D(t)\rangle$$

$$H_D(t) = U_0^\dagger(t, t_0) H_1(t) U_0(t, t_0)$$

Define  $U_D(t, t_0) = U_0^\dagger(t, t_0) U(t, t_0)$ ; then  $|\Psi_D(t)\rangle = U_D(t, t_0) |\Psi_H\rangle$

$$\Rightarrow i\hbar \frac{\partial}{\partial t} U_D(t, t_0) = H_D(t) U_D(t, t_0)$$

$$\Rightarrow U_D(t, t_0) = T \left( e^{-\frac{i}{\hbar} \int_0^t dt' H_D(t')} \right)$$

Expansion yields formally all orders of perturbation theory.

### 5.1.3 The Perturbation Series in the Interaction Picture

$$P_{if}(t) = \left| \langle f | \Psi_S(t) \rangle \right|^2$$

$$= \left| \langle f | U(t, 0) | i \rangle \right|^2$$

$$= \left| \underbrace{\langle f | U_0^\dagger(t, 0) U(t, 0) | i \rangle}_{= e^{+\frac{i}{\hbar} \epsilon_f t} \langle f |} \right|^2 = \left| \langle f | U_D(t, 0) | i \rangle \right|^2$$

With Dyson perturbation series:

$$\langle f | U_D | i \rangle = \delta_{fi} + \sum_{n=1}^{\infty} \langle f | U_D^{(n)} | i \rangle$$

1st order:

$$\langle f | U_D^{(1)} | i \rangle = -\frac{i}{\hbar} \int_0^t dt' \langle f | H_D(t') | i \rangle$$

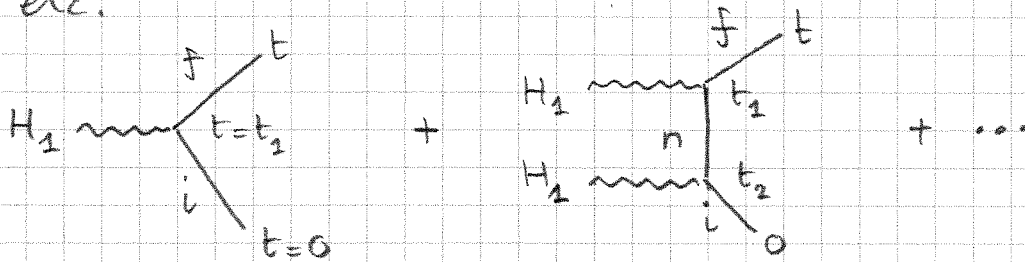
$$= -\frac{i}{\hbar} \int_0^t dt' e^{+\frac{i}{\hbar} \epsilon_f t'} \langle f | H_1(t') | i \rangle e^{-\frac{i}{\hbar} \epsilon_i t'}$$

$$= -\frac{i}{\hbar} \int_0^t dt' e^{i\omega_{fi} t'} W_{fi}(t'), \quad \omega_{fi} = \frac{\epsilon_f - \epsilon_i}{\hbar}, \text{ like before}$$

2nd order:

$$\begin{aligned}
\langle f | U_D^{(2)} | i \rangle &= \left( \frac{-i}{\hbar} \right)^2 \int_0^t dt_1 \int_0^{t_1} dt_2 \langle f | H_D(t_1) H_D(t_2) | i \rangle \\
&= \left( \frac{-i}{\hbar} \right)^2 \sum_n \int_0^t dt_1 \int_0^{t_1} dt_2 e^{+\frac{i}{\hbar} \epsilon_f t_1} \langle f | H_1(t_1) | n \rangle e^{-\frac{i}{\hbar} \epsilon_n t_1} \times \\
&\quad e^{+\frac{i}{\hbar} \epsilon_n t_2} \langle n | H_1(t_2) | i \rangle e^{-\frac{i}{\hbar} \epsilon_i t_2} \\
&= \left( \frac{-i}{\hbar} \right)^2 \sum_n \int_0^t dt_1 \int_0^{t_1} dt_2 e^{i \omega_{fn} t_1} W_{fn}(t_1) e^{i \omega_{ni} t_2} W_{ni}(t_2)
\end{aligned}$$

etc.

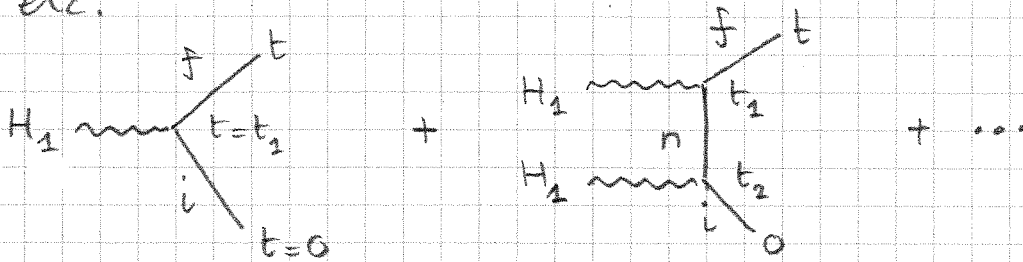


Sum over discrete scatterings; virtual intermediate states.

2nd order:

$$\begin{aligned}
 \langle f | U_D^{(2)} | i \rangle &= \left( \frac{-i}{\hbar} \right)^2 \int_0^t dt_1 \int_0^{t_1} dt_2 \langle f | H_D(t_1) H_D(t_2) | i \rangle \\
 &= \left( \frac{i}{\hbar} \right)^2 \sum_n \int_0^t dt_1 \int_0^{t_1} dt_2 e^{+\frac{i}{\hbar} \epsilon_f t_1} \langle f | H_1(t_1) | n \rangle e^{-\frac{i}{\hbar} \epsilon_n t_1} \times \\
 &\quad e^{+\frac{i}{\hbar} \epsilon_n t_2} \langle n | H_1(t_2) | i \rangle e^{-\frac{i}{\hbar} \epsilon_i t_2} \\
 &= \left( \frac{i}{\hbar} \right)^2 \sum_n \int_0^t dt_1 \int_0^{t_1} dt_2 e^{i \omega_{fn} t_1} W_{fn}(t_1) e^{i \omega_{ni} t_2} W_{ni}(t_2)
 \end{aligned}$$

etc.



Sum over discrete scatterings, virtual intermediate states.

Stopping criterium (sufficient):

$$\frac{t}{\hbar} |\langle m | H_1(t) | n \rangle| \ll 1$$

## 5.1.4 Fermi's Golden Rule

Let  $\langle f | i \rangle = 0$ , transition  $i \rightarrow f$  occurs at 1st order. $H_1$  is independent of  $t$ , or rather

$$H_1(t) = \begin{cases} H_1 & \text{for } 0 \leq t \leq T \\ 0 & \text{otherwise} \end{cases}$$

 $\rightarrow$  then it is natural to consider eigenstates of  $H_0$  at  $t=0, T$ 

$$P_{if} \approx \frac{1}{\hbar^2} |\langle f | H_1 | i \rangle|^2 f\left(T, \frac{\epsilon_f - \epsilon_i}{\hbar}\right)$$

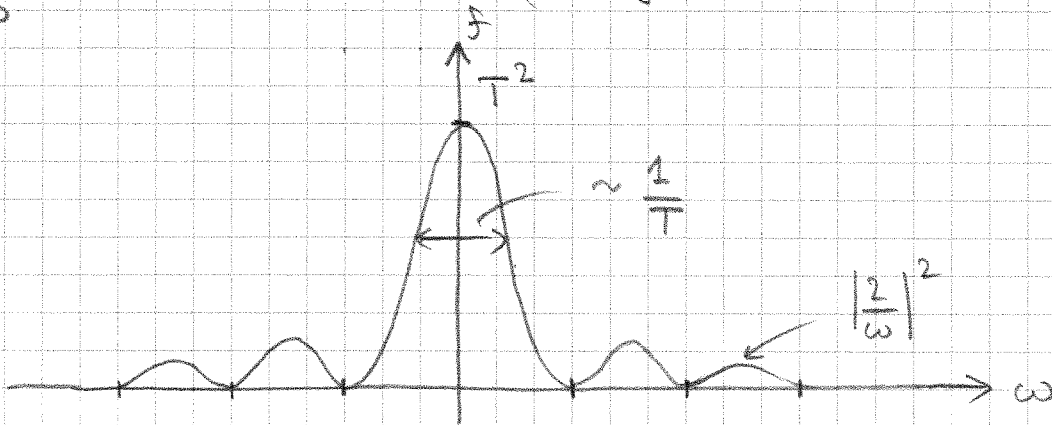
$$f(T, \omega) = \left| \int_0^T dt e^{it\omega} \right|^2 = \left| \frac{2}{\omega} \sin\left(\frac{\omega T}{2}\right) \right|^2$$

(1st order)

$$\lim_{\omega \rightarrow 0} f(T, \omega) =: f(T, 0) = T^2$$

$$f(T, \omega) = 0 \quad \text{for} \quad \omega = \frac{2\pi \cdot n}{T}$$

$$\int_{-\infty}^{\infty} d\omega f(T, \omega) = 2\pi T \quad \left( \Leftrightarrow \int d\omega e^{i(t-t')\omega} = 2\pi \delta(t-t') \right)$$



$$\Rightarrow \frac{1}{2\pi T} f(T, \omega) \xrightarrow{T \rightarrow \infty} \delta(\omega)$$

$\hookrightarrow$  representation of  $\delta$ -function with height  $\sim \frac{T}{2\pi}$   
width  $\sim \frac{1}{T}$

$\Rightarrow$  for large  $T$  transitions with

$$|\epsilon_f - \epsilon_i| \lesssim \frac{\hbar}{T} \quad \text{are favoured}$$

$$\text{where: } P_{if} \approx \frac{1}{\hbar^2} |\langle f | H_1 | i \rangle|^2 T^2$$

$\ll 1$  (validity condition of 1st order)

$\lim_{T \rightarrow \infty}$  is meaningful for transitions to a continuum of states:  
(since for discrete states it holds  $\lim_{T \rightarrow \infty} \frac{1}{T} P_{if} \sim \delta_{fi}$ )

let  $F$  be a continuum of final states with energies close to  $\epsilon_f$

$$P_F = \int d\mathcal{F} |\mathcal{F}\rangle \langle \mathcal{F}| = \text{projector on } F$$

$$T_{i \rightarrow F} = \frac{1}{T} \int_F d\mathcal{F} P_{if} \quad : \quad \frac{\text{transitions}}{\text{time}} = \text{rate}$$



$$T_{i \rightarrow f} \stackrel{T \rightarrow \infty}{=} \int_{\mathbb{F}} d\mathbb{F} \frac{1}{\hbar^2} |\langle \mathbb{F} | H_2 | i \rangle|^2 \underbrace{2\pi \delta\left(\frac{\epsilon_{\mathbb{F}} - \epsilon_i}{\hbar}\right)}_{= \frac{\mathbb{F}}{\hbar}}$$

$$= \frac{2\pi}{\hbar} |\langle \mathbb{F} | H_2 | i \rangle|^2 g(\epsilon_i) \quad \text{Fermi's golden rule}$$

$$g(\epsilon) = \int_{\mathbb{F}} d\mathbb{F} \delta(\epsilon - \epsilon_{\mathbb{F}}) = \text{density of states}$$

We used the assumption:  $|\langle \mathbb{F} | H_2 | i \rangle|$  is constant on the "shell" of  $|\mathbb{F}\rangle$  with  $\epsilon_{\mathbb{F}} = \epsilon_i$

### 5.1.5 Periodic Perturbation

Example: atom in the electromagnetic field

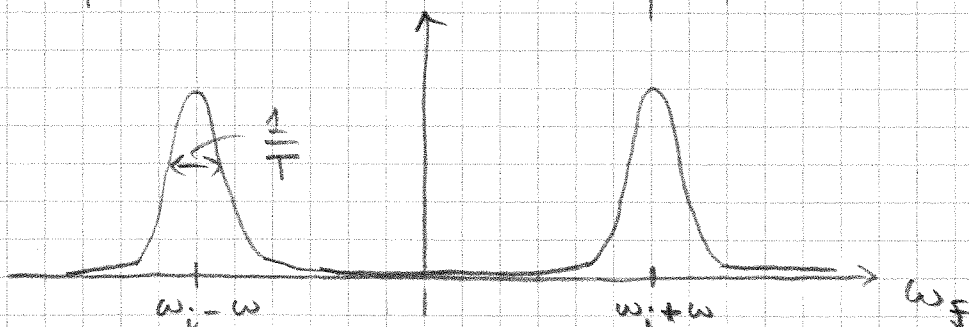
$$H_2(t) = V \cos(\omega t) \quad (\text{monochromatic perturbation})$$

$$P_{i \rightarrow f} = \frac{1}{\hbar^2} \left| \int_0^T dt e^{\frac{i}{\hbar}(\epsilon_{\mathbb{F}} - \epsilon_i) \cdot t} \cos(\omega t) \langle \mathbb{F} | V | i \rangle \right|^2$$

$$= \frac{1}{4\hbar^2} |\langle \mathbb{F} | V | i \rangle|^2 \left| \int_0^T dt \left\{ e^{i(\omega_{\mathbb{F}i} + \omega)t} + e^{i(\omega_{\mathbb{F}i} - \omega)t} \right\} \right|^2$$

$$\omega_{\mathbb{F}i} = \frac{\epsilon_{\mathbb{F}} - \epsilon_i}{\hbar} \quad \text{Bohr frequency}$$

- integrals are similar as before
- contributions when  $\omega_{\mathbb{F}i} + \omega = 0 \left(\frac{1}{T}\right)$  or  $\omega_{\mathbb{F}i} - \omega = 0 \left(\frac{1}{T}\right)$
- well separated when  $\omega \gg \frac{1}{T}$ , no mixing contributions



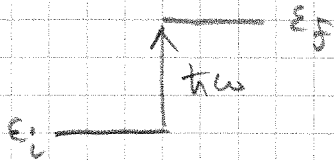
- therefore peaks when  $\epsilon_{\mathbb{F}} = \epsilon_i \pm \hbar\omega$

$$P_{if} \approx \frac{1}{4\hbar^2} |\langle f | V | i \rangle|^2 \left( f(T, \omega_{fi} + \omega) + f(T, \omega_{fi} - \omega) \right)$$

Consider groups of continuum final states  $F_{\pm}$  with energies  $\approx \epsilon_i \pm \hbar\omega$

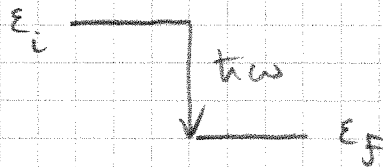
$$\Gamma_{i \rightarrow F_+} = \frac{\pi}{2\hbar} |\langle F_+ | V | i \rangle|^2 \rho(\epsilon_i + \hbar\omega)$$

absorption



$$\Gamma_{i \rightarrow F_-} = \frac{\pi}{2\hbar} |\langle F_- | V | i \rangle|^2 \rho(\epsilon_i - \hbar\omega)$$

stimulated emission



This works provided that  $\rho(\epsilon_i \pm \hbar\omega)$  is large: resonance, depends on the spectrum

## 6. Quantum Mechanics as a Path Integral

Formalism proposed by P. D. Feynman in his Ph.D. thesis

- alternative but equivalent representation of QM
- no operators! instead functional integrals
- some problems are easier to handle
- classical limit: very intuitive
- starting point for several approximation methods and numerical simulations, in particular for Quantum Field Theorie.

In the following:  $H = \frac{\hat{p}^2}{2m} + \underbrace{V(\hat{x})}_{\text{time-independent}}$ , for now  $D=1$  (dimensions)

$H|n\rangle = E_n|n\rangle$  energie-eigenstates  $|n\rangle, |m\rangle, \dots$  (notation)

$\hat{x}|y\rangle = y|y\rangle$  position-eigenstates  $|x\rangle, |y\rangle, |z\rangle, \dots$

$\hat{p}|q\rangle = q|q\rangle$  momentum-eigenstates  $|p\rangle, |q\rangle, \dots$

Normalization and completeness:

$$\langle n|m\rangle = \delta_{nm}, \quad \mathbb{1} = \sum_n |n\rangle \langle n|$$

$$\langle x|y\rangle = \delta(x-y), \quad \mathbb{1} = \int dx |x\rangle \langle x|$$

$$\langle p|q\rangle = \delta(p-q), \quad \mathbb{1} = \int dp |p\rangle \langle p|$$

Wave functions of the momentum-eigenstates: plane waves

$$\langle x|p\rangle = \frac{1}{\sqrt{2\pi\hbar}} e^{\frac{i}{\hbar} p \cdot x}$$

Bemerkung: non-normalizable continuum states like  $|x\rangle, |p\rangle$  are easier to manipulate than physical wave packets.

## 6.1. The Path Integral Representation of Transition Amplitudes

$$U(t_f, x_f; t_i, x_i) := \langle x_f | e^{-\frac{i}{\hbar} H(t_f - t_i)} | x_i \rangle$$

also called: propagator or evolution kernel

$$\text{Wave function: } \Psi(t_f, x_f) = \int dx_i U(t_f, x_f; t_i, x_i) \Psi(t_i, x_i)$$

$$\text{since } \langle x_f | \Psi(t_f) \rangle = \langle x_f | e^{-\frac{i}{\hbar} H(t_f - t_i)} | \Psi(t_i) \rangle$$

$$= \int dx_i \underbrace{\langle x_f | e^{-\frac{i}{\hbar} H(t_f - t_i)} | x_i \rangle}_U \underbrace{\langle x_i | \Psi(t_i) \rangle}_{\Psi(t_i, x_i)}$$

Example: the free particle:  $H = \frac{\hat{p}^2}{2m}$

$$U = \langle x_f | e^{-\frac{i}{\hbar} \frac{\hat{p}^2}{2m} T} | x_i \rangle, \quad T = t_f - t_i$$

$$= \int dp \langle x_f | p \rangle e^{-\frac{i}{\hbar} \frac{p^2}{2m} T} \langle p | x_i \rangle$$

$$= \frac{1}{2\pi\hbar} \int dp e^{-\frac{i}{\hbar} \frac{p^2}{2m} T + \frac{i}{\hbar} p(x_f - x_i)}$$

Gauss integral with imaginary exponent

$$\text{Classic Gauss integrals: } \int_{-\infty}^{\infty} dx e^{-\frac{x^2}{2}} = \sqrt{2\pi}$$

$$\Rightarrow \int_{-\infty}^{\infty} dx e^{-\frac{ax^2}{2}} = \sqrt{\frac{2\pi}{a}} \quad \text{for } a > 0$$

Can be analytically continued in the domain  $a \in \mathbb{C}$ ,  $\text{Re } a > 0$ .

The boundary value with  $a = iA$  is

$$\int_{-\infty}^{\infty} dx e^{-iA \frac{x^2}{2}} = \sqrt{\frac{2\pi}{iA}}, \quad A \in \mathbb{R}.$$

With an additional linear term in the exponent (completing the square)

$$\int_{-\infty}^{\infty} dx e^{-\frac{ax^2}{2} + bx} = \int_{-\infty}^{\infty} dx e^{-\frac{a}{2} \left(x - \frac{b}{a}\right)^2 + \frac{b^2}{2a}} = \sqrt{\frac{2\pi}{a}} e^{\frac{b^2}{2a}}$$

$$\Rightarrow U = \langle x_f | e^{-\frac{i}{\hbar} \frac{\hat{p}^2}{2m} T} | x_i \rangle = \left( \frac{2\pi i \hbar T}{m} \right)^{-1/2} e^{\frac{i m}{2\hbar T} (x_f - x_i)^2}$$

(  $a = i \frac{T}{\hbar m}$ ,  $b = \frac{i}{\hbar} (x_f - x_i)$  above )

Now  $V(\hat{x}) \neq 0$ :

There is no simple spectral representation for  $e^{-\frac{i}{\hbar} H T}$ .

For small  $T = \tau$ :

$$e^{-\frac{i}{\hbar} \tau \left[ \frac{\hat{p}^2}{2m} + V \right]} \approx e^{-\frac{i}{\hbar} \tau V} e^{-\frac{i}{\hbar} \frac{\hat{p}^2}{2m} \tau} e^{-\frac{i}{\hbar} \tau V} + O(\tau^3)$$

(Baker, Campbell, Hausdorff  $\rightarrow$  exercises)

$$\begin{aligned} \langle x_f | e^{-\frac{i}{\hbar} H T} | x_i \rangle &\approx \langle x_f | e^{-\frac{i}{\hbar} \tau V} e^{-\frac{i}{\hbar} \frac{\hat{p}^2}{2m} \tau} e^{-\frac{i}{\hbar} \tau V} | x_i \rangle \\ &= \left( \frac{2\pi i \hbar \tau}{m} \right)^{-1/2} e^{\frac{i m}{2\hbar \tau} (x_f - x_i)^2 - \frac{i}{\hbar} \frac{\tau}{2} (V(x_i) + V(x_f))} \end{aligned}$$

calculation as before: result valid only for  $\tau \rightarrow 0$

Partition of the time interval  $[t_i, t_f]$  in  $N$  subintervals

$$T = t_f - t_i = N \cdot \tau$$

$$\Rightarrow U = \langle x_f | e^{-\frac{i}{\hbar} H T} | x_i \rangle = \langle x_f | \left( e^{-\frac{i}{\hbar} H \tau} \right)^N | x_i \rangle$$

completeness relations

$$= \int dx_1 \dots \int dx_{N-1} \langle x_f | e^{-\frac{i}{\hbar} H \tau} | x_{N-1} \rangle \langle x_{N-1} | e^{-\frac{i}{\hbar} H \tau} | x_{N-2} \rangle$$

$$\approx \left( \frac{2\pi i \hbar \tau}{m} \right)^{-\frac{N}{2}} \left( \prod_{k=1}^{N-1} dx_k \right) \prod_{k=0}^{N-1} e^{\frac{i m}{2\hbar \tau} (x_{k+1} - x_k)^2 - \frac{i}{\hbar} \frac{\tau}{2} (V(x_{k+1}) + V(x_k))}$$

with  $|x_i\rangle \equiv |x_0\rangle$

$|x_f\rangle \equiv |x_N\rangle$

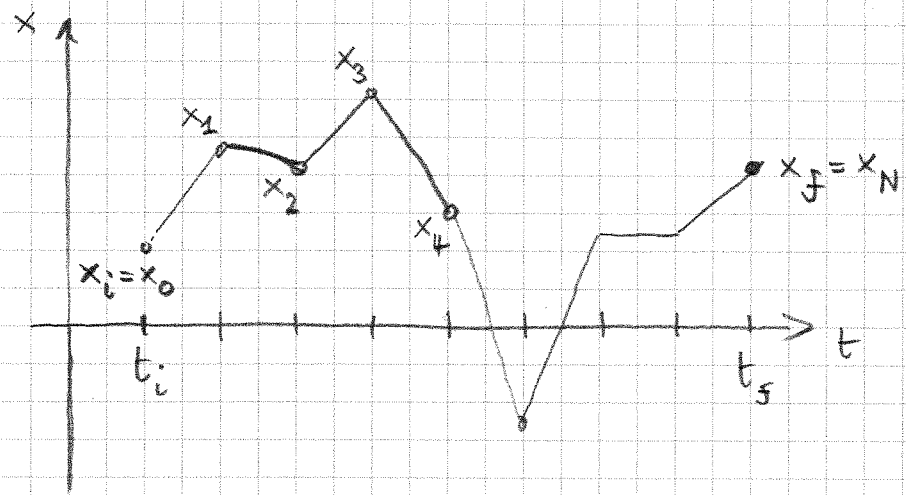
$$+ O(N \cdot \tau^3)$$

$$= O(\tau^2)$$

(Lie-Kato-Grotrian product formula)

Interpretation for  $N \rightarrow \infty$ ,  $\tau \rightarrow 0$ ,  $T = N \cdot \tau$  fixed:

(68)



$x(t) \approx$  trajectory or path

Integration over all paths:

$$\left[ \int_{-\infty}^{\infty} \prod_k dx_k \right]^{-\frac{N}{2}} =: \int \mathcal{D}x \text{ or } \int \mathcal{D}[x(t)]$$

$\tau \rightarrow 0$ : functional integral: integration over all continuous functions  $x(t)$  with  $x(t_i) = x_i$  and  $x(t_f) = x_f$

Integrand:  $\prod_k e^{\dots} = e^{\sum \dots}$

$$\sum_k \tau \cdot \mathcal{F}(x_k) \xrightarrow{\tau \rightarrow 0} \int_{t_i}^{t_f} dt \mathcal{F}(x(t))$$

$$\frac{x_{k+1} - x_k}{\tau} \xrightarrow{\tau \rightarrow 0} \dot{x}(t)$$

$$\Rightarrow U = \int \mathcal{D}x e^{\frac{i}{\hbar} \int_{t_i}^{t_f} \left( \frac{m}{2} \dot{x}^2 - V(x) \right) dt}$$

Classical mechanics: action along a trajectory

$$S[x(t)] = \int_{t_i}^{t_f} \left[ \frac{m}{2} \dot{x}^2(t) - V(x(t)) \right] dt$$

$$\Rightarrow U = \langle x_f | e^{-\frac{i}{\hbar} H(t_f - t_i)} | x_i \rangle = \int \mathcal{D}x e^{\frac{i}{\hbar} S[x(t)]}$$

- $\lim_{T \rightarrow 0}$  is mathematically problematic:
  - $x(t)$  continuous but in general not differentiable  $\rightarrow$  what is  $\dot{x}$ ?
  - in Quantum Mechanics: Wiener measure
  - for QFT the existence of  $T \rightarrow 0$  is in general not proven

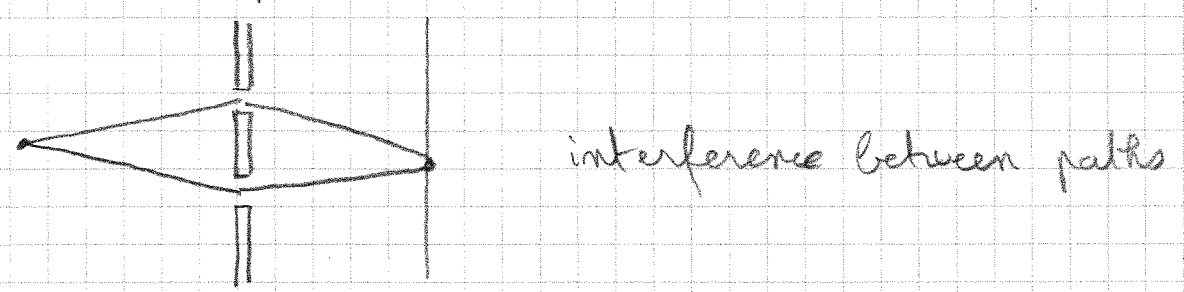
- classical limit:  $S \gg \hbar$  or " $\hbar \rightarrow 0$ "  
 dominant contribution comes from the trajectories for which  $e^{\frac{i}{\hbar}S}$  has the slowest oscillations

$\rightarrow \frac{\delta S}{\delta x(t)} \Big|_{x(t_i)=x_i, x(t_f)=x_f} = 0 =$  classical principle of least action

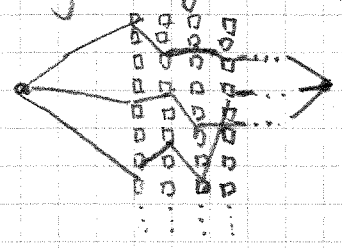
$\Rightarrow \frac{d}{dt} \frac{\partial L}{\partial \dot{x}} - \frac{\partial L}{\partial x} = 0$  ,  $L = \frac{m}{2} \dot{x}^2 - V(x)$   
 Euler-Lagrange eqs.      Lagrange function

In QM all paths contribute even the classically forbidden  $\rightarrow$  tunneling etc.

