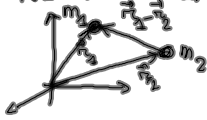


WE DEAL WITH ELEMENTARY NUMERICAL METHODS AND THEIR APPLICATION IN PHYSICS AND MATHEMATICS.

EXAMPLE: CELESTIAL MECHANICS

NEWTON'S LAW:  $\vec{F}_{12} = -\frac{Gm_1 m_2}{|\vec{r}_2 - \vec{r}_1|^2} \frac{\vec{r}_2 - \vec{r}_1}{|\vec{r}_2 - \vec{r}_1|}$



Okt 8-10:23

$$\begin{cases} m_1 \frac{d^2 \vec{r}_1}{dt^2} = \vec{F}_{12} \\ m_2 \frac{d^2 \vec{r}_2}{dt^2} = \vec{F}_{21} = -\vec{F}_{12} \end{cases}$$

2-BODY PROBLEM: EXACTLY SOLVABLE  
KEPLER'S LAWS

Okt 8-10:28

EARTH-SUN SYSTEM,



SINCE  $m_{\text{EARTH}} \approx 10^{-6} m_{\text{SUN}}$   
SUN IS APPROXIMATELY STATIONARY

ORBIT OF EARTH: ELLIPSE WITH THE SUN IN ONE OF THE FOCAL POINTS

ADD A THIRD PLANET: JUPITER  
 $m_{\text{JUPITER}} \approx 10^{-2} m_{\text{SUN}}$

Okt 8-10:31

1 AU (ASTRONOMICAL UNIT)  
= SEMIMAJOR AXIS OF EARTH'S ELLIPSE AROUND THE SUN  
=  $1.496 \times 10^{11}$  m

Okt 8-10:36

$m_{\text{JUPITER}} = m_{\text{SUN}}$  (PICTURES FROM GIORDANO AND NAKANISHI)  
TRUE 3-BODY PROBLEM  
THE SUN IS NOT STATIONARY  
ORBIT OF EARTH: UNSTABLE,  
CAN BE COMPUTED BY NUMERICAL SIMULATIONS  
CHAOS: EXTREME SENSITIVITY TO INITIAL CONDITIONS

Okt 8-10:44

MATRIX EXPONENTIAL

A  $n \times n$  COMPLEX MATRIX

$$e^A \equiv \exp(A) = \sum_{k=0}^{\infty} \frac{1}{k!} A^k$$

MATLAB: `expm(A)`

Okt 8-11:04

ARITHMETIC OPERATIONS: \* ^ /  
(MULTIPLICATION \*, EXPONENTIATION ^,  
DIVISION /)

ELEMENTWISE OPERATIONS: .\* .^ ./  
(ACT ELEMENT BY ELEMENT)

Okt 8-11:13

PREDEFINED FUNCTIONS (exp, log, ...)

ON MATRICES THEY ACT  
ELEMENT BY ELEMENT

$$\exp(A) \neq \expm(A)$$

Okt 8-11:19

FUNCTIONS IN MATLAB:

- VARIABLES CAN BE PASSED  
AS ARGUMENTS:

$$\underbrace{[a, b]}_{\text{OUTPUT}} = \text{meandiff}(\underbrace{x, y}_{\text{INPUT}})$$

- VARIABLES INSIDE FUNCTION ARE LOCAL  
(DO NOT EXIST OUTSIDE THE FUNCTION)
- WHAT MATTERS IS THE FILENAME OF A FUNCTION

Okt 8-11:37