Illustration of the principle. double-slit experiment:

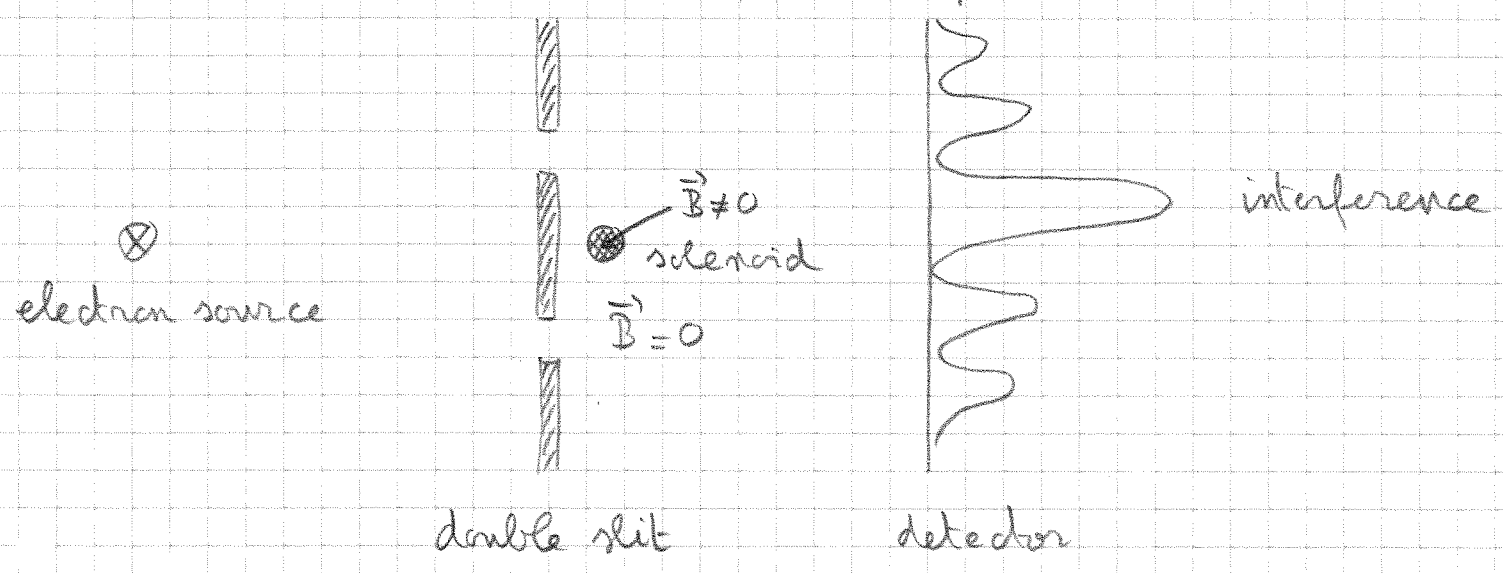


Limit of infinitely many "infinite-slit" experiments:



# 6.2 The Aharonov - Bohm Effect

D. Bohm and Y. Aharonov, 1959 prediction



- \* magnetic solenoid  $\nabla$  generates  $\vec{B} \neq 0$ ,  $\vec{B}$  parallel to slit confined inside the solenoid, outside  $\vec{B} = 0$
- \* solenoid is much smaller than distance of the slit, wave function of electron is practically zero there
- \*  $\vec{B}$  on or  $\vec{B}$  out changes the interference pattern!
- \* this effect has been observed in several experiments

$\vec{B} = \nabla \times \vec{A}$ ; outside the solenoid is  $\vec{B} = 0$  but for example

$$\vec{A}(\vec{x}) = \frac{\Phi}{2\pi} \frac{1}{x_1^2 + x_2^2} \begin{pmatrix} -x_2 \\ x_1 \\ 0 \end{pmatrix} \Rightarrow \vec{B} = \begin{pmatrix} 0 \\ 0 \\ 0 \end{pmatrix}$$

Does  $\vec{A}$  perhaps affect the interference pattern? but  $\vec{A}$  unphysical.

Classical mechanics: no effect, since  $m\ddot{\vec{x}} = e\dot{\vec{x}} \times \vec{B}$   
 $\Rightarrow \vec{B} = 0 \Leftrightarrow$  no force on the electrons



Lagrange function and action:

$$L = \frac{m}{2} \dot{\vec{x}}^2 + e \dot{\vec{x}} \cdot \vec{A}(\vec{x}) = \underbrace{L_0}_{\text{free electron}} + e \dot{\vec{x}} \cdot \vec{A}(\vec{x})$$

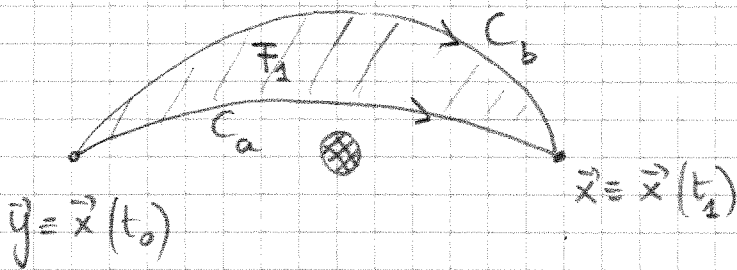
$$\Rightarrow S = S_0 + e \int_{t_1}^{t_2} \vec{A}(\vec{x}(t)) \cdot \dot{\vec{x}}(t) dt$$

$$= S_0 + e \int_{\vec{x}(t_1)}^{\vec{x}(t_2)} \vec{A}(\vec{x}) \cdot d\vec{x} \quad \text{line integral}$$

QM: all paths contribute to path integral

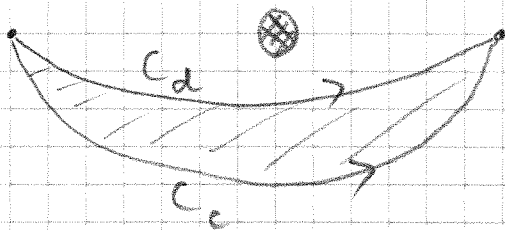
We can distinguish two categories of paths, (1) and (2), depending whether they go left or right of the solenoid

① consider two paths  $C_a$  and  $C_b$  of type (1):



$F_1 =$  enclosed surface

② consider two paths  $C_c$  and  $C_d$  of type (2):



$F_2 =$  enclosed surface

$$\textcircled{1}: \int_{C_a} \vec{A} \cdot d\vec{x} - \int_{C_b} \vec{A} \cdot d\vec{x} = \oint_{C_a - C_b} \vec{A} \cdot d\vec{x} = \int_{F_1} (\nabla \times \vec{A}) \cdot d\vec{F} \quad \text{Stokes' theorem}$$

$$= \int_{F_1} \vec{B} \cdot d\vec{F} = 0$$

$$\Rightarrow \int_{C_a} \vec{A} \cdot d\vec{x} = \int_{C_b} \vec{A} \cdot d\vec{x} =: \alpha_1$$

Paths of type (1) have the same action independent of the path

②: analogously:  $\int_{C_c} \vec{A} \cdot d\vec{x} = \int_{C_d} \vec{A} \cdot d\vec{x} =: \alpha_2$

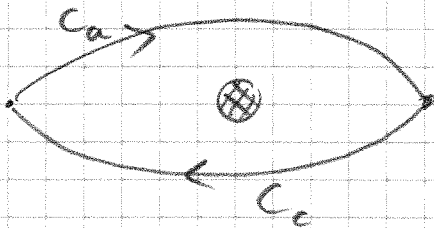
In the path integral:

$\langle \vec{x}, t_1 | \vec{y}, t_0 \rangle = \int \mathcal{D}x e^{\frac{i}{\hbar} S} = \int_{\text{paths of type (1)}} \mathcal{D}x e^{\frac{i}{\hbar} S} + \int_{\text{type (2)}} \mathcal{D}x e^{\frac{i}{\hbar} S}$   
 $= \int_{(1)} \mathcal{D}x e^{\frac{i}{\hbar} S_0} e^{\frac{ie}{\hbar} \alpha_2} + \int_{(2)} \mathcal{D}x e^{\frac{i}{\hbar} S_0} e^{\frac{ie}{\hbar} \alpha_2}$   
 $= K_1 + K_2$

$\Rightarrow \langle \vec{x}, t_1 | \vec{y}, t_0 \rangle = K_1 e^{\frac{ie}{\hbar} \alpha_2} + K_2 e^{\frac{ie}{\hbar} \alpha_2} = e^{\frac{ie}{\hbar} \alpha_2} (K_1 + K_2 e^{\frac{ie}{\hbar} (\alpha_2 - \alpha_1)})$

The interference pattern is determined by  $|\langle \rangle|^2$  and depends on

$\alpha_2 - \alpha_1 = \int_{C_a} \vec{A} \cdot d\vec{x} - \int_{C_c} \vec{A} \cdot d\vec{x} = \oint_{C_a - C_c} \vec{A} \cdot d\vec{x} = \int \vec{B} \cdot d\vec{S} = \Phi$   
magnetic flux

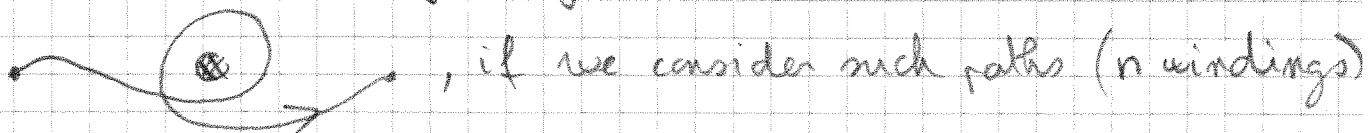


$|\langle \vec{x}, t_1 | \vec{y}, t_0 \rangle|^2 = |K_1 + K_2 e^{\frac{ie}{\hbar} \Phi}|^2$

Remarks:

\*  $\vec{A}$  depends on the gauge but  $\oint \vec{A} \cdot d\vec{x}$  is gauge invariant

\* The decomposition in  $\int^{(1)} + \int^{(2)}$  was incomplete, for example



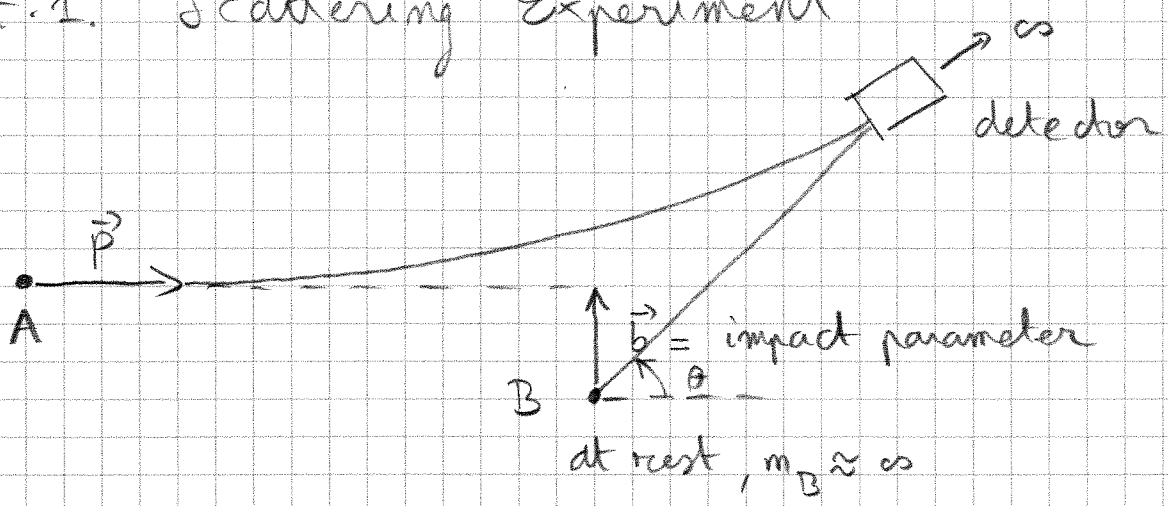
$|\langle \vec{x}, t_1 | \vec{y}, t_0 \rangle|^2 = \left| \sum_{n \in \mathbb{Z}} K_n e^{\frac{ine}{\hbar} \Phi} \right|^2$

# 7. Scattering

So far: spectroscopy:  
transitions between discrete states

Scattering: continuum of states (at beginning and end)  
- form superposition of states

## 7.1. Scattering Experiment

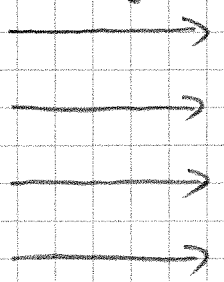


$t \rightarrow \pm \infty$ : interaction between A and B can be neglected  
 $\rightarrow$  uniform (straight line) motion

$V(\vec{x}_A - \vec{x}_B) \cong \text{constant}$  for  $|\vec{x}_A - \vec{x}_0| \rightarrow \infty$   
detector  $\rightarrow$  infinitely distant

more realistic:

$\vec{p}$   
incoming beam



thin target

impact parameter  $\vec{b}$   
and position of B  
in the plane  $\perp \vec{p}$   
are statistically  
distributed (even  
classically)

The detector counts the incoherent superposition of scatterings of particles A's in the beam at particles B's in the target without multiple scatterings

QM: wave packets,  $\vec{p}$ ,  $\vec{b}$  are not sharp, target density, ...

Choose dimensions of the system such that

- classical treatment is asymptotically meaningful
- uncertainty relation OK, i.e.  $\vec{p}$  and  $\vec{x}$  are sufficiently well determined

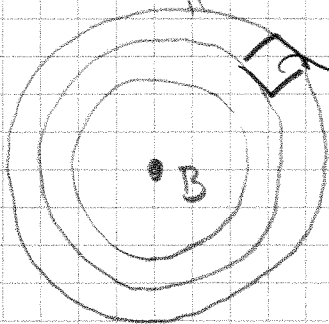
Counting rate  $\left[ \frac{\text{clicks}}{\text{sec}} \right]$  in the detector

$$R(d\Omega) \propto \gamma_{\text{in}} \cdot \# (\text{involved scattering centers})$$

$$\gamma_{\text{in}} = \frac{\# \text{ incoming particles}}{\text{sec} \cdot \text{area} (\perp \vec{p})} \left[ \frac{1}{\text{sec} \cdot \text{m}^2} \right]$$

$$d\sigma := \frac{R}{\gamma_{\text{in}} \cdot \# (\text{scattering centers})} = f(\theta, \varphi) d\Omega \quad [\text{m}^2]$$

differential cross section



$d\sigma$ : particles incoming here land asymptotically in  $d\Omega$  (classically)

- classical scattering is described statistically because of beam, scattering centers (impact parameter cannot be determined)
- same for the QM description ("rate of clicks in  $d\Omega$ "), although there are no trajectories any more

$$\sigma_{\text{tot}} = \int d\sigma = \int d(\cos\theta) d\varphi F(\theta, \varphi)$$

Total cross section

$\sigma_{\text{tot}}$ : surface around B, where particles are deflected = experience a force

$\sigma_{\text{tot}} \sim (\text{range of interaction})^2$  if finite  
otherwise: more complicated

Nuclear physics: range  $1 \text{ fm} = 10^{-13} \text{ cm}$

Unit for cross section:  $1 \text{ barn} = 10^{-24} \text{ cm}^2$

mb,  $\mu\text{b}$ , pb, fb =  $10^{-12}$  barn

Here we consider only elastic scattering: the particle changes its direction but not its kinetic energy

We do not consider inelastic scattering  $A + B \rightarrow A^* + B^*$   
different internal state, e.g. atom

Goal: infer information about the potential  $V$  from  $\frac{d\sigma(\theta, \varphi)}{d\Omega} = F$

Remark:

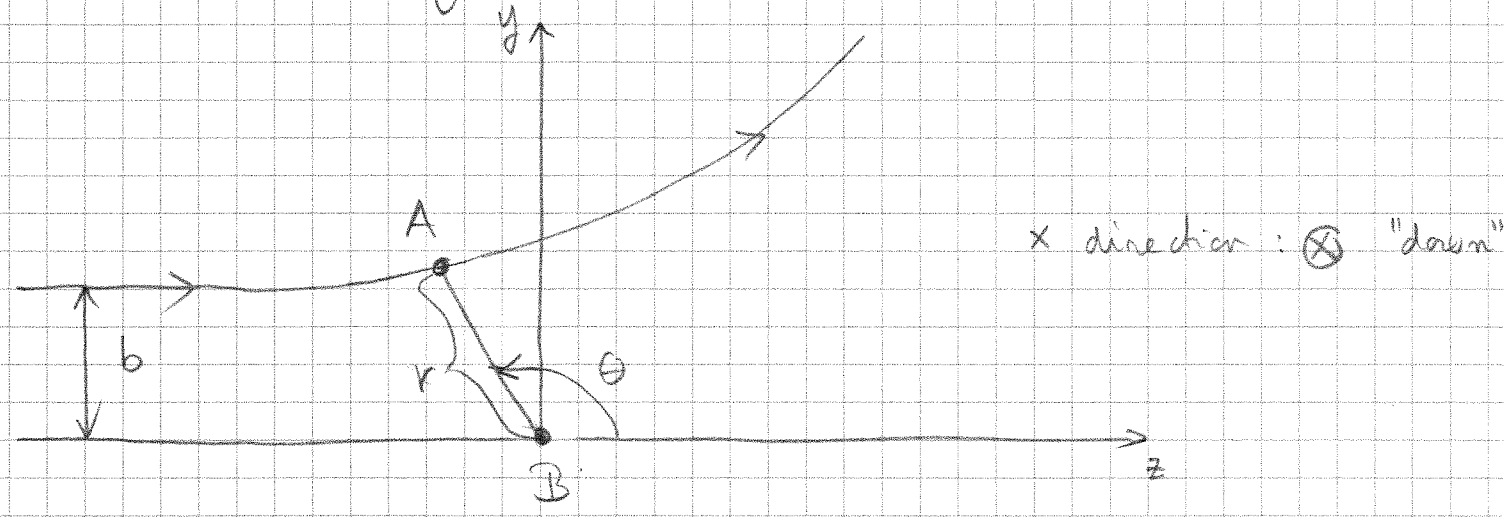
new particles (relativistic,  $E > mc^2$ ):  $A + B \rightarrow X$

$\sigma_X$  = "cross section for X-production"

$\frac{\sigma_X}{\sigma_{\text{anything}}}$  = X-branching ratio

Constraints: conservation laws (can be tested)

# 7.2 Scattering classical



effective "1-body in potential" problem

$$m \ddot{\vec{r}} = -\vec{\nabla} V(|\vec{r}|) ; \quad V = V(|\vec{r}|) \text{ central potential, well-behaved}$$

$$\vec{r}(t) \approx \vec{v}_{in} \cdot t + \vec{r}_{in} \quad \text{for } t \rightarrow -\infty$$

$$\vec{r}(t) \approx \vec{v}_{out} \cdot t + \vec{r}_{out} \quad \text{for } t \rightarrow +\infty$$

$$\vec{L} = \vec{r} \times \vec{p} = \vec{r}_{in} \times m\vec{v}_{in} = \vec{r}_{out} \times m\vec{v}_{out}$$

$\vec{L}$  is conserved  $\Rightarrow$  motion in the  $(y, z)$ -plane:

$$\vec{r} = (0, r \sin \theta, r \cos \theta)$$

$$\vec{v}_{in} = (0, 0, v)$$

$$\Rightarrow \left. \begin{aligned} r(t) &\approx -v \cdot t \\ \pi - \theta(t) &\approx \frac{b}{-vt} \end{aligned} \right\} t \rightarrow -\infty$$

Energy:

$$E = \frac{m}{2} \underbrace{(\dot{r}^2 + r^2 \dot{\theta}^2)}_{= \dot{\vec{r}}^2} + V(r) \underset{t \rightarrow -\infty}{=} \frac{m}{2} v^2$$

Angular momentum:

$$\vec{L} = \vec{r} \times \vec{p} =: \begin{pmatrix} l \\ 0 \\ 0 \end{pmatrix}, \quad l = -m r^2 \dot{\theta} \underset{t \rightarrow -\infty}{=} -m (-vt)^2 \left( \frac{-b}{vt^2} \right) = +mbv = b \sqrt{2mE}$$

⇒

$$\dot{r}^2 = \frac{2}{m} \left[ E \left( 1 - \frac{b^2}{r^2} \right) - V(r) \right]$$

↑ centrifugal term  $\frac{l^2}{2mr^2}$

for  $t \leq t_0$ :

$$\dot{r} = - \sqrt{\frac{2}{m} \left[ E \left( 1 - \frac{b^2}{r^2} \right) - V(r) \right]^{1/2}}$$

$r$  decreases till the right hand side becomes zero for  $t=t_0, r=r_0$   
afterwards

$$\dot{r} = + \sqrt{\frac{2}{m} \left[ E \left( 1 - \frac{b^2}{r^2} \right) - V(r) \right]^{1/2}} \quad \text{for } t \geq t_0$$

$$r(t_0 + \Delta t) = r(t_0 - \Delta t) \quad (\text{symmetry})$$

$$r_0 = r(t_0) : \text{turning point for } r$$

$$t \rightarrow \infty : r(t) = r(t_0 + t - t_0) = r(t_0 - (t - t_0)) = r(\underbrace{2t_0 - t}_{\rightarrow -\infty}) = v \cdot (t - 2t_0)$$

$$\text{Conservation of angular momentum} \Rightarrow \dot{\Theta} = - \frac{bv}{r(t)^2}$$

$$\Theta_{\text{out}} - \Theta_{\text{in}} = \Theta_{\text{out}} - \pi$$

$$= \int_{-\infty}^{\infty} dt \dot{\Theta}(t) = -bv \int_{-\infty}^{\infty} \frac{dt}{r(t)^2} = -2bv \int_{t_0}^{\infty} \frac{dt}{r(t)^2} \quad \left[ \frac{dr}{dt} = \dot{r} \right]$$

$$= -2bv \int_{r_0}^{\infty} \frac{dr}{\dot{r} r^2} = -2b\sqrt{E} \int_{r_0}^{\infty} \frac{dr}{r} \left[ E(r^2 - b^2) - r^2 V(r) \right]^{-1/2}$$

$\Theta_{\text{out}}$  = scattering angle

given  $b, E \rightarrow$  find  $r_0$  ( $\dot{r}=0$ )  $\rightarrow$  compute  $\Theta_{\text{out}}$

$$\Theta_{\text{out}} =: \Theta = f(b) \quad \text{for fixed } E$$

Cross section:

$$J_{\text{in}} = 2\pi b db \frac{\text{particles}}{\text{sec}} \quad \text{incoming in the ring}$$



⇒ These particles are deflected in the interval  $[\theta(b), \theta(b+db)]$

# particles in  $d\Omega$  =  $\gamma_{in} b db d\varphi$  ( $2\pi = \int d\varphi$ )

=  $\gamma_{in} b(\theta) \frac{db}{d\theta} \frac{d\Omega}{\sin\theta}$  ( $d\Omega = \sin\theta d\theta d\varphi$ )

⇒  $\frac{d\sigma}{d\Omega} = \frac{b(\theta)}{\sin\theta} \frac{db}{d\theta}$

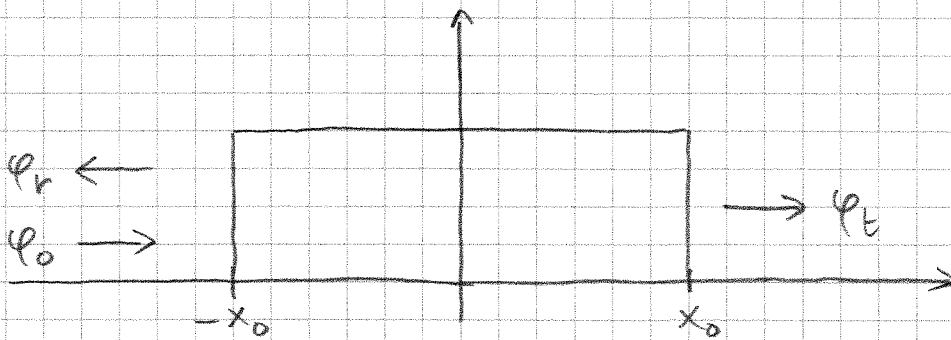
$\sigma_{tot} = \int \frac{d\sigma}{d\Omega} d\Omega = 2\pi \int_{\theta=0}^{\theta=\pi} b db = \pi b_{max}^2 = \text{surface,}$   
 where scattering occurs

$\sigma_{tot} < \infty$  only when  $V(r) = 0$  for  $r \geq R$

### 7.3 Scattering QM

#### 7.3.1 Asymptotic Solution for rotationally invariant potentials

Reminder: dimension  $d=1$



$\left. \begin{array}{l} \varphi_0 : \text{incoming wave} \\ \varphi_r : \text{reflected wave} \end{array} \right\} \text{for } x_0 < -x_0 : \varphi_0 + \varphi_r$   
 $\varphi_t : \text{transmitted wave} \quad \text{for } x_0 > x_0 : \varphi_t$

stationary scattering solution:

$\varphi_0 + \varphi_r = e^{ikx} + \alpha_-(k) e^{-ikx}$

$\varphi_t = \gamma_+(k) e^{ikx}$

$E = \frac{(\hbar k)^2}{2m}$



general scattering solution for  $x \notin [-x_0, x_0]$ :

$$\Psi(x,t) = \int dk \tilde{\varphi}(k) e^{-\frac{i}{\hbar} E(k)t} \begin{cases} \varphi_0 + \varphi_r, & x < -x_0 \\ \varphi_t & x > x_0 \end{cases}$$

- time dependent solution of the Schrödinger equation:

$$i\hbar \frac{\partial \Psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \Psi}{\partial x^2}$$

- no fixed energy, momentum
- localized, for example incoming wave packet

In  $d=3$  dimensions we look for stationary scattering solutions

$$\varphi(\vec{x}) = \varphi_0(\vec{x}) + \varphi_s(\vec{x})$$

$$\left( -\frac{\hbar^2}{2m} \Delta + V(|\vec{x}|) - E \right) \varphi(\vec{x}) = 0$$

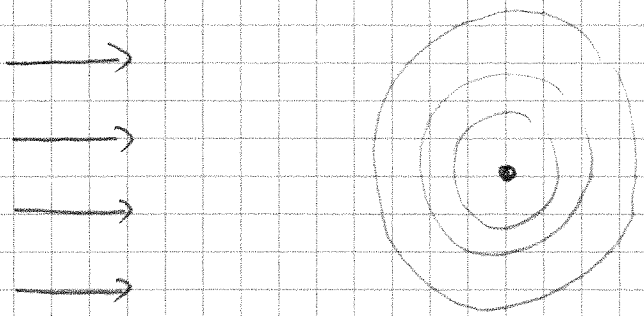
with  $\varphi_0 = e^{i\vec{k} \cdot \vec{x}}$  : incoming

Particle current:

$$\vec{j}_0 = \frac{\hbar}{2mi} \left( \varphi_0^* \vec{\nabla} \varphi_0 - \varphi_0 \vec{\nabla} \varphi_0^* \right) = \frac{\hbar \vec{k}}{m}$$

in direction of  $\vec{k} = (0, 0, k)$

to be determined:  $\varphi_s$  : scattering wave ( $\leftrightarrow V(r)$ )



$V(r) \xrightarrow{r \rightarrow \infty} 0$  sufficiently fast, such that:

$$\lim_{r \rightarrow \infty} r V(r) = 0, \text{ i.e. } V \propto \frac{1}{r^{1+\epsilon}}, \text{ not Coulomb}$$

$\Rightarrow \varphi_0 + \varphi_s$  is asymptotically ( $r \rightarrow \infty$ ) a solution of the free Schrödinger equation with suitable boundary conditions:  $\varphi_s$  outgoing scattering wave

$\varphi_0 = e^{ikz}$  = solution of the free ( $V=0$ ) Schrödinger equation

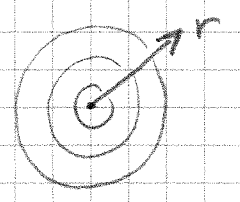
$\varphi_0 = e^{ikr \cos(\theta)}$  =  $F(r, \theta, \varphi)$  =  $\sum_{l=0}^{\infty} i^l (2l+1) j_l(kr) P_l(\cos \theta)$   
polar coordinates      symmetry

$j_l(z) = (-z)^l \left(\frac{1}{z} \frac{d}{dz}\right)^l \frac{\sin z}{z}$  : spherical Bessel function

asymptotically:

$j_l(kr) \xrightarrow{r \rightarrow \infty} \frac{1}{kr} \sin\left(kr - \frac{l\pi}{2}\right) = \frac{1}{2ikr} \left[ e^{i\left(kr - \frac{l\pi}{2}\right)} - e^{-i\left(kr - \frac{l\pi}{2}\right)} \right]$   
outgoing spherical wave      incoming spherical wave

$e^{-\frac{i}{\hbar}Et} e^{ikr}$  has constant phase for  $r = \frac{E \cdot t}{\hbar k}$



Ansatz:  $\varphi_s = f(\theta) \frac{e^{ikr}}{r}$   
= modification of the outgoing component

It is indeed an asymptotic solution for  $r \rightarrow \infty$ :

$\Delta \varphi_0 = -k^2 \varphi_0$

$\Delta \varphi_s = \left( \frac{1}{r} \frac{\partial^2}{\partial r^2} r - \frac{\vec{L}^2}{r^2 \hbar^2} \right) f(\theta) \frac{e^{ikr}}{r}$   
 $= \left( -k^2 + O\left(\frac{1}{r^2}\right) \right) \varphi_s$

$\Rightarrow \left( -\frac{\hbar^2}{2m} \Delta + V(r) \right) (\varphi_0 + \varphi_s) = \left( \frac{\hbar^2 k^2}{2m} + O\left(\frac{1}{r^2 + E}\right) \right) \varphi \stackrel{!}{=} E \varphi$   
assumption!

$\Rightarrow E = \frac{(\hbar k)^2}{2m}$  = energy of the incoming particles

Particle current of the scattering wave:

$$\vec{j}_s = \frac{\hbar}{2im} (\psi_s^* \vec{\nabla} \psi_s - \psi_s \vec{\nabla} \psi_s^*)$$

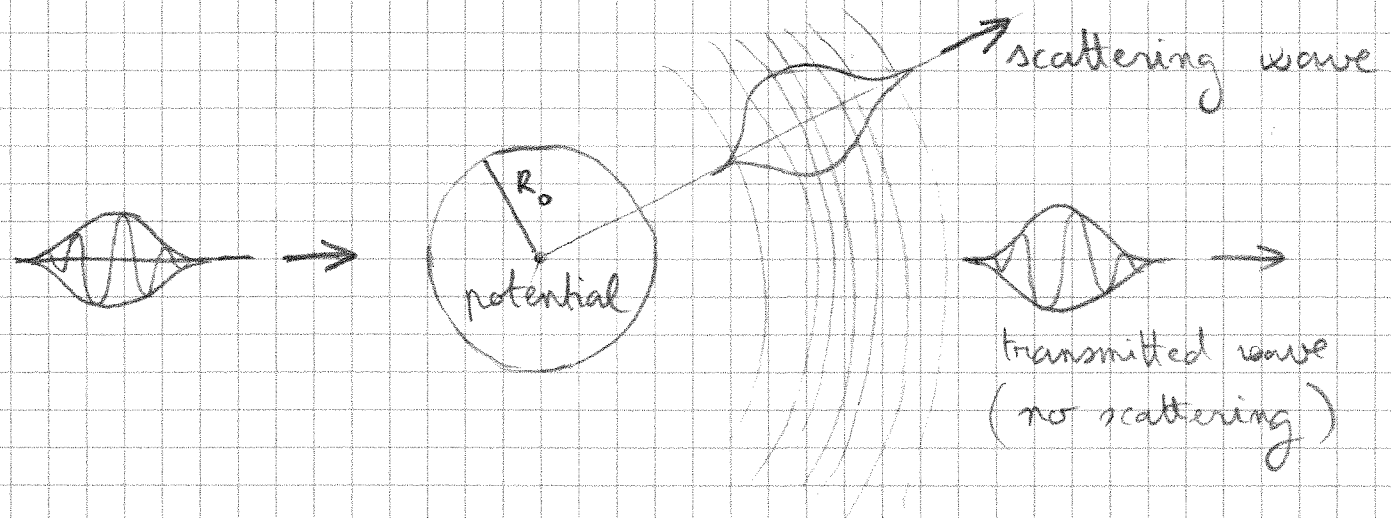
$$= \frac{\hbar k}{m} \frac{|f(\theta)|^2}{r^2} \vec{e}_r + \mathcal{O}\left(\frac{1}{r^3}\right), \quad \vec{r} = r \vec{e}_r$$

because  $\vec{\nabla} = \vec{e}_r \frac{\partial}{\partial r} + \frac{1}{r} \vec{e}_\theta \frac{\partial}{\partial \theta} + \frac{1}{r \sin \theta} \vec{e}_\varphi \frac{\partial}{\partial \varphi}$   
 $\hookrightarrow ik$

- We assume that there is no interference between  $\psi_0$  and  $\psi_s$ ,

$$\vec{j}_{0+s} = \vec{j}_0 + \vec{j}_s$$

This can be achieved through the construction of wavepackets



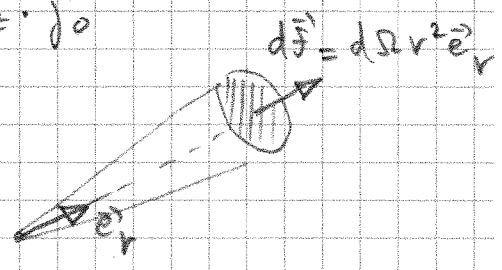
$$\frac{\# \text{ particle in detector}}{\text{sec}} = \int_{\text{sphere } (R)} \vec{j}_s \cdot d\vec{F} = \frac{2\pi \hbar k}{m} \underbrace{\int_{-1}^1 d(\cos \theta) |f(\theta)|^2}_{\text{independent of } R} + \mathcal{O}\left(\frac{1}{R}\right)$$

$\Rightarrow$  conservation of number of particles of the asymptotic solution

$$d\sigma(\theta, \varphi) = \lim_{r \rightarrow \infty} \frac{\vec{j}_s \cdot d\vec{F}}{\vec{e}_z \cdot \vec{j}_0} = \lim_{r \rightarrow \infty} \frac{\vec{j}_s \cdot (d\Omega r^2 \vec{e}_r)}{\vec{e}_z \cdot \vec{j}_0}$$

$$= |f(\theta)|^2 d\Omega$$

$$\frac{d\sigma}{d\Omega} = |f(\theta)|^2$$



- look for solutions with asymptotics  $\psi_s$

### 7.3.2 Partial Waves

From asymptotics of the solution in full space we determine  $f(\theta)$

$V(|\vec{x}|) \rightarrow$  separation in polar coordinates  $\rightarrow$  effective  $d=1$  problem

General Ansatz: 
$$\varphi(r, \theta) = \sum_{l=0}^{\infty} \frac{u_l(r)}{r} P_l(\cos \theta)$$

- only  $P_l \propto Y_{l0}$  because of symmetry in the azimuthal angle  $\phi$  ( $Y_{lm} \propto e^{im\phi}$ )

Stationary Schrödinger equation:

$$\left( -\frac{\hbar^2}{2m} \Delta + V(r) - E \right) \varphi = 0$$

Ansatz is solution, if for each  $l$  it holds

$$u_l'' + (k^2 - v_{\text{eff}}(r)) u_l = 0$$

$$k^2 = \frac{2m}{\hbar^2} E$$

$$v_{\text{eff}} = \frac{2m}{\hbar^2} \left[ V(r) + \frac{\hbar^2 l(l+1)}{2mr^2} \right] \quad \left( \Delta = \frac{1}{r} \frac{\partial^2}{\partial r^2} r - \frac{L^2}{r^2 \hbar^2} \right)$$

$\varphi(0, \theta)$  regular  $\Rightarrow u_l(0) = 0$  boundary condition (b.c.)

- unique solution up to a complex factor of the differential equation of 2nd order and b.c. at  $r=0$

- for each  $l$  one free constant (= factor in  $u_l$ ) for the general solution

$\rightarrow$  compute  $u_l$  for a given  $V(r)$

→ fit to asymptotics  $\varphi_0 + \varphi_s = e^{ikz} + f(\theta) \frac{e^{ikr}}{r}$  for  $r \rightarrow \infty$

→ determine  $f(\theta)$

For  $V=0$  the solution is  $\varphi = \varphi_0 \forall r$  with  $\frac{u_l}{r} \propto j_l(kr)$

⇒ solves  $u_l'' + \left(k^2 - \frac{l(l+1)}{r^2}\right) u_l = 0$  with  $u_l(0) = 0$

$l=0$  :  $u_0 = kr j_0(kr) = \sin(kr)$

$l > 0$  :  $u_l = kr j_l(kr) \underset{r \rightarrow \infty}{\approx} \sin\left(kr - \frac{l\pi}{2}\right)$

$-\frac{l\pi}{2}$  : phase shift due to centrifugal potential

For  $V \neq 0$  but  $V \xrightarrow{r \rightarrow \infty} 0$  fast enough it must hold

$u_l \underset{r \rightarrow \infty}{\approx} \alpha_l \sin\left(kr - \frac{l\pi}{2} + \delta_l\right)$

$\delta_l$  : phase shift due to  $V$

Determine  $\{\alpha_l\}$  through asymptotics  $\varphi_0 + \varphi_s$  :

$\alpha_l$  can be chosen such that

$u_l \underset{r \rightarrow \infty}{\approx} \frac{1}{k} i^l (2l+1) e^{i\delta_l} \sin\left(kr - \frac{l\pi}{2} + \delta_l\right)$

⇒  $\varphi(r, \theta) \underset{r \rightarrow \infty}{\approx} \sum_{l=0}^{\infty} i^l (2l+1) P_l(\cos\theta) \frac{1}{2ikr} \left[ e^{2i\delta_l} e^{i\left(kr - \frac{l\pi}{2}\right)} - e^{-i\left(kr - \frac{l\pi}{2}\right)} \right]$

- in the representation of  $e^{ikz}$  in terms of spherical waves only the outgoing wave changes by a factor  $S_l = e^{2i\delta_l} = \text{phase}$

⇒  $f(\theta) = \frac{1}{k} \sum_{l=0}^{\infty} (2l+1) \frac{1}{2i} (S_l - 1) P_l(\cos\theta)$

⇒  $S_l$  contain precisely the information about the effect of the potential

$\delta_e$  = phase shift of the  $l$  partial wave (8)

$|S_e| = 1$  (phase)  $\Rightarrow$  scattering is elastic: it holds

$$\int_{\partial V} \vec{j} \cdot d\vec{F} = 0 \quad : \quad \text{no particle of the type considered is lost}$$

$\partial V$  = sphere of radius  $R \rightarrow \infty$

Proof:

$$\vec{j} = \frac{\hbar}{m} \nabla \varphi^* (\varphi \nabla \varphi) \quad ; \quad \varphi = \sum_{l=0}^{\infty} \frac{u_l(r)}{r} P_l(\cos \theta)$$

$$\nabla \varphi = \frac{\partial \varphi}{\partial r} \vec{e}_r + \dots \quad u_l \underset{r \rightarrow \infty}{\approx} \frac{i^l (2l+1)}{2ik} \left[ S_l e^{i(kr - \frac{l\pi}{2})} - e^{-i(kr - \frac{l\pi}{2})} \right]$$

$$\int_{\partial V} \vec{j} \cdot d\vec{F} = \int d\Omega r^2 \vec{e}_r \cdot \vec{j} =$$

$$= \sum_{e, e'} \underbrace{\int d\Omega P_e(\cos \theta) P_{e'}(\cos \theta)}_{= \delta_{ee'} \frac{4\pi}{2e+1}} \frac{\hbar}{m} \nabla \varphi^* \frac{d}{dr} \varphi$$

$$= \sum_e 4\pi (2e+1) \frac{\hbar}{m} \frac{1}{4ik} \left( \underset{\substack{\uparrow \\ \text{outgoing}}}{|S_e|^2} - \underset{\substack{\uparrow \\ \text{incoming}}}{1} \right) = 0$$

$\Rightarrow$  conservation of particle current for each partial wave (9)

$$f(\theta) = \frac{1}{k} \sum_{l=0}^{\infty} (2l+1) e^{i\delta_l} \sin(\delta_l) P_l(\cos \theta)$$

classically:  $L = bp \leq R_0 \sqrt{2mE}$  ( $R_0$  = range of  $V$ )

$\rightarrow$  does this imply only a finite number of  $l$ -values in GM?

### 7.3.3 The Optical Theorem

$$\frac{d\sigma}{d\Omega} = |F(\theta)|^2$$

$$\sigma = \int d\Omega \frac{d\sigma}{d\Omega} = \frac{1}{k^2} \sum_{\ell, \ell'} (2\ell+1)(2\ell'+1) e^{i(\delta_\ell - \delta_{\ell'})} \sin(\delta_\ell) \sin(\delta_{\ell'}) \int d\Omega P_\ell(\cos\theta) P_{\ell'}(\cos\theta)$$

$$= \frac{4\pi}{2\ell+1} \delta_{\ell\ell'}$$

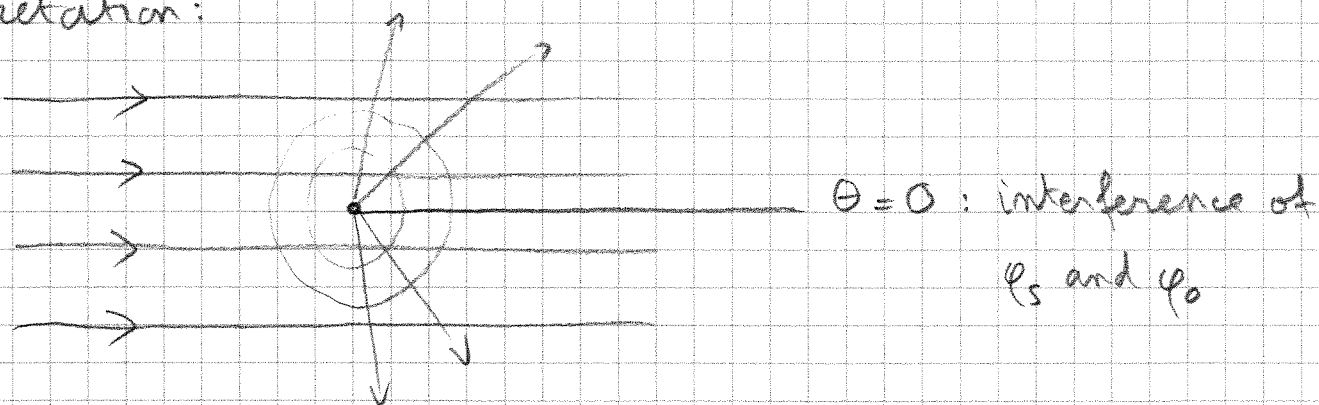
$$\Rightarrow \sigma = \frac{4\pi}{k^2} \sum_{\ell} (2\ell+1) \sin^2(\delta_\ell)$$

$$F(0) = \frac{1}{k} \sum_{\ell} (2\ell+1) e^{i\delta_\ell} \sin \delta_\ell \quad (P_\ell(1) = 1)$$

$$\Rightarrow \sigma = \frac{4\pi}{k} \text{Im} F(0)$$

This relation holds in general, even without rotational invariance.

Interpretation:



$F(\theta) \rightarrow$  scattering current deriving from  $\varphi_0$  in direction  $\theta=0$  must be cancelled by interference ("shadow") because of conservation of the total current

## 7.4 Example: The Potential Well

$$V(r) = \begin{cases} -V_0 & \text{for } r < a \\ 0 & \text{for } r > a \end{cases}, \quad q(r)^2 = \begin{cases} k_0^2 = \frac{2m}{\hbar^2} (E + V_0) & \text{for } r < a \\ k^2 = \frac{2m}{\hbar^2} E & \text{for } r > a \end{cases}$$

$$u_e'' + \left( q^2 - \frac{l(l+1)}{r^2} \right) u_e = 0, \quad \varphi = \sum_{l=0}^{\infty} \frac{u_l(r)}{r} P_l(\cos \theta)$$

Solution with  $u_e(0) = 0$ :

$$\frac{u_e(r)}{r} = \begin{cases} c_e j_l(kr) & \text{for } r < a \\ \alpha_e j_l(kr) + \beta_e n_l(kr) & \text{for } r > a \end{cases}$$

$\uparrow$   $\uparrow$  Neumann functions (irregular at  $r=0$ )  
 basis for general solution ( $\rightarrow \sin, \cos$  asymptotically)

$c_e$  are arbitrary (Normalization), then  $\alpha_e, \beta_e$  fixed by continuity at  $r=a$

For  $r \rightarrow \infty$ :

$$u_e \stackrel{r \rightarrow \infty}{\approx} \frac{\alpha_e}{k} \sin\left(kr - \frac{l\pi}{2}\right) - \frac{\beta_e}{k} \cos\left(kr - \frac{l\pi}{2}\right)$$

$$\propto \sin\left(kr - \frac{l\pi}{2} + \delta_l\right) = \cos \delta_l \sin\left(kr - \frac{l\pi}{2}\right) + \sin \delta_l \cos\left(kr - \frac{l\pi}{2}\right)$$

$$\Rightarrow \tan \delta_l = -\frac{\beta_l}{\alpha_l}$$

$\frac{d}{dr} \ln\left(\frac{u_l}{r}\right)$  continuous at  $r=a$  ( $\Leftrightarrow \frac{u_l}{r}$  and  $\frac{d}{dr}\left(\frac{u_l}{r}\right)$  continuous):

$$k_0 \left. \frac{j_l'}{j_l} \right|_{r=k_0 a} = k \frac{\alpha_l j_l' + \beta_l n_l'}{\alpha_l j_l + \beta_l n_l} \Big|_{r=ka}$$

$$\Rightarrow \tan \delta_l = \frac{k j_l'(ka) j_l(k_0 a) - k_0 j_l'(k_0 a) j_l(ka)}{k n_l'(ka) j_l(k_0 a) - k_0 j_l'(k_0 a) n_l(ka)}$$



## 7.4.1 Low-Energy Limit

$$z = k \cdot a \rightarrow 0 \quad (\text{not } k_0 \cdot a \rightarrow 0!)$$

$$\text{for small argument } z : j_l(z) = \frac{z^l}{(2l+1)!!} (1 + O(z^2))$$

$$n_l(z) = -\frac{(2l-1)!!}{z^{l+1}} (1 + O(z^2))$$

$$(2l \pm 1)!! = 1 \cdot 3 \cdot 5 \cdot \dots \cdot (2l \pm 1)$$

Therefore:

$$\begin{aligned} \tan \delta_l &= \frac{1}{(2l+1)!! (2l-1)!!} \frac{l z^l j_l(k_0 a) - z^l a k_0 j_l'(k_0 a)}{(l+1) z^{-l-1} j_l(k_0 a) + z^{-l-1} a k_0 j_l'(k_0 a)} \\ &= \frac{z^{2l+1}}{(2l+1)!! (2l-1)!!} \frac{l j_l - a k_0 j_l'}{(l+1) j_l + a k_0 j_l'} \quad \text{all arguments} = k_0 a \end{aligned}$$

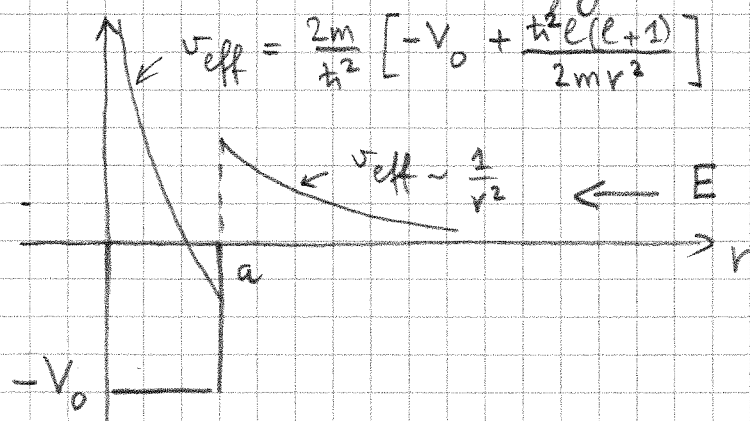
If the denominator  $\neq 0$  (see resonance):

$$\tan \delta_l \propto (ka)^{2l+1}$$

dominated by  $l=0$ , s-wave scattering

other  $l$  values: centrifugal repulsion, particles do not

"see"  $V(r)$  at low energy:



$$\tan \delta_0 = -ka \frac{a k_0 j_0'(k_0 a)}{j_0(k_0 a) + a k_0 j_0'(k_0 a)}$$

with  $j_0(z) = \frac{\sin(z)}{z}$  :  $\tan \delta_0 = ka \left( \frac{\tan(ka)}{ka} - 1 \right)$

### 7.4.2 Resonance

$$0 = (l+1) j_l(k_0 a) + k_0 a j_l'(k_0 a) \Big|_{E=E_R}, \quad k_0^2 = \frac{2m}{\hbar^2} (E + V_0)$$

In the case  $y = k_0 a \gg l \gg 1$  (deep well):

$$j_l(y) \approx \frac{1}{y} \sin\left(y - \frac{l\pi}{2}\right)$$

$$j_l'(y) \approx -\frac{1}{y^2} \sin\left(y - \frac{l\pi}{2}\right) + \frac{1}{y} \cos\left(y - \frac{l\pi}{2}\right)$$

$$0 = \frac{l}{k_0 a} \sin\left(k_0 a - \frac{l\pi}{2}\right) + \cos\left(k_0 a - \frac{l\pi}{2}\right)$$

$$-\cot\left(k_0 a - \frac{l\pi}{2}\right) = \frac{l}{k_0 a} \ll 1$$

$$= \tan\left(\underbrace{k_0 a - \frac{(l+1)\pi}{2}}_{\text{close to } n \cdot \pi}\right)$$

$$\Rightarrow \underbrace{k_0 a - \frac{l\pi}{2} \approx \left(n + \frac{1}{2}\right)\pi + \frac{l}{k_0 a}}_{\substack{\text{condition for binding energy} \\ \text{in the deep 3d potential well}}}$$

$n = 0, 1, 2, \dots$   
 since  $\frac{l}{k_0 a} \ll 1$ ,  
 $n < 0$  is not possible

Scattering phase at  $E_R$ :

for  $E \neq E_R$ ,  $\tan \delta_l \approx (ka)^{2l+1} \ll 1$ ,  $\delta_l \approx m \cdot \pi$

at  $E_R$ :  $\tan \delta_l$  diverges,  $\delta_l(E_R) = \left(m \pm \frac{1}{2}\right)\pi$ ,  
 $\delta_l$  makes a jump  $\delta_l \rightarrow (m \pm 1) \cdot \pi$

Taylor expansion of denominator at  $E \approx E_R$ , replace  $E \rightarrow E_R$  in the numerator  $\Rightarrow$

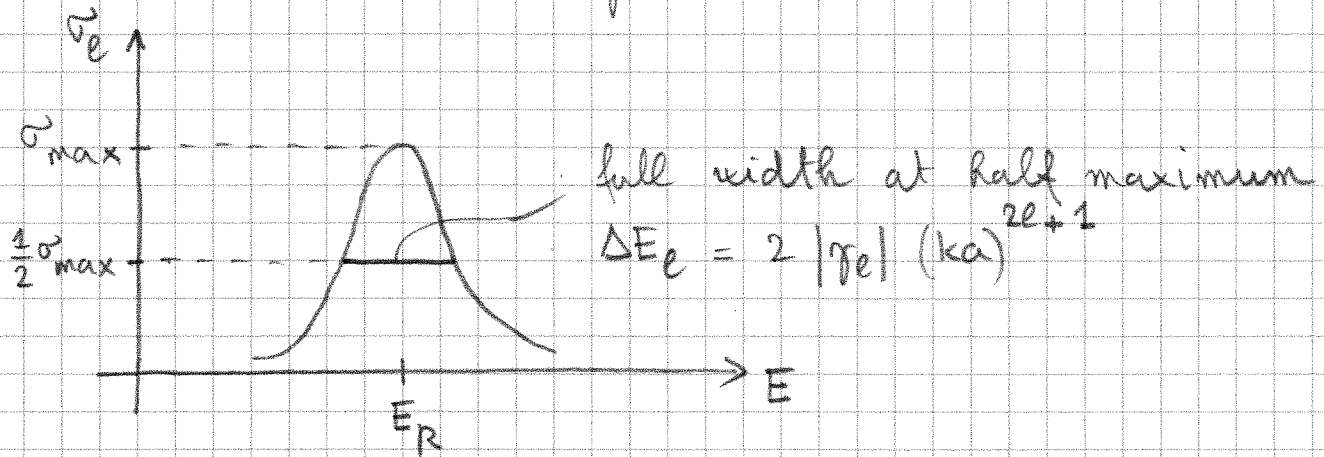
$$\tan \delta_l \approx \underbrace{r_l}_{\text{known explicitly}} \frac{(ka)^{2l+1}}{\underbrace{E - E_R}_{\text{linear term}}}$$

Partial cross section:

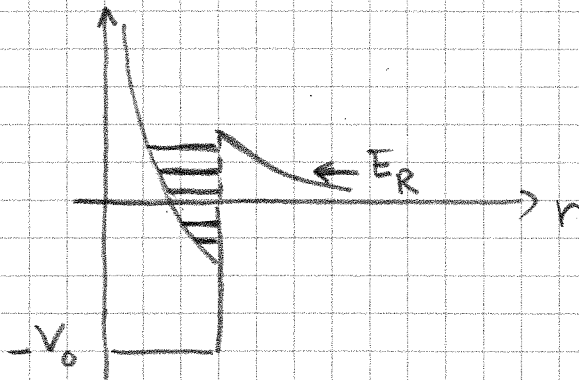
$$\sigma_e = \frac{4\pi(2l+1)}{k^2} \sin^2 \delta_e, \quad \sin^2 = \frac{\tan^2}{1 + \tan^2}$$

$$\sigma_e = \frac{4\pi(2l+1)}{k^2} \frac{\gamma_e^2 (ka)^{4l+2}}{(E - E_R)^2 + \gamma_e^2 (ka)^{4l+2}}$$

Breit-Wigner formula



- sharp resonance,  $ka$  is small
- other  $\sigma_e$  negligible (do not fulfill resonance condition)
- large part of the  $l$ -wave tunnels through centrifugal barrier
- quasi-bound state, metastable, lifetime  $\tau$ :  $\Delta E_e \sim \frac{\hbar}{\tau}$ ,  $l$ -wave emitted again  $\rightarrow$  large  $\sigma_e$



$l=0$ , s-wave scattering

Low-energy limit:  $\tan \delta_0 = -ka_s$   
 $a_s = a \left( 1 - \frac{\tan(ka)}{ka} \right)$  scattering length

$$\Rightarrow \sigma_0 = \frac{4\pi}{k^2} \sin^2 \delta_0 \xrightarrow{k \rightarrow 0} 4\pi a_s^2 + O(k^2)$$

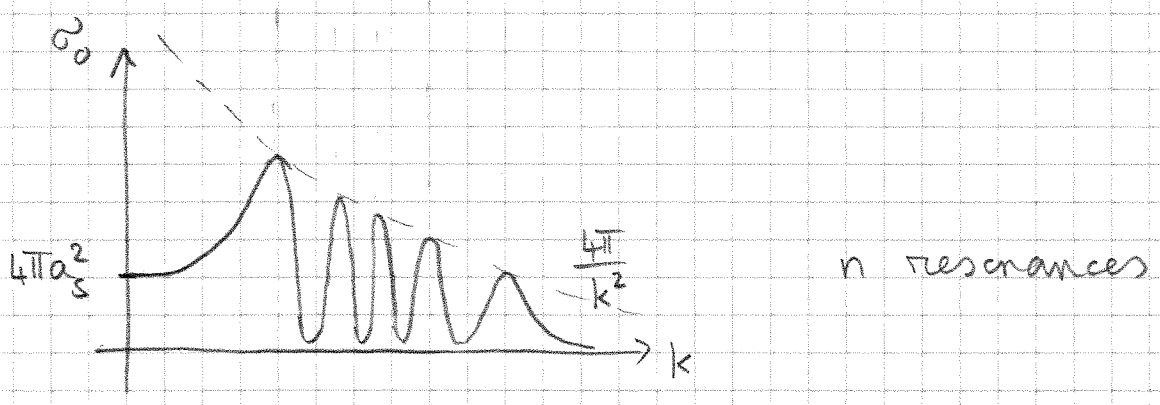
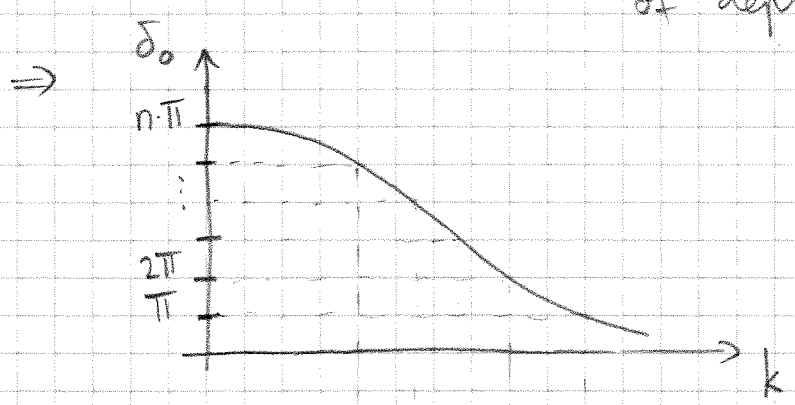
Analysis of the exact expression for  $\tan \delta_0$  yields

$$0 = \lim_{k \rightarrow 0} \tan \delta_0(k) = \lim_{k \rightarrow \infty} \tan \delta_0(k)$$

$\delta_0$  was introduced in  $e^{2i\delta_0} \rightarrow$  undetermined by  $n \cdot \pi$

Convention:  $\delta_0(\infty) = 0$ ,  $\delta_0(k)$  continuous

$\Rightarrow \delta_0(0) = n \cdot \pi$ ,  $n = \#$  bound states in the potential well of depth  $V_0$



### 7.4.4 Born Approximation for Scattering Phases

We had:

$$u_l'' + \left( k^2 - \frac{2m}{\hbar^2} V(r) - \frac{l(l+1)}{r^2} \right) u_l = 0$$

$$u_l(0) = 0$$

$$u_l \xrightarrow{r \rightarrow \infty} \frac{1}{k} i^l (2l+1) e^{i\delta_l} \sin \left( kr - \frac{l\pi}{2} + \delta_l \right) \xrightarrow{\text{abstract}} \delta_l$$

Let  $u_e^{(0)}$  be the above solution for  $V=0$ :

$$u_e^{(0)} = i^l (2l+1) r j_l(kr)$$

Using the differential equations it can be shown

$$\int_0^\infty dr \left( \underset{\uparrow}{u_e''} \underset{\downarrow}{u_e^{(0)}} - \underset{\downarrow}{u_e} \underset{\uparrow}{u_e^{(0)''}} \right) = \frac{2m}{\hbar^2} \int_0^\infty dr V u_e u_e^{(0)}$$

partial integration

$$= \left( u_e' u_e^{(0)} - u_e u_e^{(0)'} \right) \Big|_0^\infty = 0$$

$$= \lim_{r \rightarrow \infty} \left( u_e' u_e^{(0)} - u_e u_e^{(0)'} \right)$$

$$= -k \left( \frac{1}{k} i^l (2l+1) \right)^2 e^{i\delta_l} \sin \delta_l$$

$$\Rightarrow i^l (2l+1) e^{i\delta_l} \sin \delta_l = -\frac{2m}{\hbar^2} \int_0^\infty dr V kr j_l(kr) u_e$$

- exact relation (integral representation)

- not an explicit solution:  $u_e$  on the right-hand side!

Born approximation: replace

$u_e \rightarrow u_e^{(0)}$  on the right-hand side

valid when:

- scattering  $\propto V$  is a small effect

-  $\delta_l(k)$  is small

Not in the resonance region, but for  $k \rightarrow \infty$ . Then:

$$\delta_l \stackrel{k \rightarrow \infty}{\approx} -\frac{2m}{\hbar^2} \frac{1}{k} \int_0^\infty dr V(r) \underbrace{[kr j_l(kr)]^2}_{\approx \sin^2(kr) \text{ when } kr \gg l}$$

# 7.5 Solution via Green Function

## 7.5.1 Schrödinger Equation as an Integral Equation

Stationary Schrödinger equation for scattering solution:

$$(\Delta_x + k^2) \varphi(\vec{x}) = v(\vec{x}) \varphi(\vec{x}) \quad , \quad \Delta_x = \sum_{i=1}^3 \frac{\partial^2}{\partial x_i^2}$$

$$k^2 = \frac{2m E}{\hbar^2} \quad , \quad v(\vec{x}) = \frac{2m}{\hbar^2} V(\vec{x})$$

Green function:

$$(\Delta_x + k^2) G(\vec{x}-\vec{y}) = \delta^{(3)}(\vec{x}-\vec{y})$$

Then, the solution  $\varphi(\vec{x})$  fulfills the integral equation

$$\varphi(\vec{x}) = \varphi_0(\vec{x}) + \int d^3y G(\vec{x}-\vec{y}) v(\vec{y}) \varphi(\vec{y}) \quad (*)$$

where  $(\Delta_x + k^2) \varphi_0(\vec{x}) = 0$

\*  $G(\vec{x}-\vec{y})$  is an integral kernel, distribution like  $\delta^{(3)}(\vec{x}-\vec{y})$

\* Boundary conditions must be imposed:  $\varphi_0 + \varphi_s$  for  $|\vec{x}| \rightarrow \infty$

Born series:

Ansatz:  $\varphi = \sum_{n=0}^{\infty} \varphi_n$  ,  $\varphi_0$ : homogeneous solution  $(\Delta_x + k^2) \varphi_0 = 0$

insert in (\*)

$\Rightarrow$   
power series in  
 $v$ , compare coeffs.

$$\varphi_n = \int d^3y G(\vec{x}-\vec{y}) v(\vec{y}) \varphi_{n-1}(\vec{y}) \quad , \quad n \geq 1$$

Mostly:  $\varphi \approx \varphi_0 + \varphi_1$

$f(\theta, \phi)$  is determined from  $\varphi_1$ .

## 7.5.2 Green Function

Solution of  $(\Delta^2 + k^2) G = \delta$  is not unique  $\rightarrow$  boundary conditions are needed

We take:  $\varphi_0 = e^{ikx_3}$  = incoming wave

$\Rightarrow \int d^3y G(\vec{x}-\vec{y}) v(\vec{y}) \varphi(\vec{y}) \xrightarrow{|\vec{x}| \rightarrow \infty}$  outgoing spherical wave

If we assume:  $v(\vec{y}) = 0$  for  $|\vec{y}| > R_0$ , then it must hold

$G(\vec{x}) \xrightarrow{|\vec{x}| \rightarrow \infty}$  outgoing spherical wave

Construction of  $G$ :

cf. electrostatics, there  $\Delta_{\vec{x}} \frac{-1}{4\pi} \frac{1}{|\vec{x}-\vec{y}|} = \delta^{(3)}(\vec{x}-\vec{y})$

Coulomb potential of a point charge, i.e. the case  $k^2 = 0$ .

Method for general  $k$ : Fourier:

$$\left. \begin{aligned} G(\vec{x}) &= \int \frac{d^3p}{(2\pi)^3} \tilde{G}(\vec{p}) e^{i\vec{p}\cdot\vec{x}} \\ \delta^{(3)}(\vec{x}) &= \int \frac{d^3p}{(2\pi)^3} e^{i\vec{p}\cdot\vec{x}} \\ (\Delta_{\vec{x}} + k^2) G(\vec{x}) &= \int \frac{d^3p}{(2\pi)^3} (-\vec{p}^2 + k^2) \tilde{G}(\vec{p}) e^{i\vec{p}\cdot\vec{x}} \end{aligned} \right\} \Rightarrow$$

$$\tilde{G}(\vec{p}) = \frac{1}{-\vec{p}^2 + k^2}$$

Where is the non-uniqueness?  $\tilde{G}$  is singular for  $\vec{p}^2 = k^2$ , the inverse transformation  $G(\vec{x}) = \int \frac{d^3p}{(2\pi)^3} \tilde{G}(\vec{p}) e^{i\vec{p}\cdot\vec{x}}$  is ill-defined.

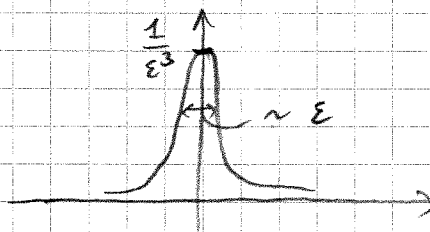
We define

$$G_{\varepsilon}(\vec{x}) = \int \frac{d^3p}{(2\pi)^3} \frac{e^{i\vec{p}\cdot\vec{x}}}{-\vec{p}^2 + (k+i\varepsilon)^2} \quad (k \rightarrow k+i\varepsilon)$$

denominator  $\neq 0$ ,  $G_{\varepsilon}(\vec{x})$  exists as a normal function, like

$$\lim_{\varepsilon \rightarrow 0^+} \frac{1}{\varepsilon^3} \frac{e^{-\frac{|\vec{x}|^2}{2\varepsilon^2}}}{(2\pi)^{3/2}} = \delta^{(3)}(\vec{x})$$

$$= \delta_{\varepsilon}^{(3)}(\vec{x})$$

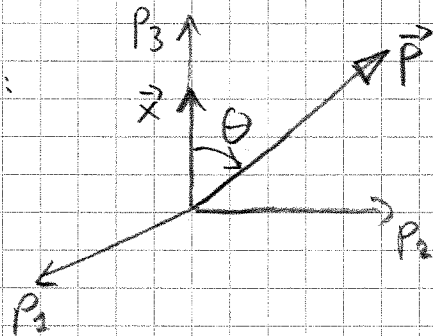


We will see that  $\lim_{\epsilon \rightarrow 0^+} G_\epsilon \neq \lim_{\epsilon \rightarrow 0^-} G_\epsilon \Leftrightarrow$  ambiguity

Computation of  $G_\epsilon(\vec{x})$ :

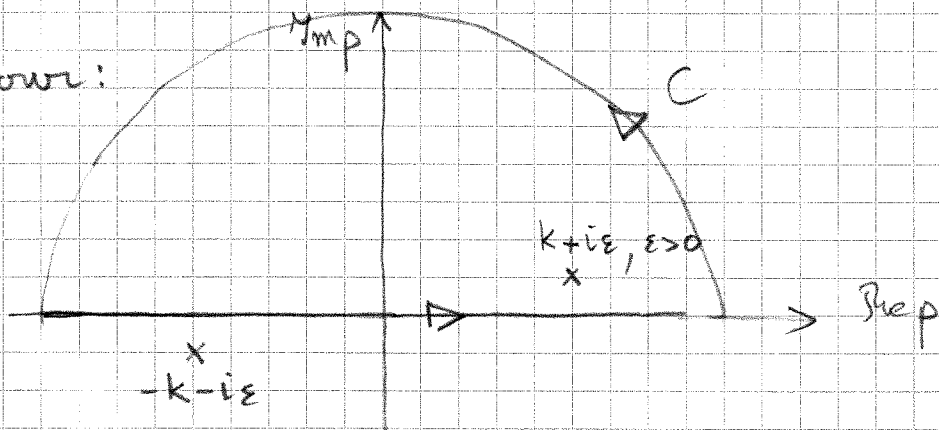
polar coordinates,  $p_3$  along  $\vec{x}$ :

$\vec{p} \cdot \vec{x} = p |\vec{x}| \cos \theta$ ,  $p = |\vec{p}|$



$$\begin{aligned}
 G_\epsilon(\vec{x}) &= \frac{1}{(2\pi)^3} \int_0^\infty dp p^2 \int_{-1}^1 d(\cos \theta) \int_0^{2\pi} d\phi \frac{e^{ip|\vec{x}| \cos \theta}}{-p^2 + (k+i\epsilon)^2} \\
 &= \frac{1}{4\pi^2} \int_0^\infty dp \frac{p^2}{-p^2 + (k+i\epsilon)^2} \left[ \frac{e^{ip|\vec{x}|u}}{ip|\vec{x}|} \right]_{u=-1}^{u=+1} \\
 &= \frac{1}{2\pi^2} \int_0^\infty dp \underbrace{\frac{\sin(p|\vec{x}|)}{|\vec{x}|}}_{\text{odd functions of } p} \underbrace{\frac{p}{-p^2 + (k+i\epsilon)^2}}_{\text{odd functions of } p} \\
 &= \frac{1}{4\pi^2 i |\vec{x}|} \int_{-\infty}^{\infty} dp e^{ip|\vec{x}|} \frac{-p}{(p-k-i\epsilon)(p+k+i\epsilon)}
 \end{aligned}$$

closed contour:



$\int_{-\infty}^{\infty} dp \dots = \oint_C dp \dots$  since  $e^{ip|\vec{x}|} = e^{i \text{Re } p |\vec{x}|} \cdot e^{-\text{Im } p |\vec{x}|} \xrightarrow{p \rightarrow \infty} 0$  on the circle with  $\text{Im } p > 0$

$\epsilon > 0$  :  $= 2\pi i$  residuum at  $p = k+i\epsilon$

$\epsilon < 0$  :  $= 2\pi i$  residuum at  $p = -k-i\epsilon$



$$\begin{aligned} \text{pole } p = k + i\varepsilon, \text{ residuum} &= \frac{1}{4\pi^2 i |\vec{x}|} e^{ik|\vec{x}| - \varepsilon|\vec{x}|} \left(-\frac{1}{2}\right) \\ (\varepsilon > 0) &= \frac{-1}{8\pi^2 i |\vec{x}|} e^{ik|\vec{x}|} + O(\varepsilon) \end{aligned}$$

$$\begin{aligned} \text{pole } p = -k - i\varepsilon, \text{ residuum} &= \frac{1}{4\pi^2 i |\vec{x}|} e^{-ik|\vec{x}| + \varepsilon|\vec{x}|} \left(-\frac{1}{2}\right) \\ (\varepsilon < 0) &= \frac{-1}{8\pi^2 i |\vec{x}|} e^{-ik|\vec{x}|} + O(\varepsilon) \end{aligned}$$

$$\Rightarrow G_{\pm}(\vec{x}) = \lim_{\varepsilon \rightarrow 0_{\pm}} G_{\varepsilon}(\vec{x}) = \underbrace{\frac{-1}{4\pi |\vec{x}|}}_{\text{Coulomb!}} e^{\pm ik|\vec{x}|}$$

$G_{\pm}$  correspond to outgoing (+) / incoming (-) spherical waves

$$\Rightarrow G_{+}(\vec{x}) = \frac{-1}{4\pi |\vec{x}|} e^{ik|\vec{x}|} \text{ for scattering}$$

### 7.5.3 Born Approximation

$$\varphi(\vec{x}) = \varphi_0(\vec{x}) + \int d^3y G_{+}(\vec{x}-\vec{y}) v(\vec{y}) \varphi(\vec{y})$$

$$\text{For } |\vec{x}| \rightarrow \infty: |\vec{x}-\vec{y}| = \sqrt{|\vec{x}|^2 + |\vec{y}|^2 - 2\vec{x}\cdot\vec{y}} \approx |\vec{x}| \left(1 - \frac{2\vec{x}\cdot\vec{y}}{|\vec{x}|^2}\right)^{1/2} \approx |\vec{x}| - \vec{y}\cdot\hat{x}, \hat{x} = \frac{\vec{x}}{|\vec{x}|}$$

$$\int_{|\vec{y}| \leq R_0} d^3y \frac{-1}{4\pi |\vec{x}|} e^{ik|\vec{x}-\vec{y}|} v(\vec{y}) \varphi(\vec{y}) \approx \frac{-e^{ik|\vec{x}|}}{4\pi |\vec{x}|} \int d^3y e^{-ik\vec{y}\cdot\hat{x}} v(\vec{y}) \varphi(\vec{y})$$

Comparison with  $\frac{e^{ikr}}{r} f(\theta, \phi)$  yields

$$f(\theta, \phi) = -\frac{m}{2\pi\hbar^2} \int d^3y e^{-ik\vec{y}\cdot\hat{x}} V(\vec{y}) \varphi(\vec{y})$$

Lowest order of the Born approximation:

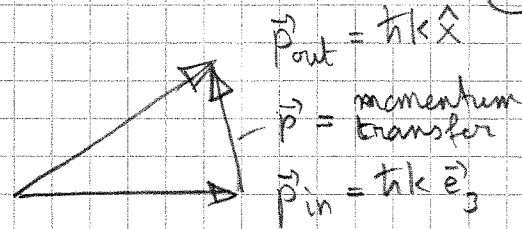
$$\varphi(\vec{y}) \approx \varphi_0(\vec{y}) = e^{ik\vec{y}\cdot\vec{e}_3}$$

$$\Rightarrow f(\theta, \phi) \approx -\frac{m}{2\pi\hbar^2} \int d^3y V(\vec{y}) e^{-ik\vec{y}\cdot(\vec{x} - \vec{e}_3)}$$

In this approximation:

Fourier transformation

$$f \sim \tilde{V}(\vec{p} = k(\hat{x} - \vec{e}_3) \hbar)$$



direction  $\hat{x} \leftrightarrow (\theta, \phi)$ , incoming wave parallel to  $\vec{e}_3$ ,  
 no symmetry  $V(|\vec{y}|)$  assumed  $\Rightarrow f(\theta, \phi)$  depends on  $\theta$  and  $\phi$

Validity?

it should be  $|e_1| \leq |e_0| = 1$  for large  $|\vec{x}|$

Analysis ( $\rightarrow$  Nolting) shows that the approximation is plausible for large energy  $kR_0 \gg 1$  and  $V = V(|\vec{y}|)$  when

$$\left| \int_0^\infty dr V(r) \right| \ll \frac{\hbar^2 k}{m}$$

### 7.5.4 Rutherford Scattering Formula

First: Yukawa potential

$$V(\vec{x}) = V_0 \frac{e^{-\alpha|\vec{x}|}}{|\vec{x}|} \quad (\alpha > 0)$$

Range  $R_0 = \frac{1}{\alpha}$  (Coulomb limit:  $\alpha \rightarrow 0$ )

Scattering amplitude in first (lowest) order Born approximation

$$f(\theta) = -\frac{mV_0}{2\pi\hbar^2} \int d^3y \frac{e^{-\alpha|\vec{y}|}}{|\vec{y}|} e^{-i\vec{K} \cdot \vec{y}}$$

where  $\vec{K} = k(\hat{x} - \vec{e}_3) = k(\sin\theta \cdot c, \sin\theta \cdot s, \cos\theta - 1)$   
 $c = \cos\phi, s = \sin\phi$

$$\hbar\vec{K} = \underbrace{\vec{p}_{out}}_{\hbar k \hat{x}} - \underbrace{\vec{p}_{in}}_{\hbar k \vec{e}_3}$$

Evaluate in polar coordinates ( $\rightarrow$  exercise):

$$\Rightarrow F(\theta) = -\frac{2mV_0}{\hbar^2} \frac{1}{\alpha^2 + \vec{K}^2}$$

$$\vec{K}^2 = k^2 (\hat{x} - \hat{e}_3)^2 = k^2 (2 - 2\hat{x}_3) = k^2 2 (1 - \cos\theta) = 4k^2 \sin^2 \frac{\theta}{2}$$

$$\frac{d\sigma}{d\Omega} = |F|^2 = \frac{4m^2 V_0^2}{\hbar^4} \frac{1}{(\alpha^2 + 4k^2 \sin^2 \frac{\theta}{2})^2}$$

$$\sigma_{\text{tot}} = \int d\Omega \frac{d\sigma}{d\Omega} = \frac{4m^2 V_0^2}{\hbar^4} \frac{4\pi}{\alpha^2 (\alpha^2 + 4k^2)}$$

Coulomb limit:  $\alpha \downarrow 0$ ,  $V_0 = \frac{q_1 q_2}{4\pi \epsilon_0}$

exact Coulomb scattering amplitude:

$$f_c(\theta) = -\frac{\gamma}{4k \sin^2 \frac{\theta}{2}} e^{(2i\sigma_0 - i\gamma \ln(\sin^2 \frac{\theta}{2}))}$$

$$\gamma = \frac{V_0 m}{\hbar^2 k}$$

differs from Born amplitude only by a phase factor, although Born approximation assumes a finite range ( $\rightarrow$  more complicated treatment is needed).

$$\frac{d\sigma}{d\Omega} = \frac{V_0}{E^2} \frac{1}{16 \sin^4 \frac{\theta}{2}} \quad (\text{Rutherford})$$

$\sigma_{\text{tot}} = \infty$ : in Coulomb potential there is a small deflection even for arbitrary large impact parameter  $b$   
 $\rightarrow$  unphysical, since charge is always screened at very large distances

# 7.6 Formalism of Scattering Theory

up to now: stationary scattering problem in position space  $\rightarrow$  more general formalism

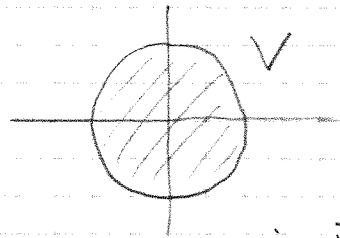
$$H = H_0 + V(\vec{x})$$

$V \xrightarrow{|\vec{x}| \rightarrow \infty} 0$  fast "enough" to be able to use asymptotically the eigenstates of  $H_0$

## 7.6.1 Wave Operator (Møller Operator)

$$i\hbar \frac{d}{dt} |\varphi_-(t)\rangle = H_0 |\varphi_-(t)\rangle$$

represents free incoming state in the limit  $t \rightarrow -\infty$ ,



for example:  $\langle \vec{p} | \varphi_-(t) \rangle = e^{-\frac{i}{\hbar} \frac{\vec{p}^2}{2m} t} \tilde{\varphi}(\vec{p})$

$$i\hbar \frac{d}{dt} |\psi(t)\rangle = H |\psi(t)\rangle$$

represents the full exact solution with

$$\lim_{t \rightarrow -\infty} | \psi(t) \rangle - | \varphi_-(t) \rangle = 0$$

$$\begin{aligned} |\psi(t)\rangle &= e^{-\frac{i}{\hbar} (t-\tau) H} |\psi(\tau)\rangle \\ &= \lim_{\tau \rightarrow -\infty} e^{-\frac{i}{\hbar} (t-\tau) H} |\varphi_-(\tau)\rangle \\ &= \lim_{\tau \rightarrow -\infty} \underbrace{e^{-\frac{i}{\hbar} (t-\tau) H}}_{U(t, \tau)} \underbrace{e^{+\frac{i}{\hbar} (t-\tau) H_0}}_{U_0(\tau, t)} |\varphi_-(\tau)\rangle \end{aligned}$$

$$|\psi(t)\rangle = \Omega_{in} |\varphi_-(t)\rangle$$

$$\begin{aligned} \Omega_{in} &:= \lim_{\tau \rightarrow -\infty} e^{-\frac{i}{\hbar}(t-\tau)H} e^{\frac{i}{\hbar}(t-\tau)H_0} \\ &= \lim_{\tau \rightarrow -\infty} e^{+\frac{i}{\hbar}\tau H} e^{-\frac{i}{\hbar}\tau H_0} \end{aligned}$$

$\Omega_{in}$  (free "incoming" solution) = (full scattering solution)

Analogously: free "outgoing" solution  $|\varphi_+(t)\rangle$ :

$$|\psi(t)\rangle = \Omega_{out} |\varphi_+(t)\rangle$$

$$\Omega_{out} = \lim_{\tau \rightarrow +\infty} e^{\frac{i}{\hbar}\tau H} e^{-\frac{i}{\hbar}\tau H_0}$$

Convergence of operator for  $\tau \rightarrow \pm\infty$  difficult to show, theorems, depend on  $V$

Physical expectation: for a given  $|\varphi_-(t)\rangle$  there exists  $|\varphi_+(t)\rangle$  such that

$$\Omega_{in} |\varphi_-(t)\rangle = \Omega_{out} |\varphi_+(t)\rangle$$

$$\Omega_{in} \mathcal{M} = \Omega_{out} \mathcal{M} = \mathcal{M}_S = \text{space of scattering states}$$

$\mathcal{M}_B$ : space of bound states: it holds

$$\mathcal{M}_B \perp \mathcal{M}_S$$

heuristic "proof":

$$H|n\rangle = E_n|n\rangle, \quad E_n < 0$$

$$e^{+\frac{i}{\hbar}E_n t} \langle n | \Omega_{in} |\varphi_-(t)\rangle = \langle n | \underbrace{e^{+\frac{i}{\hbar}Ht} \Omega_{in} |\varphi_-(t)\rangle}_{\text{localized incoming packet for } t \rightarrow -\infty} \xrightarrow{t \rightarrow -\infty} 0$$

rigorous proof for suitable potentials.

## 7.6.2 S-Matrix, T-Matrix

Probability to find a specific outgoing state  $|\varphi_+(t)\rangle$  given an incoming state  $|\varphi_-(t)\rangle$

$$W = \left| \langle \varphi_+(t) | \Omega_{\text{out}}^+ \Omega_{\text{in}} | \varphi_-(t) \rangle \right|^2$$

$W$  is independent of  $t$ , since

$$i\hbar \frac{d}{dt} \underbrace{\Omega_{\text{in}} | \varphi_-(t) \rangle}_{= |\psi(t)\rangle} = H \Omega_{\text{in}} | \varphi_-(t) \rangle$$

$$i\hbar \frac{d}{dt} \underbrace{\langle \varphi_+(t) | \Omega_{\text{out}}^+}_{= \langle \psi(t) |} = - \langle \varphi_+(t) | \Omega_{\text{out}}^+ H$$

$|\varphi_{\pm}\rangle := |\varphi_{\pm}(t=0)\rangle \Leftrightarrow$  Heisenberg states

$$W = \left| \langle \varphi_+ | S | \varphi_- \rangle \right|^2$$

$S = \Omega_{\text{out}}^+ \Omega_{\text{in}}$  : S-matrix, scattering matrix

$$S S^+ = \mathbb{1} = S^+ S \quad (\text{since } \Omega_x \Omega_x^+ = \mathbb{1} = \Omega_x^+ \Omega_x, x = \text{in, out})$$

$$S = \lim_{\substack{\tau_+ \rightarrow \infty \\ \tau_- \rightarrow -\infty}} e^{\frac{i}{\hbar} \tau_+ H_0} e^{-\frac{i}{\hbar} (\tau_+ - \tau_-) H} e^{-\frac{i}{\hbar} \tau_- H_0}$$

$$= \lim_{\substack{\tau_+ \rightarrow \infty \\ \tau_- \rightarrow -\infty}} U_{\rightarrow}(\tau_+, \tau_-)$$

$\hookrightarrow$  interaction picture

Limit of sharp energy:

$$|\varphi\rangle = \int d^3p \tilde{\varphi}(\vec{p}) |\vec{p}\rangle \rightarrow \text{discussion for } |\vec{p}\rangle$$

$$\langle \vec{p} | \vec{p}' \rangle = \delta^{(3)}(\vec{p} - \vec{p}') ; H_0 |\vec{p}\rangle = E_{\vec{p}} |\vec{p}\rangle ; E_{\vec{p}} = \frac{\vec{p}^2}{2m}$$

$$|\vec{p} \text{ in}\rangle = \Omega_{\text{in}} |\vec{p}\rangle$$

$$\text{Proposition: } H |\vec{p} \text{ in}\rangle = E_{\vec{p}} |\vec{p} \text{ in}\rangle$$

↓  
full Hamiltonian!

Proof:

$$e^{\frac{i}{\hbar} \tau H} e^{-\frac{i}{\hbar} \tau H_0} |\vec{p}\rangle \xrightarrow{\tau \rightarrow -\infty} |\vec{p} \text{ in}\rangle \text{ should converge}$$

$$\Rightarrow i\hbar \frac{d}{d\tau} \left( \quad \right) \xrightarrow{\tau \rightarrow -\infty} 0$$

$$= -H |\vec{p} \text{ in}\rangle + \Omega_{\text{in}} H_0 |\vec{p}\rangle$$

$$= (E_{\vec{p}} - H) |\vec{p} \text{ in}\rangle \quad \square$$

$$\langle \vec{p} \text{ in} | \vec{p}' \text{ in} \rangle = \delta^{(3)}(\vec{p} - \vec{p}')$$

$$\text{Likewise: } |\vec{p} \text{ out}\rangle = \Omega_{\text{out}} |\vec{p}\rangle$$

$$\langle \vec{p}' | S | \vec{p} \rangle = \langle \vec{p}' | \Omega_{\text{out}}^\dagger \Omega_{\text{in}} | \vec{p} \rangle$$

$$= \langle \vec{p}' \text{ out} | \vec{p} \text{ in} \rangle$$

$$=: \delta^{(3)}(\vec{p}' - \vec{p}) - 2\pi i \delta(E_{\vec{p}'} - E_{\vec{p}}) T(\vec{p}', \vec{p})$$

$\uparrow$  contribution for  $V=0$        $\uparrow$  eigenstates of  $H$        $\nwarrow$   $T$ -matrix, defined here with  
 orthogonal when  $E_{\vec{p}'} \neq E_{\vec{p}}$

It can be shown:

$$\vec{p} = (0, 0, p)$$

$$\vec{p}' = p (\sin\theta \cos\varphi, \sin\theta \sin\varphi, \cos\theta)$$

$$\Rightarrow \frac{d\Omega}{d\Omega} = (2\pi)^4 (m\hbar)^2 |T(\vec{p}', \vec{p})|^2$$

### 7.6.3 Lippmann-Schwinger Equations

$$\Omega_{out} = \int_0^\infty dt \frac{d}{dt} e^{\frac{i}{\hbar} H t} e^{-\frac{i}{\hbar} H_0 t} + \mathbb{1}$$

$$= \mathbb{1} + \frac{i}{\hbar} \int_0^\infty dt e^{\frac{i}{\hbar} H t} \underbrace{V}_{=H-H_0} e^{-\frac{i}{\hbar} H_0 t}$$

$$\Omega_{out} |\varphi\rangle = |\varphi\rangle + \frac{i}{\hbar} \int_0^\infty dt \int d^3p \tilde{\varphi}(\vec{p}) e^{\frac{i}{\hbar} H t} V e^{-\frac{i}{\hbar} H_0 t} |\vec{p}\rangle$$

$\uparrow$   
 $e^{-\frac{i}{\hbar} E_p t}$

$$= |\varphi\rangle - \lim_{\varepsilon \rightarrow 0} \int d^3p \tilde{\varphi}(\vec{p}) (H - E_p + i\varepsilon)^{-1} V |\vec{p}\rangle$$

$$\frac{-i}{\hbar} \int_0^\infty dt e^{\frac{i}{\hbar} (H - E_p + i\varepsilon) t} = (H - E_p + i\varepsilon)^{-1}$$

$\varepsilon > 0$  : damps the integrand for  $t \rightarrow \infty$   
 needed to interchange  $\int dt \leftrightarrow \int d^3p$ , i.e. to act on  $|\vec{p}\rangle$

$$|\vec{p} \text{ out}\rangle = |\vec{p}\rangle - (H - E_p + i\varepsilon)^{-1} V |\vec{p}\rangle$$

$$|\vec{p} \text{ in}\rangle = |\vec{p}\rangle - (H - E_p - i\varepsilon)^{-1} V |\vec{p}\rangle$$

$\uparrow$  ! convergence for  $t \rightarrow -\infty$

$$|\vec{p} \text{ out}\rangle = \left( \mathbb{1} - \underbrace{(H - E_p + i\varepsilon)^{-1} V}_{(H-z)^{-1}} \right) |\vec{p}\rangle$$

$$= (H-z)^{-1}$$

Invert the relation:

$$\left( \mathbb{1} + (H_0 - z)^{-1} V \right) \left( \mathbb{1} - (H - z)^{-1} V \right) =$$

$$= \mathbb{1} + (H_0 - z)^{-1} V - (H - z)^{-1} V - (H_0 - z)^{-1} \underbrace{(H - H_0)}_V (H - z)^{-1} V$$

$$= \mathbb{1} + (H_0 - z)^{-1} V - (H - z)^{-1} V - (H_0 - z)^{-1} (H - z)^{-1} V$$

$$= \mathbb{1}$$



$$\Rightarrow |\vec{p}\rangle = \left( \mathbb{1} + (H_0 - E_{\vec{p}} + i\epsilon)^{-1} V \right) |\vec{p}\text{ out}\rangle$$

$$|\vec{p}\rangle = |\vec{p}\text{ out}\rangle + (H_0 - E_{\vec{p}} + i\epsilon)^{-1} V |\vec{p}\text{ out}\rangle$$

$$|\vec{p}\rangle = |\vec{p}\text{ in}\rangle + (H_0 - E_{\vec{p}} - i\epsilon)^{-1} V |\vec{p}\text{ in}\rangle$$

In position space: identity

$$\left( -\frac{\hbar^2}{2m} \Delta_x - \frac{\vec{p}^2}{2m} \pm i\epsilon \right) \langle \vec{x} | (H_0 - E_{\vec{p}} \pm i\epsilon)^{-1} | \vec{x}' \rangle = \delta^{(3)}(\vec{x} - \vec{x}')$$

From section 7.5.2 we know:

$$\left( \Delta + \underbrace{(k \pm i\epsilon)^2}_{= k^2 \pm 2i\epsilon k} \right) G_{\pm}(\vec{x}) = \delta^{(3)}(\vec{x})$$

Comparison  $\Rightarrow$

$$\begin{aligned} \langle \vec{x} | (H_0 - E_{\vec{p}} \pm i\epsilon)^{-1} | \vec{x}' \rangle &= -\frac{2m}{\hbar^2} G_{\mp}(\vec{x} - \vec{x}') \\ &= \frac{m}{2\pi\hbar^2 |\vec{x} - \vec{x}'|} e^{\mp \frac{i}{\hbar} |\vec{p}| |\vec{x} - \vec{x}'|} \end{aligned}$$

### 7.6.4 Born Series

$$\begin{aligned} \langle \vec{p}' \text{ out} | \vec{p} \text{ in} \rangle &= \delta^{(3)}(\vec{p}' - \vec{p}) + \left( \langle \vec{p}' \text{ out} | - \langle \vec{p}' \text{ in} | \right) |\vec{p} \text{ in}\rangle \\ &= \delta^{(3)}(\vec{p}' - \vec{p}) - \langle \vec{p}' | V \left[ \underbrace{(H - E_{\vec{p}'} - i\epsilon)^{-1}}_{\langle \vec{p}' \text{ out} |} - \underbrace{(H - E_{\vec{p}'} + i\epsilon)^{-1}}_{\langle \vec{p}' \text{ in} |} \right] |\vec{p} \text{ in}\rangle \end{aligned}$$

$$H |\vec{p} \text{ in}\rangle = E_{\vec{p}} |\vec{p} \text{ in}\rangle$$

$$\lim_{\epsilon \rightarrow 0} \left[ (u - i\epsilon)^{-1} - (u + i\epsilon)^{-1} \right] = 2\pi i \delta(u)$$

$$\Rightarrow [ \quad ] \rightarrow 2\pi i \delta(E_{\vec{p}} - E_{\vec{p}'})$$

Comparison with the definition of the T-matrix =>

$$T(\vec{p}, \vec{p}') = \langle \vec{p}' | V | \vec{p} \text{ in} \rangle$$

$$= \langle \vec{p}' | V [ | \vec{p} \rangle - (H_0 - E_{\vec{p}} - i\epsilon)^{-1} V | \vec{p} \text{ in} \rangle ]$$

↑  
! see top page (103)

Series is obtained by recursively replacing

$$| \vec{p} \text{ in} \rangle \rightarrow | \vec{p} \rangle - \dots$$

Leading order:

$$T(\vec{p}, \vec{p}') \approx \langle \vec{p}' | V | \vec{p} \rangle = \tilde{V}(\vec{p}' - \vec{p})$$

$$= (2\pi\hbar)^{-3} \int d^3x e^{-\frac{i}{\hbar}(\vec{p}' - \vec{p}) \cdot \vec{x}} V(\vec{x})$$

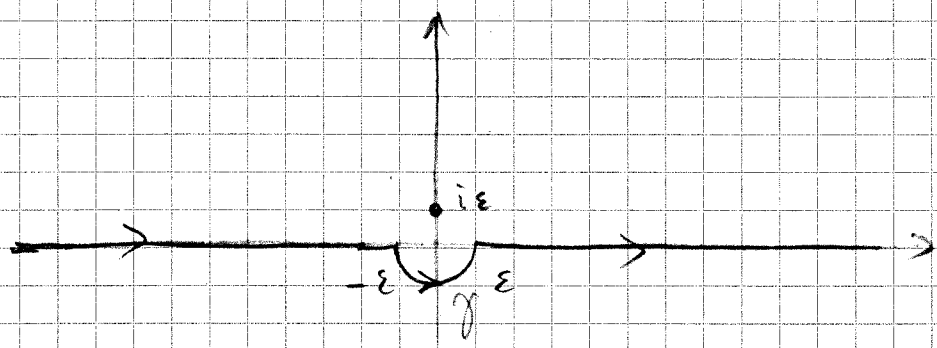
reproduce the result of section 7.5.3.

Proposition:  $\lim_{\epsilon \rightarrow 0} \left[ \frac{1}{u-i\epsilon} - \frac{1}{u+i\epsilon} \right] = 2\pi i \delta(u)$

Proof: we show:

$$\int_{-\infty}^{\infty} du f(u) \left[ \frac{1}{u-i\epsilon} - \frac{1}{u+i\epsilon} \right] \xrightarrow{\epsilon \rightarrow 0} 2\pi i \cdot f(0)$$

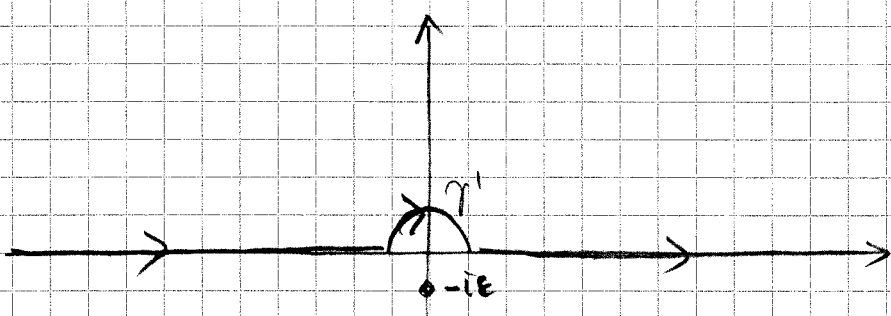
$\frac{1}{u-i\epsilon}$ :



$$\int_{-\infty}^{\infty} du f(u) \frac{1}{u-i\epsilon} = \int_{-\infty}^{-\epsilon} + \int_{\gamma} + \int_{\epsilon}^{+\infty}$$

$$\int_{\gamma} du f(u) \frac{1}{u-i\epsilon} \xrightarrow{\epsilon \rightarrow 0} + i \cdot (\text{residuum } u=0) \pi = i\pi f(0)$$

$\frac{1}{u+i\epsilon}$ :



$$\int_{-\infty}^{\infty} du f(u) \frac{1}{u+i\epsilon} = \int_{-\infty}^{-\epsilon} + \int_{\gamma'} + \int_{\epsilon}^{\infty}$$

$$\int_{\gamma'} du f(u) \frac{1}{u+i\epsilon} \xrightarrow{\epsilon \rightarrow 0} -i \cdot (\text{residuum } u=0) \pi = -i\pi f(0)$$

↑  
orientation

$$\Rightarrow \int_{-\infty}^{\infty} du f(u) \left[ \frac{1}{u-i\epsilon} - \frac{1}{u+i\epsilon} \right] = \int_{\gamma} - \int_{\gamma'} = 2\pi i f(0) \quad \square$$

# 8. Relativistic Quantum Mechanics

## 8.1 Classical Relativistic Mechanics

Non-relativistic mechanics and quantum mechanics (Schrödinger-, Pauli-equation, ...) hold only when

$$|\vec{v}| = \frac{|\vec{p}|}{m} \ll c \approx 3 \times 10^8 \frac{\text{m}}{\text{sec}}$$

only then:  $E = \frac{p^2}{2m}$

and  $\hat{H} = \frac{\hat{p}^2}{2m} \dots$

Otherwise:

$p^\mu = \left( \frac{E}{c}, \vec{p} \right)$  : four-vector,  $\mu = 0, 1, 2, 3$   
time - components      space - components

$p^2 = p^\mu p^\nu g_{\mu\nu} = \frac{E^2}{c^2} - \vec{p}^2 = m^2 c^2$  : Lorentz-scalar (see below)

$g_{\mu\nu} = \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{pmatrix}$  diagonal matrix

$\vec{p} = \frac{m\vec{v}}{\sqrt{1-\beta^2}}$  ← consistency with non-relativistic limit, see Jackson

$E = \sqrt{m^2 c^4 + c^2 \vec{p}^2} = \frac{m c^2}{\sqrt{1-\beta^2}}, \quad \beta^2 = \frac{v^2}{c^2} < 1$

When  $\beta \ll 1$ :

$E = m c^2 \left( 1 + \frac{1}{2} \beta^2 + \frac{3}{8} \beta^4 + O(\beta^6) \right) = m c^2 + \frac{m}{2} v^2 + \dots$   
rest-energy      non-relat. term

$x^\mu = (ct, \vec{x})$  space-time vector

Lorentz-transformation:  $x^\mu \rightarrow x'^\mu = \Lambda^\mu_\nu x^\nu$

leaves scalar product of 4-vectors invariant:

$$x' \cdot y' = x \cdot y \Leftrightarrow g_{\mu\nu} \Lambda^\mu_\kappa \Lambda^\nu_\lambda = g_{\kappa\lambda} = \begin{pmatrix} 1 & -1 & 0 \\ 0 & -1 & -1 \end{pmatrix}$$

$$(x \cdot y \equiv g_{\mu\nu} x^\mu y^\nu)$$

"generalized rotations", pseudoorthogonal  
= Lorentz-group, contains:

\* rotations:

$$x'^0 = x^0$$

$$x'^i = R^i_k x^k, \quad i, k = 1, 2, 3, \quad R^i_k: \text{rotation matrix as before}$$

\* boosts, for example:

$$p'^1 = p^1$$

$$p'^2 = p^2$$

$$\begin{pmatrix} p'^0 \\ p'^3 \end{pmatrix} = \begin{pmatrix} \cosh \alpha & \sinh \alpha \\ \sinh \alpha & \cosh \alpha \end{pmatrix} \begin{pmatrix} p^0 \\ p^3 \end{pmatrix}; \quad \alpha \in \mathbb{R}$$

$$p'^2 \equiv p' \cdot p' = p^2 \quad [ \cosh^2 \alpha - \sinh^2 \alpha = 1 ]$$

$$p = mc \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix}$$

↑  
particle at rest

$$\xrightarrow{\text{boost}} p' = mc \begin{pmatrix} \cosh \alpha \\ 0 \\ 0 \\ \sinh \alpha \end{pmatrix}$$

↓  
particle (or observer) moving

Equation of motion: (holds independently of the reference frame)

$$\frac{dx^\mu}{d\tau} = \frac{1}{m} p^\mu \Rightarrow dx^0 = c dt = \frac{p^0}{m} d\tau \Rightarrow \frac{d\vec{x}}{dt} = c \frac{\vec{p}}{p^0}$$

$$d\vec{x} = \frac{\vec{p}}{m} d\tau$$

In moving frame:  $\frac{d\vec{x}'}{dt'} = c \frac{\vec{p}'}{p'^0} = \underbrace{c \tanh \alpha}_{=v < c} \vec{e}_3$

$$\left[ d\tau = \frac{mc}{p^0} dt = \sqrt{1-\beta^2} dt : \text{proper time, Lorentz-invariant} \right] = v < c$$

General infinitesimal Lorentz - transformation:

$$\Lambda^\mu_\kappa = \delta^\mu_\kappa + \omega^\mu_\kappa \quad (\omega^2 \approx 0)$$

$$g_{\mu\nu} \Lambda^\mu_\kappa \Lambda^\nu_\lambda = g_{\kappa\lambda} \Rightarrow g_{\mu\nu} (\omega^\mu_\kappa \delta^\nu_\lambda + \delta^\mu_\kappa \omega^\nu_\lambda) \stackrel{!}{=} 0$$
$$\omega_{\nu\kappa} \delta^\nu_\lambda + \delta^\mu_\kappa \omega_{\mu\lambda} = \omega_{\lambda\kappa} + \omega_{\kappa\lambda} \stackrel{!}{=} 0$$

$\Rightarrow \omega_{\mu\nu} = g_{\mu\lambda} \omega^\lambda_\nu$  must be anti-symmetric!

$\rightarrow \omega_{\mu\nu}$  has  $\frac{1}{2} \cdot 4 \cdot 3 = 6$  free parameters

\* rotations:

$$(\omega^\cdot) = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & \varphi \vec{n} \cdot \vec{T} \\ 0 & & & \\ 0 & & & \end{pmatrix}$$

$T^i$ : generators of infinitesimal rotations in  $\mathbb{R}^3$ :

$$(\varphi \vec{n} \cdot \vec{T} \vec{x})_k = \varphi (\vec{n} \times \vec{x})_k, \quad (T^i)_{ke} = \epsilon_{kie}$$

$$\omega_{ne} = g_{nk} \omega^k_e = -(\varphi \vec{n} \cdot \vec{T})_{ne} \text{ anti-symmetric}$$

\* boosts:  $\sinh \alpha \approx \alpha, \cosh \alpha \approx 1$

$$(\omega^\cdot) = \begin{pmatrix} 0 & \alpha \\ \alpha & 0 \end{pmatrix} \quad (\text{only } \nu=0,3 \text{ shown})$$

$$(\omega^{\cdot\cdot}) = \begin{pmatrix} 0 & \alpha \\ -\alpha & 0 \end{pmatrix} \quad \text{anti-symmetric}$$

Gradient:

$$x'^\mu = \Lambda^\mu_\nu x^\nu \Rightarrow \frac{\partial x'^\mu}{\partial x^\nu} = \Lambda^\mu_\nu$$

$$\partial_\mu = \left( \frac{\partial}{\partial(ct)}, \vec{\nabla} \right) = \frac{\partial}{\partial x^\mu}$$

$$\Rightarrow \partial_\mu = \frac{\partial x'^\nu}{\partial x^\mu} \frac{\partial}{\partial x'^\nu} \quad (\text{chain rule})$$

$\partial_\mu = \Lambda^\nu_\mu \partial'_\nu \rightarrow$  covariant transformation rule

$x^\mu$  has a contravariant transformation rule  $x'^\mu = \Lambda^\mu_\nu x^\nu$

$\partial'_\mu = \frac{\partial x^\nu}{\partial x'^\mu} \partial_\nu = \underbrace{\Lambda^\nu_\mu}_{\text{new, introduced here}} \partial_\nu$

$(\Lambda^\cdot_\cdot)^T, (\Lambda^\cdot_\cdot)$  are inverse to each other

$\Lambda^\kappa_\mu \Lambda^\nu_\kappa = \delta_\mu^\nu$

Since  $\Lambda$  is pseudoorthogonal the inversion is easy,  $\approx$  transpose

Definition:  $g_{\mu\kappa} g^{\kappa\nu} = \delta_\mu^\nu \Rightarrow g^{\cdot\cdot} = \begin{pmatrix} +1 & & \\ & -1 & 0 \\ & & -1 \end{pmatrix}$

Definition of Lorentz-transf.:  $g_{\kappa\lambda} \Lambda^\kappa_\mu \Lambda^\lambda_\sigma = g_{\mu\sigma}$   
 $\Leftrightarrow g_{\kappa\lambda} \Lambda^\kappa_\mu \Lambda^\lambda_\sigma g^{\sigma\nu} = \delta_\mu^\nu$

$\Rightarrow \Lambda^\nu_\kappa = g_{\kappa\lambda} \Lambda^\lambda_\sigma g^{\sigma\nu}$

## 8.2 Preliminary Considerations

### 8.2.1 Invariance of Wave Equations

Schrödinger- and Pauli-equations are not Lorentz-invariant:

$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi$

1st order, 2nd order, asymmetric space  $\leftrightarrow$  time

$\psi(\vec{x}, t) = \psi(x^\mu)$  solution  $\rightarrow \psi(x')$ ,  $x'^\mu = \Lambda^\mu_\nu x^\nu$  is in general not a solution

It was different for Maxwell equations:

Maxwell equations  $\Rightarrow$  discovery of Lorentz-invariance

$\Rightarrow$  look for a new wave equation

-  $\frac{\partial}{\partial t}$ ,  $\nabla$  both first or both second order

in particular for second order: interpretation? need

$\psi$  and  $\frac{\partial \psi}{\partial t}$  as initial conditions

- Lorentz-invariant!

- yields Schrödinger- or Pauli-equation in non-relativistic limit

### 8.2.2 Difficulty with One-particle Quantum Mechanics

Is for the particle number was conserved:

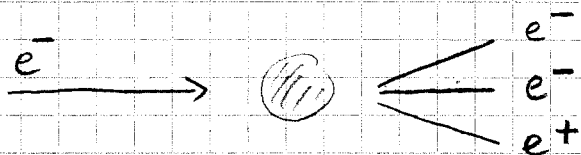
$$\rho(\vec{x}) = \psi^*(\vec{x})\psi(\vec{x}) \sim \text{probability to find the particle at } \vec{x}$$
$$\int d^3x \rho(\vec{x}) = \text{fixed (for example } = 1)$$

$$\frac{d}{dt} \int d^3x \rho(\vec{x}) = 0 \iff \text{Schrödinger equation}$$

$\psi(t, \vec{x}_1, \vec{x}_2)$ : two particles (for example 2 electrons in He atom)

relativistic:  $E = mc^2$  for a particle at rest

$\Rightarrow$  creation, annihilation possible when  $E_{kin} > mc^2$



Problem of principles with  $\psi(t, \vec{x}) \xrightarrow{?} \psi(t, \vec{x}_1, \vec{x}_2, \vec{x}_3)$

Solution: Quantum Field Theory:

$$\mathcal{H} = \mathcal{H}_{\text{vacuum}} \oplus \mathcal{H}_{\text{one-particle}} \oplus \mathcal{H}_{\text{two-particles}} \oplus \dots$$



Fock-space, transitions between different "sectors" are possible

but: small relativistic corrections, before particle production plays a role can be described by a modified wave equation:



$$E < n \cdot mc^2$$

like for example  $E = mc^2 + \frac{p^2}{2m} - \frac{p^4}{8m^3c^2} + \dots$

↑  
relat. correction

but in general get inconsistency when trying to formulate (fully) relativistic one-particle theory

### 8.3 Klein Gordon Equation

$$\partial_\mu = \frac{\partial}{\partial x^\mu} = \left( \frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla} \right) \text{ covariant 4-vector operator}$$

Scalar:  $g^{\mu\nu} \partial_\mu \partial_\nu = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \Delta = \partial^2 = \square$

d'Alembert  $\leftrightarrow$  wave equation of electrodynamics

We try to preserve the correspondences

$$i\hbar \frac{\partial}{\partial t} \leftrightarrow \text{energy}$$

$$\frac{\hbar}{i} \vec{\nabla} \leftrightarrow \text{momentum}$$

$$\rightarrow -\hbar^2 c^2 \partial^2 \leftrightarrow E^2 - c^2 \vec{p}^2 = m^2 c^4$$

Trial:

$$\left( -\hbar^2 \frac{\partial^2}{\partial t^2} + \hbar^2 c^2 \Delta \right) \psi = (mc^2)^2 \psi \quad (\Delta \equiv \vec{\nabla}^2)$$

$$\Leftrightarrow -\partial^2 \psi = \left( \frac{mc}{\hbar} \right)^2 \psi \quad \text{Klein Gordon equation (K.G.)}$$

Solution: plain wave:

$$\psi = e^{-\frac{i}{\hbar} (Et - \vec{p} \cdot \vec{x})} = e^{-\frac{i}{\hbar} p^\mu x_\mu}$$

Schrödinger equation $\Rightarrow E = \frac{p^2}{2m}$	} for free particles
Klein Gordon equation $\Rightarrow E^2 = \vec{p}^2 c^2 + m^2 c^4$	

$$E = \pm \sqrt{\vec{p}^2 c^2 + m^2 c^4} : \text{negative energy?}$$

Interpretation:

current, probability density:

$$\partial^\mu (\psi^* \overleftrightarrow{\partial}_\mu \psi) = \psi^* \partial^2 \psi - (\partial^2 \psi)^* \psi \stackrel{\text{K.G.}}{=} 0$$

$$[\text{definitions: } \partial^\mu = g^{\mu\nu} \partial_\nu = \left( \frac{1}{c} \frac{\partial}{\partial t}, -\vec{\nabla} \right)]$$

$$\int \overleftrightarrow{\partial}_\mu g = \int (\partial_\mu g) - (\partial_\mu g)$$

$$(c\rho, \vec{j}) = j^\mu := \frac{i\hbar}{2m} \psi^* \overleftrightarrow{\partial}^\mu \psi$$

$$\Rightarrow 0 = \partial_\mu j^\mu = \frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j}$$

$$\vec{j} = \frac{\hbar}{2mi} \psi^* \overleftrightarrow{\nabla} \psi \quad \text{like Schrödinger}$$

$$\rho = \frac{i\hbar}{2mc^2} \psi^* \overleftrightarrow{\partial} \psi \neq \psi^* \psi, \text{ in general } \underline{\text{not}} \geq 0!$$

$$\text{However when } \psi = e^{-\frac{i}{\hbar} E t} \varphi(\vec{x})$$

$$\Rightarrow \rho = \frac{E}{mc^2} \varphi^* \varphi$$

positive (negative) for positive (negative) energy!

- reinterpretation as charge density  $\rho \rightarrow q \cdot \rho$ ,  
 $E < 0 \rightarrow$  antiparticle, charge  $= -q$

- small relativistic corrections: waves with only  
 $E > 0$  components

- coupling to electromagnetic field

$$\left. \begin{array}{l} \frac{\partial}{\partial t} \rightarrow \frac{\partial}{\partial t} + \frac{iq}{\hbar} \phi \\ \vec{\nabla} \rightarrow \vec{\nabla} - \frac{iq}{\hbar} \vec{A} \end{array} \right\} \text{minimal substitution}$$

$$\partial^\mu \rightarrow \partial^\mu + \frac{iq}{\hbar} A^\mu ; \quad A^\mu = \left( \frac{\phi}{c}, \vec{A} \right)$$

because of gauge invariance.  $q = \text{charge}$ .

$$-\partial^2 \psi = \left( \frac{mc}{\hbar} \right)^2 \psi \rightarrow - \left( \partial^\mu + \frac{iq}{\hbar} A^\mu \right) \left( \partial_\mu + \frac{iq}{\hbar} A_\mu \right) \psi = \left( \frac{mc}{\hbar} \right)^2 \psi$$

K.G. equation in electromagnetic potential

If  $\psi$  is solution for charge  $q$

$\rightarrow \psi^*$  is solution for charge  $-q$

- eigenstates in Coulomb potential can be computed exactly
- $E_{n,l}$ :  $l$ -degeneracy of hydrogen atom is lifted but does not reproduce measurements
- better: relativistic Dirac equation
- Klein Gordon: spin 0 - mesons, for example  $\pi^+, \pi^-$
- negative energy  $\rightarrow$  Quantum Field Theory, antiparticles

## 8.4 Dirac Equation

### 8.4.1 Pauli Equation Revisited

Idea: first order differential equation

$$i\hbar \frac{\partial \psi}{\partial t} \rightarrow i\hbar \partial_\mu \psi$$

$\rightarrow$  with what can index  $\mu$  of  $\partial_\mu$  be contracted, compare  $\partial_\mu \partial^\mu$ , to obtain a covariant equation?

$\rightarrow$  Spin, Pauli:

$$\psi(\vec{x}) = \begin{pmatrix} \psi_+(\vec{x}) \\ \psi_-(\vec{x}) \end{pmatrix}, \quad \text{spin } \frac{1}{2}$$

(active) rotation

$$\psi(\vec{x}) \rightarrow \psi'(\vec{x}) = e^{-\frac{i}{\hbar} \alpha \vec{n} \cdot \vec{S}} \psi(R^{-1} \vec{x})$$

$\alpha$ : angle,  $\vec{n}$ : axis,  $\vec{S} = \frac{\hbar}{2} \vec{\sigma}$

$$\vec{\sigma} = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \text{ Pauli matrices}$$

infinitesimal:  $\delta\psi = -\frac{i}{\hbar} \delta \vec{n} \cdot \vec{J} \psi$

$$\vec{J} = \vec{L} + \vec{S}, \quad \vec{L} = \vec{x} \times \vec{p}$$

$\vec{L}, \vec{S}$  vector operators

Schrodinger equation:  $i\hbar \frac{\partial}{\partial t} \psi = \left( \frac{1}{2m} [\vec{\sigma} \cdot (\frac{\hbar}{i} \vec{\nabla} - q\vec{A})]^2 + q\phi \right) \psi$

The operator  $\vec{\sigma} \cdot \vec{\nabla}$  appears. It is a scalar:

$$[J_k, \vec{\sigma} \cdot \vec{\nabla}] = 0, \text{ therefore}$$

$$\vec{\sigma} \cdot \vec{\nabla} e^{\frac{i}{\hbar} \alpha \vec{n} \cdot \vec{J}} \psi = e^{\frac{i}{\hbar} \alpha \vec{n} \cdot \vec{J}} \vec{\sigma} \cdot \vec{\nabla} \psi$$

Proof:  $[J_k, \sigma_i \nabla_i] = [L_k + S_k, \sigma_i \nabla_i]$

$$= \sigma_i [L_k, \nabla_i] + [S_k, \sigma_i] \nabla_i$$

$$[L_k, \nabla_i] = \frac{\hbar}{i} \epsilon_{k\ell m} [\hat{x}_\ell \nabla_m, \nabla_i] = \frac{\hbar}{i} \epsilon_{k\ell m} [\hat{x}_\ell, \nabla_i] \nabla_m$$

$$= \frac{\hbar}{i} \epsilon_{k\ell m} (-\delta_{\ell i}) \nabla_m = i\hbar \epsilon_{kim} \nabla_m$$

$$[S_k, \sigma_i] = \frac{\hbar}{2} [\sigma_k, \sigma_i] = i\hbar \epsilon_{kim} \sigma_m$$

$$\Rightarrow [J_k, \vec{\sigma} \cdot \vec{\nabla}] = i\hbar \epsilon_{kim} \left( \underbrace{\sigma_i \nabla_m}_{\text{anti-symmetric}} + \underbrace{\sigma_m \nabla_i}_{\text{symmetric } i \leftrightarrow m} \right) = 0$$

$\vec{\sigma}, \vec{\nabla}$  are both vector operators  $\rightarrow$  scalar product. □

# 8.4.2 Lorentz-invariant Differential Equation of first order

We want to generalize:

rotation group  $\rightarrow$  Lorentz-group  
3d  $\rightarrow$  4d

$\Psi$ : several components

$\vec{\sigma}$   $\rightarrow$   $\gamma^\mu$  matrices such that

$\vec{\sigma} \cdot \vec{\nabla}$   $\rightarrow$   $\gamma^\mu \partial_\mu$   
scalar operator under rotations

scalar operator under Lorentz-transformations: that is,  $\gamma^\mu$  is a 4-vector-operator

$$x^\mu \rightarrow x'^\mu = \Lambda^\mu_\nu x^\nu \stackrel{\text{infinitesimal}}{\approx} x^\mu + \omega^\mu_\nu x^\nu ; \omega^2 \approx 0 ; \omega_{\mu\nu} = -\omega_{\nu\mu}$$

Finite Lorentz-transformation:

$$\lim_{n \rightarrow \infty} \left(1 + \frac{\omega}{n}\right)^n = e^\omega$$

$$\Psi'(x) = e^{-\frac{i}{\hbar} \frac{1}{2} \omega_{\mu\nu} S^{\mu\nu}} \Psi(\Lambda^{-1} x)$$

$$\Psi'(x) \approx \left(1 - \frac{i}{2\hbar} \omega_{\mu\nu} \gamma^{\mu\nu}\right) \Psi(x) \text{ infinitesimal}$$

$$\gamma^{\mu\nu} = S^{\mu\nu} + L^{\mu\nu}$$

$$L^{\mu\nu} = +i\hbar (x^\mu \partial^\nu - x^\nu \partial^\mu)$$

$$\left(1 - \frac{i}{2\hbar} \omega_{\mu\nu} L^{\mu\nu}\right) \Psi(x) = \left(1 + \omega_{\mu\nu} x^\mu \partial^\nu\right) \Psi(x)$$

$$= \left(1 - \omega^\mu_\nu x^\nu \partial_\mu\right) \Psi(x) \approx \Psi(x^\mu - \omega^\mu_\nu x^\nu) = \Psi(\Lambda^{-1} x)$$

$$\Rightarrow \Psi'(x) \approx \Psi(\Lambda^{-1} x) - \frac{i}{2\hbar} \omega_{\mu\nu} S^{\mu\nu} \Psi(x) \approx \left(1 - \frac{i}{2\hbar} \omega_{\mu\nu} S^{\mu\nu}\right) \Psi(\Lambda^{-1} x)$$

infinitesimal

$\gamma^\mu \partial_\mu$  is a Lorentz-scalar when

$$[\gamma^{\kappa\lambda}, \gamma^\mu \partial_\mu] = 0$$

in addition it must be:  $\gamma^\mu$  is a 4-vector, in particular it transforms in the exact same way as  $\partial^\mu$

$$\frac{i}{\hbar} \left[ \frac{1}{2} \omega_{\kappa\lambda} \gamma^{\kappa\lambda}, \partial^\mu \right] = - \left[ \omega_{\kappa\lambda} x^\kappa \partial^\lambda, \partial^\mu \right]$$

$$= - \omega_{\kappa\lambda} \underbrace{[x^\kappa, \partial^\mu]}_{\delta^{\mu\kappa}} \partial^\lambda = \omega^\mu{}_\lambda \partial^\lambda$$

$$= g^{\mu\nu} [x^\kappa, \frac{\partial}{\partial x^\nu}] = -g^{\mu\kappa}$$

$$\frac{i}{\hbar} \left[ \frac{1}{2} \omega_{\kappa\lambda} S^{\kappa\lambda}, \gamma^\mu \right] \stackrel{!}{=} \omega^\mu{}_\lambda \gamma^\lambda$$

Solution: Pauli:  $\{\sigma_i, \sigma_j\} = \sigma_i \sigma_j + \sigma_j \sigma_i = 2\delta_{ij}$

$$\rightarrow \{\gamma^\mu, \gamma^\nu\} = \gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu = 2g^{\mu\nu}$$

Proposition:  $S^{\kappa\lambda} = \frac{i\hbar}{4} [\gamma^\kappa, \gamma^\lambda]$  is solution

$$\text{Proof: } \frac{i}{\hbar} \left[ \frac{1}{2} \omega_{\kappa\lambda} S^{\kappa\lambda}, \gamma^\mu \right] = -\frac{1}{4} [\omega_{\kappa\lambda} \gamma^\kappa \gamma^\lambda, \gamma^\mu]$$

$$= -\frac{1}{4} \omega_{\kappa\lambda} (\gamma^\kappa \{\gamma^\lambda, \gamma^\mu\} - \{\gamma^\mu, \gamma^\kappa\} \gamma^\lambda)$$

$$[AB, C] = A\{B, C\} - \{C, A\}B$$

$$= -\frac{1}{4} \omega_{\kappa\lambda} (\gamma^\kappa 2g^{\lambda\mu} - 2g^{\mu\kappa} \gamma^\lambda)$$

$$= -\frac{1}{2} (\omega_{\kappa}{}^\mu \gamma^\kappa - \omega^\mu{}_\lambda \gamma^\lambda)$$

$$= \omega^\mu{}_\lambda \gamma^\lambda$$

$$\Gamma \omega_{\kappa}{}^\mu = g^{\mu\lambda} \omega_{\kappa\lambda} = -g^{\mu\lambda} \omega_{\lambda\kappa} = -\omega^\mu{}_\kappa \quad \square$$

Under the assumption that the matrices  $\gamma^\mu$  exist, the equation

$$i\hbar \gamma^\mu \partial_\mu \psi(x) = mc \psi(x) \quad (\text{Dirac equation})$$

is a Lorentz-covariant equation, that is

$\psi(x)$  solution  $\Rightarrow \psi'(x)$  solution,

$$\psi'(x) = e^{\frac{i}{8} \omega_{\mu\nu} [\gamma^\mu, \gamma^\nu]} \psi(\Lambda^{-1}x)$$

↑ generalized spin

$$\Rightarrow (i\hbar \gamma^\mu \partial_\mu + mc) (i\hbar \gamma^\mu \partial_\mu - mc) \psi = 0$$

$$[\hbar^2 (\gamma^\mu \partial_\mu)^2 + m^2 c^2] \psi = 0$$

$$\hookrightarrow \gamma^\mu \gamma^\nu \partial_\mu \partial_\nu = \frac{1}{2} \{\gamma^\mu, \gamma^\nu\} \partial_\mu \partial_\nu = \partial^2$$

$(\hbar^2 \partial^2 + m^2 c^2) \psi = 0 \Leftrightarrow$  Klein Gordon equation for each component

$$\psi = u e^{-\frac{i}{\hbar} p \cdot x} \quad : \quad -E^2 + c^2 \vec{p}^2 + m^2 c^4 = 0$$

### 8.4.3 The Dirac Matrices $\gamma^\mu$

$$\left. \begin{aligned} \{\gamma^\mu, \gamma^\nu\} &= 2g^{\mu\nu} \\ \{\sigma^i, \sigma^j\} &= 2\delta_{ij} \end{aligned} \right\} \text{pairwise anticommuting}$$

$$\sigma = \left( \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \right)$$

There are no four anticommuting  $2 \times 2$  matrices

- smallest representation is 4-dimensional:

$$\gamma^0 = \begin{pmatrix} \underbrace{1}_{2} & \underbrace{0}_{2} \\ \underbrace{0}_{2} & \underbrace{-1}_{2} \end{pmatrix} \quad \gamma^k = \begin{pmatrix} \underbrace{0}_{2} & \underbrace{\sigma_k}_{2} \\ \underbrace{-\sigma_k}_{2} & \underbrace{0}_{2} \end{pmatrix}$$

- other 4x4 representations are unitary equivalent

$$\{\tilde{\gamma}^\mu, \tilde{\gamma}^\nu\} = 2g^{\mu\nu} \Rightarrow \tilde{\gamma}^\mu = U \gamma^\mu U^\dagger, \quad U U^\dagger = 1 = U^\dagger U$$

$$\psi = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \end{pmatrix}$$

Rotation:  $x'^0 = x^0$

$$x'^i = x^i + \omega^i_j x^j = x^i + \varphi (\vec{n} \times \vec{x})^i$$

$$\Rightarrow \omega_{ik} = -\omega^i_k = \varphi n^j \epsilon_{jik}, \quad \omega_{0i} = \omega_{i0} = 0$$

$$\psi'(x^0, \vec{x}') = e^{\frac{1}{8} \omega_{ij} [\gamma^i, \gamma^j]} \psi(x^0, R^{-1} \vec{x})$$

$$\frac{1}{8} \omega_{ij} [\gamma^i, \gamma^j] = \frac{\varphi}{4} n^k \epsilon_{kij} \gamma^i \gamma^j$$

$$\frac{1}{2} \epsilon_{kij} \gamma^i \gamma^j = -i \Sigma_k$$

$$\text{for example } \frac{1}{2} \epsilon_{1ij} \gamma^i \gamma^j = \gamma^2 \gamma^3 = \begin{pmatrix} 0 & \sigma_2 \\ -\sigma_2 & 0 \end{pmatrix} \begin{pmatrix} 0 & \sigma_3 \\ -\sigma_3 & 0 \end{pmatrix} = \begin{pmatrix} -i\sigma_1 & 0 \\ 0 & -i\sigma_1 \end{pmatrix}$$

$$\Rightarrow \Sigma_k = \begin{pmatrix} \sigma_k & 0 \\ 0 & \sigma_k \end{pmatrix} \rightarrow \text{Pauli doubled}$$

$$\psi' = e^{-i \frac{\varphi}{2} n^k \Sigma_k} \psi$$

$$\Rightarrow \psi_{\text{Dirac}} = \begin{pmatrix} \psi_{\text{Pauli, a}} \\ \psi_{\text{Pauli, b}} \end{pmatrix}$$

as far as the transformation under rotation is concerned, doubled Pauli-spinors

### 8.4.4 Hamiltonian Form of the Dirac Equation, free solution

$$c \gamma^0 / \text{th } \left( \gamma^0 \frac{\partial}{\partial x^0} + \gamma^k \nabla_k \right) \psi = mc \psi$$



$$\Rightarrow i\hbar \frac{\partial}{\partial t} \psi = c \frac{\hbar}{i} \vec{\alpha} \cdot \vec{\nabla} \psi + mc^2 \beta \psi$$

$$\text{with } \alpha^k = \gamma^0 \gamma^k = \begin{pmatrix} 0 & \sigma_k \\ \sigma_k & 0 \end{pmatrix}, \quad \beta = \gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$

$$\text{Hamiltonian operator: } H = c \frac{\hbar}{i} \vec{\alpha} \cdot \vec{\nabla} + mc^2 \beta$$

A) Solutions for particles at rest ( $\vec{p} = 0$ )

$$\psi = u e^{-\frac{i}{\hbar} E \cdot t} \Rightarrow (E - mc^2 \beta) \psi = 0$$

$$\begin{pmatrix} E - mc^2 & & & \\ & E - mc^2 & & \\ & & E + mc^2 & \\ & & & E + mc^2 \end{pmatrix} u = 0$$

$$\Rightarrow u^{(1)} = \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix}, \quad u^{(2)} = \begin{pmatrix} 0 \\ 1 \\ 0 \\ 0 \end{pmatrix} \quad \text{with } E = mc^2$$

$$u^{(3)} = \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix}, \quad u^{(4)} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 1 \end{pmatrix} \quad \text{with } E = -mc^2$$

Since  $[\beta, \Sigma_3] = 0$  (diagonal matrices),  $\Sigma_3$  (spin) can be simultaneously diagonalized.

$$\frac{\hbar}{2} \Sigma_3 u^{(k)} = \begin{cases} +\hbar/2 \\ -\hbar/2 \end{cases} u^{(k)} \quad \text{for } \begin{cases} k=1,3 \\ k=2,4 \end{cases}$$

	k=1	2	3	4
energy	+	+	-	-
spin	+	-	+	-

$E = -mc^2 \rightarrow$  positron at rest in Quantum Field Theory

B) now: momentum  $\vec{p} = (0, 0, p)$  imparted by boost

$$\begin{pmatrix} \hat{c} & \hat{s} \\ \hat{s} & \hat{c} \end{pmatrix} \begin{pmatrix} \pm mc \\ 0 \end{pmatrix} = \pm mc \begin{pmatrix} \hat{c} \\ \hat{s} \end{pmatrix} \stackrel{!}{=} \begin{pmatrix} E'/c \\ p \end{pmatrix}$$

$$= \Delta^\mu_\nu, \mu, \nu \in \{0, 3\}$$

$\hat{c} = \cosh \varphi$ ,  $\hat{s} = \sinh \varphi$ ; with  $E_p = +\sqrt{m^2 c^4 + p^2 c^2}$ ;

$k=1, 2$ :  $E' = +E_p$ ,  $\tanh \varphi = \frac{pc}{E_p}$

$k=3, 4$ :  $E' = -E_p$ ,  $\tanh \varphi = -\frac{pc}{E_p}$

$\psi^{(k)} = u^{(k)}(p) e^{-\frac{i}{\hbar} (E' t - p x^3)}$

$u^{(k)}(p) = e^{\frac{1}{8} \omega_{\mu\nu} [\gamma^\mu, \gamma^\nu]} u^{(k)}(0)$

Boost with "angle"  $\varphi$  in 3-direction = sequence of infinitesimal boosts with "angle"  $\delta$  and spin transformation

$u \rightarrow \left( 1 + \frac{\omega_{\mu\nu}}{8} [\gamma^\mu, \gamma^\nu] \right) u = \left( 1 + \frac{\omega_{03}}{2} \gamma^0 \gamma^3 \right) u$   
 $\approx e^{\frac{\delta}{2} \gamma^0 \gamma^3} u \quad (\omega_{03} = +\delta)$

finite boost:  $u \rightarrow e^{\frac{\varphi}{2} \gamma^0 \gamma^3} u = e^{\frac{\varphi}{2} \alpha^3} u$

$e^{\frac{\varphi}{2} \alpha^3} = \begin{pmatrix} \hat{c}_{1/2} & 0 & \hat{s}_{1/2} & 0 \\ 0 & \hat{c}_{1/2} & 0 & -\hat{s}_{1/2} \\ \hat{s}_{1/2} & 0 & \hat{c}_{1/2} & 0 \\ 0 & -\hat{s}_{1/2} & 0 & \hat{c}_{1/2} \end{pmatrix}, \hat{c}_{1/2} = \cosh \frac{\varphi}{2}, \hat{s}_{1/2} = \sinh \frac{\varphi}{2}.$

$\Rightarrow$  the boosted spinors of the 4 free solutions are:

$$u^{(1)}(p) = A \begin{pmatrix} 1 \\ 0 \\ r \\ 0 \end{pmatrix}, \quad u^{(2)}(p) = A \begin{pmatrix} 0 \\ 1 \\ 0 \\ -r \end{pmatrix}, \quad A = \text{normalization}$$

$$k=1,2: \tanh \varphi = \frac{pc}{E_p} \Rightarrow \tanh \frac{\varphi}{2} = \frac{\sinh \varphi}{\cosh \varphi + 1} = \frac{pc}{E_p + mc^2} =: r$$

$$k=3,4: \tanh \varphi = \frac{-pc}{E_p} \Rightarrow \tanh \frac{\varphi}{2} = -r \Rightarrow$$

$$u^{(3)}(p) = A \begin{pmatrix} -r \\ 0 \\ 1 \\ 0 \end{pmatrix}, \quad u^{(4)}(p) = A \begin{pmatrix} 0 \\ r \\ 0 \\ 1 \end{pmatrix}$$

$$u^{(k)}(p)^\dagger u^{(e)}(p) = |A|^2 (1+r^2) \delta_{ke}$$

$$\text{normalize} \Rightarrow A = (1+r^2)^{-1/2} = \left( \frac{2E_p}{E_p + mc^2} \right)^{-1/2}$$

# 8.5 Non-relativistic Limit of the Dirac Equation (122)

## 8.5.1 Leading Order

Free solution for  $E > 0$  ( $k=1,2$ ):

$$|u_{3,4}^{(k)}| = r |u_{1,2}^{(k)}|, \quad r = \frac{pc}{E_p + mc^2} \stackrel{\text{non-relat.}}{\sim} \frac{m v c}{2mc^2} = O\left(\frac{v}{c}\right)$$

large "upper", small "lower" components

More in general: stationary solution in the electromagnetic potential

$$E \psi = \left[ -e\phi + c \vec{\alpha} \left( \frac{\hbar}{i} \vec{\nabla} + e \vec{A} \right) + mc^2 \beta \right] \psi$$

with  $A^\mu = \left( \frac{\phi}{c}, \vec{A} \right)$ ,  $q = -e$  (electron).

$$\psi = \begin{pmatrix} \varphi \\ \chi \end{pmatrix}; \quad \varphi = \begin{pmatrix} \varphi_1 \\ \varphi_2 \end{pmatrix}; \quad \chi = \begin{pmatrix} \chi_1 \\ \chi_2 \end{pmatrix}$$

$$\beta = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}; \quad \vec{\alpha} = \begin{pmatrix} 0 & \vec{\sigma} \\ \vec{\sigma} & 0 \end{pmatrix}$$

$$\Rightarrow E \varphi = (-e\phi + mc^2) \varphi + c \vec{\alpha} \left( \frac{\hbar}{i} \vec{\nabla} + e \vec{A} \right) \chi$$

$$E \chi = (-e\phi - mc^2) \chi + c \vec{\alpha} \left( \frac{\hbar}{i} \vec{\nabla} + e \vec{A} \right) \varphi$$

$=: \vec{\pi} = \vec{p} + e \vec{A}$

Non-relativistic,  $E > 0$ :  $E = mc^2 + E'$

$$0 \leq E' \ll mc^2$$

$$|e\phi| \ll mc^2$$

$$\text{2nd equation} \Rightarrow \chi = \frac{1}{2mc^2 + E' + e\phi} c \vec{\sigma} \cdot \vec{\pi} \varphi \approx \frac{c \vec{\sigma} \cdot \vec{\pi}}{2mc^2} \varphi$$

$$|\chi| \ll |\varphi|$$

$$\begin{aligned} \text{1st equation} \Rightarrow E' \varphi &\approx -e\phi \varphi + c \vec{\sigma} \cdot \vec{\pi} \frac{\vec{\sigma} \cdot \vec{\pi}}{2mc} \varphi \\ &= \left( -e\phi + \frac{(\vec{\sigma} \cdot \vec{\pi})^2}{2m} \right) \varphi \approx \chi \end{aligned}$$

$\Leftrightarrow$  Pauli equation

$$(\vec{\sigma} \cdot \vec{\pi})^2 = \sigma_k \sigma_l \pi_k \pi_l = \pi^2 + i \sigma_m \epsilon_{klm} \pi_k \pi_l = \pi^2 + i \vec{\sigma} \cdot (\vec{\pi} \times \vec{\pi})$$

$$\vec{\pi} \times \vec{\pi} = \left( \frac{\hbar}{i} \vec{\nabla} + e\vec{A} \right) \times \left( \frac{\hbar}{i} \vec{\nabla} + e\vec{A} \right) = \frac{e\hbar}{i} \vec{\nabla} \times \vec{A} = \frac{e\hbar}{i} \vec{B}$$

$$\Rightarrow E' \varphi = \frac{1}{2m} \left( \frac{\hbar}{i} \vec{\nabla} + e\vec{A} \right)^2 \varphi - e\phi \varphi + \frac{e\hbar}{2m} \vec{B} \cdot \vec{\sigma} \varphi$$

$\underset{= \mu_B}{\phantom{\frac{e\hbar}{2m}}}$

$$E' \varphi = \frac{1}{2m} \left( \frac{\hbar}{i} \vec{\nabla} + e\vec{A} \right)^2 \varphi - e\phi \varphi + 2 \frac{\mu_B}{\hbar} \vec{B} \cdot \vec{S} \varphi$$

$\Rightarrow$  magnetic moment associated with spin with Landé factor  $g=2$  (so far empirically determined, orbital angular momentum  $\Leftrightarrow g=1$ ) from the Dirac equation!

## 8.5.2 Expansion in $\frac{v}{c}$

Special case:  $\vec{A} = 0$ ,  $-e\phi = V(|\vec{x}|)$

again:  $E = +mc^2 + E'$

$$(E' - V) \varphi = c \vec{\sigma} \cdot \hat{p} \chi$$

$$(E' - V + 2mc^2) \chi = c \vec{\sigma} \cdot \hat{p} \varphi$$

$$\Rightarrow \left[ E' - V - c^2 \vec{\sigma} \cdot \hat{p} (E' - V + 2mc^2)^{-1} \vec{\sigma} \cdot \hat{p} \right] \varphi = 0$$

still exact, not an eigenvalue equation for  $E'$ , instead

$$A(E') \varphi = 0$$

$$\text{Operator } A(E') = E' - V - \frac{1}{2m} \vec{\sigma} \cdot \hat{p} \left( 1 + \frac{E' - V}{2mc^2} \right)^{-1} \vec{\sigma} \cdot \hat{p}$$

Expansion in  $\frac{1}{c^2}$  respectively  $\frac{v^2}{c^2}$ , where  $v = \text{typical velocity}$ :  
up to and including  $O(\frac{1}{c^2})$ :

$$A = E' - V - \frac{\hat{p}^2}{2m} + \frac{1}{4m^2c^2} \vec{\sigma} \cdot \hat{p} (E' - V) \vec{\sigma} \cdot \hat{p} + O(\frac{1}{c^4})$$

$$= E' \left( 1 + \frac{\hat{p}^2}{4m^2c^2} \right) - \frac{\hat{p}^2}{2m} - V - \frac{1}{4m^2c^2} \vec{\sigma} \cdot \hat{p} V \vec{\sigma} \cdot \hat{p}$$

$$=: E' H_1 - H_2$$

$$H_1 = 1 + \frac{\hat{p}^2}{4m^2c^2} ; H_2 = \frac{\hat{p}^2}{2m} + V + \frac{1}{4m^2c^2} \vec{\sigma} \cdot \hat{p} V \vec{\sigma} \cdot \hat{p} ; H_i^+ = H_i$$

$A\psi = 0 \Leftrightarrow E' H_1 \psi = H_2 \psi$   
generalized eigenvalue problem

$$\rightarrow E' H_1^{1/2} \psi = \underbrace{H_1^{-1/2} H_2 H_1^{-1/2}}_{=: H'} \underbrace{H_1^{1/2} \psi}_{\text{eigenvector}}$$

$E' \Leftrightarrow$  eigenvalues of  $H'$  (effective theory!)

$$H_1^{\pm 1/2} = 1 \pm \frac{\hat{p}^2}{8m^2c^2} + O(\frac{1}{c^4})$$

$$\left( 1 - \frac{\hat{p}^2}{8m^2c^2} \right) \frac{\hat{p}^2}{2m} \left( 1 - \frac{\hat{p}^2}{8m^2c^2} \right) = \frac{\hat{p}^2}{2m} - \frac{(\hat{p}^2)^2}{8m^3c^2} + O(\frac{1}{c^4})$$

$$\left( \text{"} \right) V \left( \text{"} \right) = V - \frac{1}{8m^2c^2} \{V, \hat{p}^2\} + O(\frac{1}{c^4})$$

$$\vec{\sigma} \cdot \hat{p} V \vec{\sigma} \cdot \hat{p} = \frac{1}{2} \left( \hat{p}^2 V - \vec{\sigma} \cdot \hat{p} [\vec{\sigma} \cdot \hat{p}, V] + V \hat{p}^2 - [V, \vec{\sigma} \cdot \hat{p}] \vec{\sigma} \cdot \hat{p} \right)$$

$$= \frac{1}{2} \left( \{V, \hat{p}^2\} - [\vec{\sigma} \cdot \hat{p}, [\vec{\sigma} \cdot \hat{p}, V]] \right)$$

$$\Rightarrow H' = \frac{\hat{p}^2}{2m} - \frac{(\hat{p}^2)^2}{8m^3c^2} + V - \frac{1}{8m^2c^2} [\vec{\sigma} \cdot \hat{p}, [\vec{\sigma} \cdot \hat{p}, V]] + O(\frac{1}{c^4})$$

$$[\vec{\sigma} \cdot \hat{p}, V] = \frac{\hbar}{i} (\vec{\nabla} V) \cdot \vec{\sigma}$$

$$\begin{aligned}
[\vec{\sigma} \cdot \hat{p}, (\vec{\nabla} V) \cdot \vec{\sigma}] &= \sigma_k [\hat{p}_k, (\vec{\nabla} V) \cdot \vec{\sigma}] + [\sigma_k, (\vec{\nabla} V) \cdot \vec{\sigma}] \hat{p}_k \\
&= \frac{\hbar}{i} \Delta V + i 2 \epsilon_{k\ell m} \sigma_m (\nabla_\ell V) \hat{p}_k \\
&= \frac{\hbar}{i} \Delta V - 2i (\vec{\nabla} V \times \hat{p}) \cdot \vec{\sigma}
\end{aligned}$$

$$[AB, C] = A[B, C] + [A, C]B;$$

$$\sigma_k [\hat{p}_k, (\nabla_e V)] \sigma_e = \frac{\hbar}{i} (\nabla_k \nabla_e V) \sigma_k \sigma_e = \frac{\hbar}{i} (\nabla_k \nabla_e V) \frac{1}{2} \{\sigma_k, \sigma_e\} = \frac{\hbar}{i} \Delta V$$

$$H' \approx \frac{\hat{p}^2}{2m} - \frac{(\hat{p}^2)^2}{8m^3c^2} + V + \frac{\hbar^2}{8m^2c^2} \Delta V + \frac{\hbar}{4m^2c^2} (\vec{\nabla} V \times \hat{p}) \cdot \vec{\sigma}$$

relativistic energy-momentum relation

Darwin

$$V = V(|\vec{x}|), \quad \vec{\nabla} V = V' \frac{\vec{x}}{|\vec{x}|}$$

$$\text{last term} = \frac{\hbar V'}{4m^2c^2 |\vec{x}|} (\hat{x} \times \hat{p}) \cdot \vec{\sigma}$$

$$= \frac{V'}{2m^2c^2 |\vec{x}|} \vec{L} \cdot \vec{S} \quad \text{spin-orbit coupling}$$

$\vec{L} \cdot \vec{S}$  commutes with  $\vec{J} = \vec{L} + \vec{S}$ ; not with  $\vec{L}, \vec{S}$  individually  
→ compare Clebsch-Gordan

# 8.6 The Relativistic Coulomb Problem

Relativistic Coulomb problem can be solved exactly, see book of G. Münster.

Stationary solution in the potential  $A^\mu = (\frac{\phi}{c}, \vec{A}=0)$  with

$$-e\phi(|\vec{x}|) = -\frac{\gamma}{|\vec{x}|} = V(|\vec{x}|), \quad q = -e \text{ (electron)}, \quad \gamma = \frac{Ze^2}{4\pi\epsilon_0}$$

$$E\psi = (V + c\vec{\alpha} \cdot \vec{p} + mc^2\beta)\psi$$

$$\rightarrow E_{n\gamma} = mc^2 \left[ 1 + \frac{Z^2\alpha^2}{(n - \epsilon_\gamma)^2} \right]^{-1/2}$$

$$\text{with } \epsilon_\gamma = \gamma + \frac{1}{2} - \sqrt{\left(\gamma + \frac{1}{2}\right)^2 - Z^2\alpha^2} = \mathcal{O}(c^{-2})$$

$$\gamma = \text{total angular momentum: } \vec{J} = \vec{L} + \vec{S} = \vec{x} \times \vec{p} + \frac{\hbar}{2} \vec{\Sigma}$$

$$\alpha = \frac{e^2}{4\pi\epsilon_0\hbar c} \approx \frac{1}{137} \text{ fine-structure constant}$$

$$n = 1, 2, \dots$$

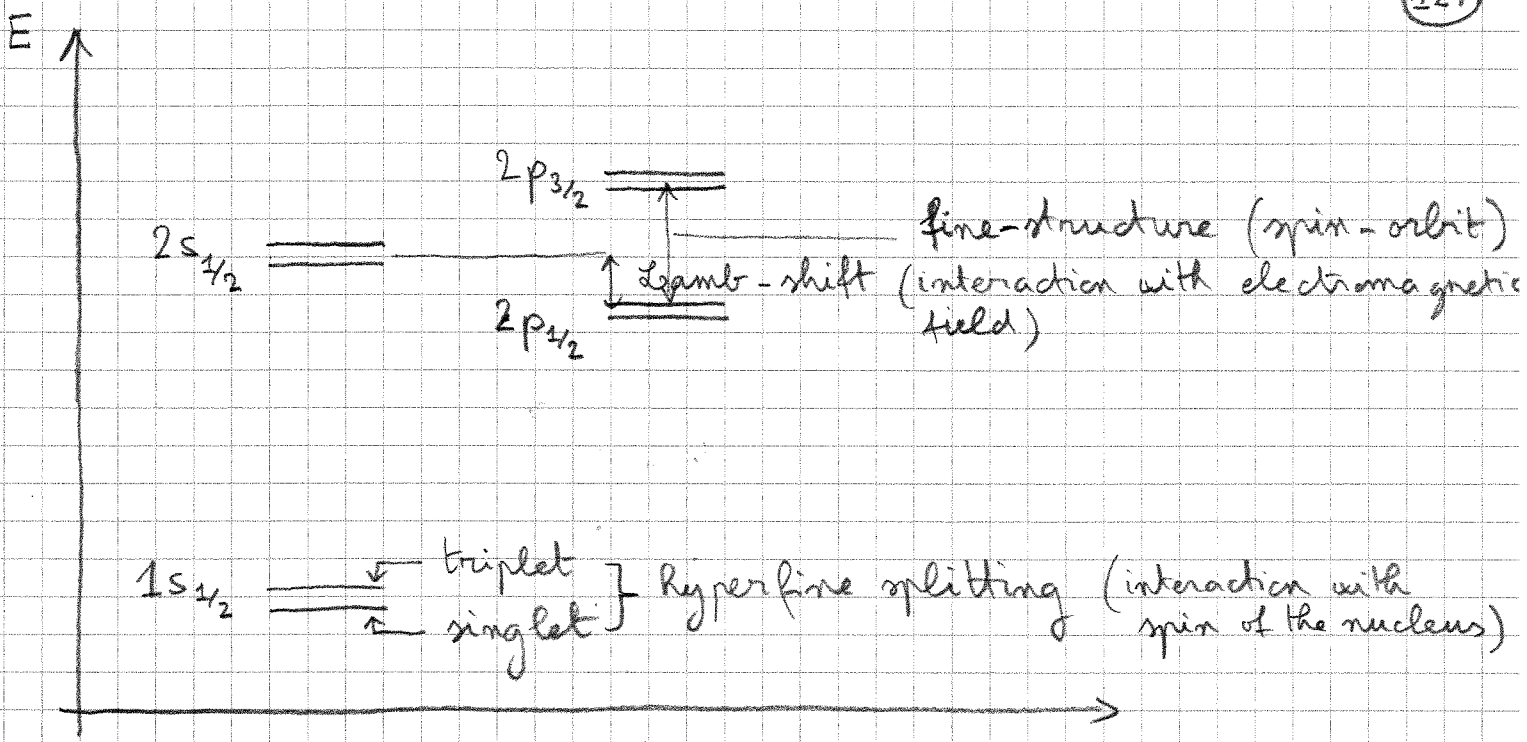
$$\gamma = \frac{1}{2}, \frac{3}{2}, \dots, n - \frac{1}{2}$$

each with  $2 \times (2\gamma + 1)$  solutions except  $\gamma = n - \frac{1}{2}$ , there  
 from  $l = \gamma \pm \frac{1}{2}$  only  $1 \times (2\gamma + 1)$   
 $l \leq n - 1$ , therefore  
 only  $\gamma = l + \frac{1}{2}$  is possible

Energy level diagram: (notation:  $n l_\gamma$ )

$n = 1$	$\gamma = \frac{1}{2}$	$1s_{\frac{1}{2}}$	
$n = 2$	$\gamma = \frac{1}{2}$	$2s_{\frac{1}{2}}$	$2p_{\frac{1}{2}}$ (degenerate)
	$\gamma = \frac{3}{2}$		$2p_{\frac{3}{2}}$





(diagram is not to scale)

Expand:

$$E_j = \left(j + \frac{1}{2}\right) \left(1 - \sqrt{1 - \left(\frac{Z\alpha}{j + \frac{1}{2}}\right)^2}\right) = \frac{1}{2} \frac{Z^2 \alpha^2}{j + \frac{1}{2}} + O(\alpha^4)$$

$$E_{nj} = mc^2 \left[ 1 - \frac{1}{2} \frac{Z^2 \alpha^2}{n^2} \left(1 + 2 \frac{E_j}{n}\right) + \frac{1}{2} \frac{3}{4} \frac{Z^4 \alpha^4}{n^4} + O(\alpha^6) \right]$$

$$= mc^2 \left[ 1 - \frac{1}{2} \left(\frac{Z\alpha}{n}\right)^2 + \left(\frac{Z\alpha}{n}\right)^4 \left(\frac{3}{8} - \frac{n}{2j+1}\right) + O(\alpha^6) \right]$$

$$\frac{1}{2} mc^2 \alpha^2 = \frac{1}{2} mc^2 \left(\frac{e^2}{4\pi\epsilon_0}\right)^2 \frac{1}{\hbar^2 c^2} = \frac{\hbar^2}{2m \left(\frac{4\pi\epsilon_0 \hbar^2}{e^2 m}\right)^2} = \frac{\hbar^2}{2ma_B^2} = E_R$$

= Rydberg energy

$$mc^2 \frac{\alpha^4}{n^4} \left(\frac{3}{8} - \frac{n}{2j+1}\right) \text{ is } = O(c^{-2}) \text{ w.r.t. } E_R$$

For  $Z \rightarrow 137$  the non-relativistic approximation becomes problematic because of large velocities of the electron.

# 9. Many Particle Systems

Atoms, nuclei, solid states

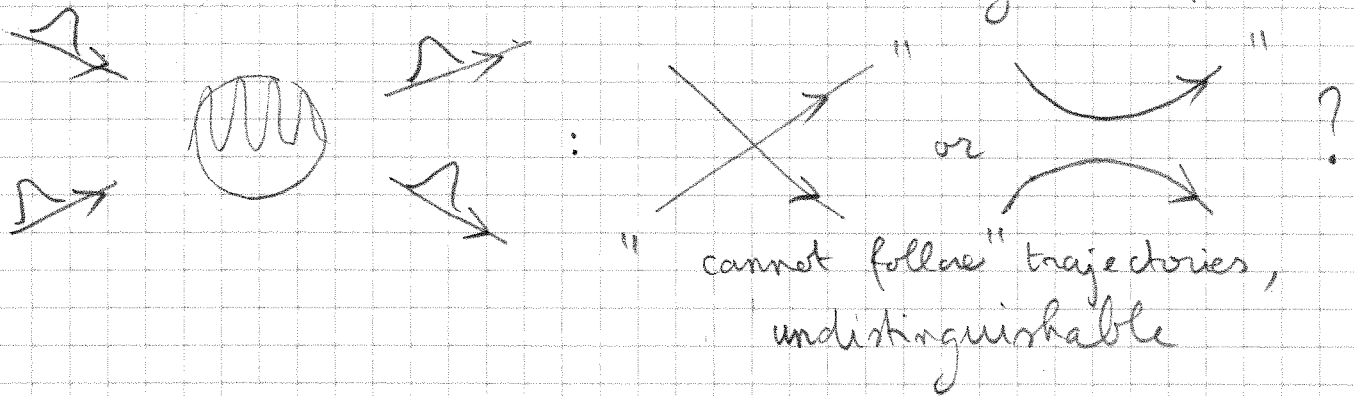
## 9.1 Identical Particles

For example  $N$   $e^-$  in the atom, same mass, spin, etc.

Classical physics:

- number the particles (somehow)
- determine  $\vec{x}_i(0), \vec{p}_i(0)$
- solve  $\rightarrow \vec{x}_i(t), i=1, \dots, N$
- identity of the particles is preserved through continuous trajectories
- different numbering at  $t=0 \Leftrightarrow$  later different numbering

Quantum mechanics, for example scattering of two particles:



$$|\Psi(t)\rangle = \int d^3x_1 d^3x_2 |\vec{x}_1\rangle |\vec{x}_2\rangle \varphi(\vec{x}_1, \vec{x}_2, t)$$

Further problem:

$\mathcal{H}$  is spanned by basis states  $\{ |\vec{x}_1 \vec{x}_2\rangle = |\vec{x}_1\rangle |\vec{x}_2\rangle \}$

Describe two identical  $e^-$

$$\langle \vec{x}_2 \vec{x}_2 | \vec{x}'_1 \vec{x}'_2 \rangle = \delta^{(3)}(\vec{x}_1 - \vec{x}'_1) \delta^{(3)}(\vec{x}_2 - \vec{x}'_2)$$

Measurement/preparation:

$e^-$  in  $V_1$  and  $e^-$  in  $V_2$  (simultaneously),  
 $V_1 \cap V_2 = \emptyset$  disjoint

Wave function  $\varphi_i \neq 0$  only in  $V_i$

State:

$$|A\rangle = \int d^3x \int d^3y \varphi_1(\vec{x}) \varphi_2(\vec{y}) |\vec{x} \vec{y}\rangle \quad \text{or} \quad \begin{array}{c} V_1 \\ \textcircled{1} \end{array} \quad \begin{array}{c} V_2 \\ \textcircled{2} \end{array}$$

$$|B\rangle = \int d^3x \int d^3y \varphi_2(\vec{x}) \varphi_1(\vec{y}) |\vec{x} \vec{y}\rangle \quad ? \quad \begin{array}{c} \textcircled{2} \\ V_1 \end{array} \quad \begin{array}{c} \textcircled{1} \\ V_2 \end{array}$$

Since  $V_1 \cap V_2 = \emptyset$  it holds  $\langle A | B \rangle = 0$

$\rightarrow$  different rays in the 2-particle Hilbert space

Experience tells: undistinguishable particles  $\rightarrow$  no other observable exist, which distinguishes  $|A\rangle, |B\rangle$

$\rightarrow n!$  states with  $n$  identical particles

QM principles? inconsistency?

## 9.2 Permutations, Bosons, Fermions

Problem  $\Leftrightarrow$  Exchange of (arbitrary but unavoidable) particle indices (cf. choice of coordinates, necessary but physics independent  $\rightarrow$  structural consequences like Galilei / Lorentz invariance)

Permutations  $\pi \in S_n : (1, 2, \dots, n) \rightarrow (\pi(1), \pi(2), \dots, \pi(n))$

$\pi(j) \in \{1, \dots, n\}$

$\pi$  is invertible  $\rightarrow 1, \dots, n$  appear only once in the range  $\{\pi(j)\}$

$\rightarrow |S_n| = n!$  different permutations

Transposition  $\pi_{ij} (i \neq j)$

$$\pi_{ij}(k) = \begin{cases} j & \text{if } k=i \\ i & \text{if } k=j \\ k & \text{otherwise} \end{cases}$$

$$\pi = \pi_{i_2 j_2} \circ \pi_{i_1 j_1} \circ \dots \circ \pi_{i_n j_n}$$

- $\pi$  can always be decomposed as composition of permutations
- not unique
- $(-1)^N$  is uniquely determined by  $\pi$ ,  $= \text{sig}(\pi)$

$S_n$  is a group

in general  $\pi \circ \pi' \neq \pi' \circ \pi$  (not abelian)

Induced unitary representation  $P$  in the Product-Hilbert-space:

$$\mathcal{H} = \underbrace{\mathcal{H}_1 \otimes \mathcal{H}_1 \otimes \dots \otimes \mathcal{H}_1}_{n \text{ factors}}$$

$$|\varphi\rangle = \sum_{i_1, \dots, i_n} \varphi_{i_1 \dots i_n} \underbrace{|i_1\rangle |i_2\rangle \dots |i_n\rangle}_{= |i_1 \dots i_n\rangle}$$

or

$$|\varphi\rangle = \int d^3x_1 \dots d^3x_n \varphi(\vec{x}_1, \dots, \vec{x}_n) |\vec{x}_1\rangle |\vec{x}_2\rangle \dots |\vec{x}_n\rangle$$

$$\langle i_1 \dots i_n | P(\pi) | \varphi \rangle = \langle i_{\pi^{-1}(1)} \dots i_{\pi^{-1}(n)} | \varphi \rangle$$

$P(\pi)$  is a unitary representation of  $S_n$

(cf. for example rotations  $\langle \vec{x} | U(R) | \varphi \rangle = \langle R^{-1} \vec{x} | \varphi \rangle$ )

Transposition:  $P(\pi_{ij}) = P_{ij}$

$$\langle k_1 \dots k_i \dots k_j \dots k_n | P_{ij} | \varphi \rangle = \langle k_1 \dots k_j \dots k_i \dots k_n | \varphi \rangle$$

$\uparrow \quad \uparrow \quad \quad \quad \quad \quad \quad \quad \uparrow \quad \uparrow$

$$P_{ij} |k_1\rangle \dots |k_i\rangle \dots |k_j\rangle \dots |k_n\rangle = |k_1\rangle \dots |k_j\rangle \dots |k_i\rangle \dots |k_n\rangle$$

$$P_{12} |\vec{x}_1 \vec{x}_2\rangle = |\vec{x}_2 \vec{x}_1\rangle$$

→ previous example :  $P_{12} |A\rangle = |B\rangle$

Postulate of symmetrization:

Physical states of  $n$  identical particles are invariant under permutations:  $P(\pi) |\psi\rangle = \lambda |\psi\rangle, \lambda \in \mathbb{C}$

- it is enough to consider  $P_{ij}$

- it holds  $(P_{ij})^2 = 1$

-  $P_{ij} |\psi\rangle = \pm |\psi\rangle$

$$\Rightarrow P(\pi) |\psi\rangle = \begin{cases} 1 \\ \text{sig}(\pi) \end{cases} |\psi\rangle$$

Case "+": bosons

Case "-": fermions (Pauli principle)

For the localization of two particles in  $V_1 \cap V_2 = \emptyset$  this implies not  $|A\rangle$ , not  $|B\rangle$  but instead

$$|\psi_+\rangle = |A\rangle + |B\rangle \rightarrow P_{12} |\psi_+\rangle = |\psi_+\rangle \text{ or}$$

$$|\psi_-\rangle = |A\rangle - |B\rangle \rightarrow P_{12} |\psi_-\rangle = -|\psi_-\rangle$$

$e^-$  are fermions, therefore  $|\psi_-\rangle!$

Relativistic Quantum Field Theory  $\Rightarrow$  (spin-statistics theorem, Pauli, 1940)  
half-integer spin  $\leftrightarrow$  fermions  
integer spin  $\leftrightarrow$  bosons

spin  $\frac{1}{2}$ :  $e, p, n, \text{quarks, neutrinos,}$

spin 1:  $\gamma, Z, W\text{-bosons, gluons}$ ; spin 0: Higgs

Nuclei, Atoms can be fermion or boson  $\rightarrow$  behavior is essentially different

Permutations acts on all the quantum numbers:

$$P_{12} |\vec{x}_1 \epsilon_1; \vec{x}_2 \epsilon_2\rangle = |\vec{x}_2 \epsilon_2; \vec{x}_1 \epsilon_1\rangle \quad (\epsilon_i = \pm)$$

$$\text{physical state of two } e^- : |\vec{x}_1 \epsilon_1; \vec{x}_2 \epsilon_2\rangle - |\vec{x}_2 \epsilon_2; \vec{x}_1 \epsilon_1\rangle$$

### 9.3 (Anti) symmetric Space of States

In  $\mathcal{H} = \underbrace{\mathcal{H}_\pm \otimes \dots \otimes \mathcal{H}_\pm}_{n \text{ factor}}$  the (anti) symmetric states form

a subspace  $|\varphi_\pm\rangle \in \mathcal{H}_\pm \subset \mathcal{H}$ ,  $P_{ij} |\varphi_\pm\rangle = \pm |\varphi_\pm\rangle$

$\mathcal{H}_\pm$ : physical Hilbert space of bosons (+) or fermions (-)

Matrix elements of observables:

$$\langle \varphi_\pm | \underbrace{A}_{\text{observable}} | \chi_\pm \rangle$$

$A |\chi_\pm\rangle$  is again projected onto  $\mathcal{H}_\pm$  through scalar product

$\Rightarrow$  symmetrization postulate for observables:

only observables  $A$  are admissible (as physical observables) for which  $A: \mathcal{H}_+ \rightarrow \mathcal{H}_+$  or  $\mathcal{H}_- \rightarrow \mathcal{H}_-$

$\Rightarrow A$  is fully determined through matrix elements

$$\langle \varphi_\pm | A | \chi_\pm \rangle = \langle \varphi_\pm | P_{ij}^\dagger A P_{ij} | \varphi_\pm \rangle$$

$$\Rightarrow A = P_{ij}^\dagger A P_{ij} = P_{ij} A P_{ij} \Leftrightarrow [A, P_{ij}] = 0$$

This does not hold for example for  $\hat{x} \otimes 1$ :

$$(\hat{x} \otimes 1) (|\vec{y}_1 \vec{y}_2\rangle - |\vec{y}_2 \vec{y}_1\rangle) = \vec{y}_1 |\vec{y}_2 \vec{y}_2\rangle - \vec{y}_2 |\vec{y}_2 \vec{y}_1\rangle$$

$$P_{12} (\hat{x} \otimes 1) P_{12} (|\vec{y}_1 \vec{y}_2\rangle - |\vec{y}_2 \vec{y}_1\rangle) = \text{different}$$

$$- P_{12} (\vec{y}_1 |\vec{y}_2 \vec{y}_2\rangle - \vec{y}_2 |\vec{y}_2 \vec{y}_1\rangle) = \vec{y}_2 |\vec{y}_1 \vec{y}_2\rangle - \vec{y}_1 |\vec{y}_2 \vec{y}_2\rangle$$

but as we can see the following holds:

$$P_{12} (\hat{x} \otimes 1) P_{12} = 1 \otimes \hat{x}$$

↑ exchange    ↑ exchange    ↑ measure 2nd factor  
↑ measure 1st factor

⇒  $\hat{x} \otimes 1 + 1 \otimes \hat{x}$  is an admissible observable  
 plus sign for  $\mathcal{H}_-$  as well as for  $\mathcal{H}_+$

⇒ observable is symmetric in the particle indices  
 for indistinguishable particles

Projectors  $\mathcal{H} \rightarrow \mathcal{H}_\pm$

$$S^+ = \frac{1}{n!} \sum_{\pi \in S_n} P(\pi)$$

$$S^- = \frac{1}{n!} \sum_{\pi \in S_n} \text{sig}(\pi) P(\pi)$$

Properties of projectors:

$$(S^\pm)^\dagger = S^\pm, \quad (S^\pm)^2 = S^\pm$$

$$S^+ S^- = S^- S^+ = 0, \quad \mathcal{H}_+ \perp \mathcal{H}_-$$

Easy to show:  $P_{ij} S^\pm |\varphi\rangle = \pm S^\pm |\varphi\rangle$

In general it does not hold  $S^+ + S^- = 1_{\mathcal{H}}$  (holds only for  $n=2$ )

→ There exist also subspaces of  $\mathcal{H}$ , which are neither symmetric nor antisymmetric → unphysical

Basis of  $\mathcal{H}_\pm$ :

$$S^\pm |\vec{x}_1 \vec{x}_2\rangle = \frac{1}{2} (|\vec{x}_1 \vec{x}_2\rangle \pm |\vec{x}_2 \vec{x}_1\rangle)$$

$$|\varphi_\pm\rangle = \int d^3x_1 d^3x_2 \varphi(\vec{x}_1, \vec{x}_2) \frac{1}{2} (|\vec{x}_1 \vec{x}_2\rangle \pm |\vec{x}_2 \vec{x}_1\rangle)$$

⇒ only (anti) symmetric part of the wave function contributes  
→ choose (anti) symmetric wave functions  $\varphi$

Discrete basis  $|i_1 \dots i_n\rangle = |i_1\rangle |i_2\rangle \dots |i_n\rangle$

$S^\pm |i_1 \dots i_n\rangle$  only depends on which  $\{i_1 \dots i_n\}$  appear but not on their ordering

Occupation-number basis:

1-particle state  $|0\rangle$ : appears  $N_0$  times

1-particle state  $|1\rangle$ : appears  $N_1$  times

⋮  
etc

appearing in  $\{i_1 \dots i_n\}$  with  $\sum_j N_j = n = \text{number of particles}$   
sum over all 1-particle states forming basis of  $\mathcal{H}_1$

$\mathcal{H}_+$ : all  $|N_0 N_1 N_2 \dots\rangle$  are linear independent, build a basis

$\mathcal{H}_-$ : like above, but only  $N_j = 0, 1$  admissible because

$S^- |i_1 \dots i_n\rangle = 0$  when the same index appears twice, for example  $i_1 = i_2$



Pauli exclusion principle:

each 1-particle state can be occupied by at most one  $e^-$

Do we need to antisymmetrize between all  $e^-$  in the universe?

$$|\varphi\rangle = |\varphi_E\rangle |\varphi_M\rangle \in \mathcal{H} = \mathcal{H}_E \otimes \mathcal{H}_M$$

$\uparrow$  Earth       $\uparrow$  Moon

$$\langle \varphi_E | \varphi_M \rangle = 0 \text{ no overlap}$$

$$|\varphi_-\rangle = \frac{1}{\sqrt{2}} (|\varphi_E\rangle |\varphi_M\rangle - |\varphi_M\rangle |\varphi_E\rangle)$$

Symmetric (= admissible) observable:

$$P_\Delta = \int_\Delta d^3x (|\vec{x}\rangle \langle \vec{x}| \otimes 1 + 1 \otimes |\vec{x}\rangle \langle \vec{x}|)$$

$$\langle \varphi_- | P_\Delta | \varphi_- \rangle = \text{probability for } e^- \text{ in volume } \Delta \text{ near the Earth}$$

$$= \frac{1}{2} \int_\Delta d^3x \left\{ 2 |\langle \vec{x} | \varphi_E \rangle|^2 \langle \varphi_M | \varphi_M \rangle + 2 |\langle \vec{x} | \varphi_M \rangle|^2 \langle \varphi_E | \varphi_E \rangle \right\}$$

$\approx 0$

$$= \int_\Delta d^3x |\langle \vec{x} | \varphi_E \rangle|^2$$

same result as for separate Earth / Moon systems

## 9.4 Atoms with Many Electrons

nucleus with charge  $Z$ , number  $Z$  of  $e^-$  with spin  $\frac{1}{2}$

$$H = \sum_{i=1}^Z H_i + \sum_{i < j} V_{ij}$$

$$H_i = \frac{\hat{p}_i^2}{2m} - \frac{Ze^2}{4\pi\epsilon_0} \frac{1}{|\hat{x}_i|}$$

$$V_{ij} = \frac{e^2}{4\pi\epsilon_0} \frac{1}{|\hat{x}_i - \hat{x}_j|}$$

H is symmetric,  $[H, P(\pi)] = 0$

looking for: bound states in  $\mathcal{H}_-$

very complicated!  $\rightarrow$  approximations.

Principle: first set  $V_{ij} \rightarrow 0$ .

Each  $H_i$ : 1-particle operator with eigenstates  $|n l m m_s\rangle$ ,

$$E_n = -\frac{Z^2 E_R}{n^2}$$

in  $\mathcal{H}_-$  occupation numbers  $N_{nlmms} = 1$  for the lowest Z energy eigenstates

$\rightarrow$  ground state  $E = -2E_1 - 8E_2 \dots$

No degeneracy when shells are fully occupied:

$$Z = 2, 2+8, 2+8+18, \dots$$

Better approximation: Central field approximation

$$H \approx \sum_{i=1}^Z (H_i + \bar{V}(|\vec{x}_i|))$$

$\bar{V}$  shall approximate  $\sum_{i < j} V_{ij}$

Heuristically:

$$\bar{V}(|\vec{x}_i|) \approx \begin{cases} 0 & \text{for } |\vec{x}_i| \rightarrow 0 \text{ (nucleus dominates)} \\ + \frac{(Z-1)e^2}{4\pi\epsilon_0 |\vec{x}_i|} & \text{for } |\vec{x}_i| \rightarrow \infty \text{ (when added to } \frac{-Ze^2}{4\pi\epsilon_0 |\vec{x}_i|} \text{, nucleus charge is screened } Z \rightarrow 1) \end{cases}$$

Specification of  $\bar{V}$  such that  $\sum_{i < j} V_{ij} - \sum_i \bar{V}(|\vec{x}_i|)$  is a small perturbation

Eigenstates with nucleus and  $\bar{V}$

$|n l m m_s\rangle$ ,  $E_{ne}$   
↑  
accidental degeneracy for  $\frac{1}{r}$ -potential is lifted

$E_{ne}$  can be occupied with at most  $2 \cdot (2l+1) e^-$ .

For the sequence of occupation there are heuristic and empirical rules:

	# $e^-$	sum $e^-$	
1s ( $l=0, n=1$ )	2	2	He
2s ( $l=0, n=2$ ), 2p ( $l=1, n=2$ )	8	10	Ne
3s, 3p	8	18	Ar
4s, 3d, 4p	18	36	Kr
5s, 4d, 5p	18	54	Xe
6s, 4f, 5d, 6p	32	86	Rn

Each row in the above table which starts with s- $e^-$  is called a shell. Shells are particularly stable.

### 9.5 Hartree-Fock Approximation

Look for ground state as the optimal antisymmetrized product (Slater determinant) of 1-particle states; determine their wave function!

Ansatz:  $|\varphi\rangle = \sqrt{Z!} S^{-1} |\varphi_1\rangle |\varphi_2\rangle \dots |\varphi_Z\rangle$

with  $\langle \varphi_i | \varphi_j \rangle = \delta_{ij}$

Minimize  $E = \frac{\langle \varphi | H | \varphi \rangle}{\langle \varphi | \varphi \rangle}$  through the choice of  $|\varphi_i\rangle$  above.

$\Rightarrow$  Hartree-Fock equation for  $\{|\varphi_i\rangle, i=1, \dots, Z\}$

This is an approximation. Because of  $V_{ij}$  the exact state is not a product state!

Result:

Basis  $|\vec{x}\sigma\rangle$ ,  $\sigma = \pm$  of 1-particle states

$$|\varphi_i\rangle = \sum_{\sigma} \int d^3x \varphi_{i\sigma}(\vec{x}) |\vec{x}\sigma\rangle$$

Hartree-Fock equation

$$\left( -\frac{\hbar^2}{2m} \Delta - \frac{Ze^2}{4\pi\epsilon_0 |\hat{x}|} + W_i(\hat{x}) \right) \varphi_{i\sigma}(\vec{x})$$

$$- \sum_{\sigma'} \int d^3y A_{i\sigma\sigma'}(\vec{x}, \vec{y}) \varphi_{i\sigma'}(\vec{y}) = \epsilon_i \varphi_{i\sigma}(\vec{x})$$

$$\text{with } W_i(\vec{x}) = \sum_{j \neq i} \sum_{\sigma} \int d^3y \frac{|\varphi_{j\sigma}(\vec{y})|^2}{4\pi\epsilon_0 |\vec{x} - \vec{y}|}$$

$$A_{i\sigma\sigma'}(\vec{x}, \vec{y}) = \sum_{j \neq i} \frac{\varphi_{j\sigma}(\vec{x}) \varphi_{j\sigma'}^*(\vec{y})}{4\pi\epsilon_0 |\vec{x} - \vec{y}|}$$

Solve by iteration:

$\{\varphi_{i\sigma}\} \rightarrow$  compute  $W_i, A_{i\sigma\sigma'} \rightarrow$  eigenvalue equation  
for each  $\varphi_{i\sigma}(\vec{x}) \rightarrow$  new  $W_i, A_{i\sigma\sigma'} \rightarrow$  etc. Convergence

$W_i$ : average repulsion

$A_{i\sigma\sigma'}$ : exchange term  $\Leftrightarrow$  Fermi statistics,  
Pauli principle

### Proof of Hartree-Fock equation:

norm:

$$\begin{aligned} \langle \varphi | \varphi \rangle &= Z! \langle \varphi_1 | \langle \varphi_2 | \dots \langle \varphi_n | \frac{1}{Z!} \sum_{\Pi} \text{sig}(\Pi) | \varphi_{\Pi(1)} \rangle \dots | \varphi_{\Pi(Z)} \rangle \\ &= \sum_{\Pi} \text{sig}(\Pi) \langle \varphi_1 | \varphi_{\Pi(1)} \rangle \dots \langle \varphi_Z | \varphi_{\Pi(Z)} \rangle = 1 \end{aligned}$$

where we used:  $-(S^{-})^{\dagger} = S^{-} = (S^{-})^2$

- only contribution from  $\Pi = \text{identity}$

Computation of  $\langle \varphi | H | \varphi \rangle = E$ :

$H$  is symmetric  $\Rightarrow [S^{-}, H] = 0$

same applies to  $\sum H_i, \sum_{i < j} V_{ij}$  individually

$$\begin{aligned} \sum_i \langle \varphi | H_i | \varphi \rangle &= \sum_{i=1}^Z \sum_{\Pi} \text{sig}(\Pi) \langle \varphi_1 | \dots \langle \varphi_i | H_i | \varphi_{\Pi(i)} \rangle \dots | \varphi_{\Pi(Z)} \rangle \\ &= \sum_{i=1}^Z \langle \varphi_i | H_i | \varphi_i \rangle \quad (\langle \varphi_i | \varphi_j \rangle = \delta_{ij} \rightarrow \text{only } \Pi = \text{identity contributes}) \end{aligned}$$

$$\begin{aligned} \sum_{i < j} \langle \varphi | V_{ij} | \varphi \rangle &= \sum_{i < j} \sum_{\Pi} \text{sig}(\Pi) \langle \varphi_i | \langle \varphi_j | V_{ij} | \varphi_{\Pi(i)} \rangle | \varphi_{\Pi(j)} \rangle \prod_{k \neq i, j} \langle \varphi_k | \varphi_{\Pi(k)} \rangle \\ &= \sum_{i < j} \left( \langle \varphi_i | \langle \varphi_j | V_{ij} | \varphi_i \rangle | \varphi_j \rangle - \langle \varphi_i | \langle \varphi_j | V_{ij} | \varphi_j \rangle | \varphi_i \rangle \right) \end{aligned}$$

- contributions from  $\Pi = \text{identity}$  and  $\Pi = \Pi_{ij}$  (transposition) with  $\text{sig}(\Pi_{ij}) = -1$

- exchange term  $\rightarrow$  important effects from Fermi statistics

Minimization of  $E$ : consider variations  $|\varphi_i\rangle \rightarrow |\varphi_i\rangle + |\delta_i\rangle$

and require:  $\delta E = 0$ , constraint:  $\langle \varphi_i | \varphi_j \rangle = \delta_{ij}$ !

In order to fulfill the constraint introduce Lagrange multipliers  $\epsilon_{ij} = \epsilon_{ji}^*$  and add to the energy functional

$$E = \underbrace{\sum_{i,j} \epsilon_{ij} (\langle \varphi_i | \varphi_j \rangle - \delta_{ij})}_{\text{real}}$$

→ stationary under  $|\delta_i\rangle$  (unconstrained!) and  $\delta\epsilon_{ij} \Rightarrow$  constraint conditions

$$0 = \sum_i 2 \text{Re} \langle \delta_i | H_i | \varphi_i \rangle \left[ \left( \langle \varphi_i | + \langle \delta_i | \right) H_i \left( | \varphi_i \rangle + | \delta_i \rangle \right) \right]$$

$$+ \sum_{i < j} 2 \text{Re} \left\{ \langle \delta_i | \langle \varphi_j | V_{ij} ( | \varphi_i \rangle | \varphi_j \rangle - | \varphi_j \rangle | \varphi_i \rangle ) + \langle \varphi_i | \langle \delta_j | \text{ " ( " " ) } \right\} \quad (*)$$

$$- \sum_{i,j} 2 \text{Re} (\epsilon_{ij} \langle \delta_i | \varphi_j \rangle)$$

$$[*]: \langle \delta_i | \langle \varphi_j | V_{ij} | \varphi_i \rangle | \varphi_j \rangle + \langle \varphi_i | \langle \varphi_j | V_{ij} | \delta_i \rangle | \varphi_j \rangle = \text{Re} \{ \langle \delta_i | \langle \varphi_j | V_{ij} | \varphi_i \rangle | \varphi_j \rangle \}$$

similarly for exchange term and  $|\delta_j\rangle$

$$\sum_{i < j} ( \langle \delta_i | \langle \varphi_j | X_{ij} \rangle + \langle \varphi_i | \langle \delta_j | X_{ij} \rangle ) = \sum_{i < j} \langle \delta_i | \langle \varphi_j | X_{ij} \rangle + \sum_{i > j} \langle \varphi_j | \langle \delta_i | X_{ji} \rangle = \langle \delta_i | \langle \varphi_j | X_{ij} \rangle$$

because the observable  $V_{ij}$  is symmetric

$$= \sum_i \sum_{j \neq i} \langle \delta_i | \langle \varphi_j | X_{ij} \rangle$$

$\langle \delta_i |$  is arbitrary  $\Rightarrow$

$$0 = H_i |\varphi_i\rangle + \sum_{j \neq i} \left( \underbrace{\langle \varphi_j | V_{ij} | \varphi_j \rangle}_{\text{operator!}} |\varphi_i\rangle - \langle \varphi_j | V_{ij} | \varphi_i \rangle |\varphi_j\rangle \right)$$

$\langle \varphi_j | \dots | \varphi_j \rangle$  : second tensor product factors

The solution is covariant under unitary transformation

$$|\tilde{\varphi}_i\rangle = \sum_j U_{ij} |\varphi_j\rangle \quad U_{ij}: \mathbb{Z} \times \mathbb{Z}, \text{ unitary}$$

$\{|\varphi_i\rangle\}$  solution  $\Rightarrow \{|\tilde{\varphi}_i\rangle\}$  also solution with

$$\tilde{\epsilon}_{ij} = \sum_{k \in \mathbb{Z}} U_{ik} \epsilon_{ke} U_{ke}^+$$

Choose  $U$  such that  $\tilde{\epsilon}_{ij} = \epsilon_i \delta_{ij}$

$$|\tilde{\varphi}\rangle = S^{-1} |\tilde{\varphi}_1\rangle \dots |\tilde{\varphi}_Z\rangle$$

$$= \frac{1}{Z!} \sum_{\pi} \text{sig } \pi \sum_{j_1 \dots j_Z} U_{\pi(1)j_1} \dots U_{\pi(Z)j_Z} |\varphi_{j_1}\rangle \dots |\varphi_{j_Z}\rangle$$

$$= \frac{1}{Z!} \sum_{\pi, \pi'} \text{sig } \pi U_{\pi(1)\pi'(1)} \dots U_{\pi(Z)\pi'(Z)} |\varphi_{\pi'(1)}\rangle \dots |\varphi_{\pi'(Z)}\rangle$$

$$= \det U S^{-1} |\varphi_1\rangle \dots |\varphi_Z\rangle = \det U |\varphi\rangle$$

$\Rightarrow$  same ray in  $\mathcal{H}$